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# Application of the sextic oscillator with a centrifugal barrier and the spheroidal equation for some $X(5)$ candidate nuclei 

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#### Abstract

The eigenvalue equation associated with the Bohr-Mottelson Hamiltonian is considered in the intrinsic reference frame and amended by replacing the harmonic oscillator potential in the $\beta$ variable with a sextic oscillator potential with centrifugal barrier plus a periodic potential for the $\gamma$ variable. After the separation of variables, the $\beta$ equation is quasi-exactly solved, while the solutions for the $\gamma$ equation are just the angular spheroidal functions. An anharmonic transition operator is used to determine the reduced $E 2$ transition probabilities. The formalism is conventionally called the sextic and spheroidal approach (SSA) and applied for several $\mathrm{X}(5)$ candidate nuclei: ${ }^{176,178,180,188,190} \mathrm{Os},{ }^{150} \mathrm{Nd},{ }^{170} \mathrm{~W},{ }^{156} \mathrm{Dy}$ and ${ }^{166,168} \mathrm{Hf}$. The SSA predictions are in good agreement with the experimental data of the mentioned nuclei. The comparison of the SSA results with those yielded by other models, such as X(5) (Iachello 2001 Phys. Rev. Lett. 87 052502), infinite square well (Raduta et al 2009 Nucl. Phys. A 819 46) and Davidson like potential (Raduta et al) for the $\beta$, otherwise keeping the spheroidal functions for the $\gamma$, and the coherent state model (Raduta et al 1981 Phys. Lett. B 99 444, Raduta et al 1982 Nucl. Phys. A 381 253, Raduta et al 1987 Phys. Rev. C 36 2111, Raduta et al 1983 Z. Phys. A 313 69, Raduta et al 1997 Phys. Rev. C 55 1747, Raduta et al 2002 Phys. Rev. C 65064322 , Raduta and Sabac 1983 Ann. Phys., NY 148 1, Raduta 2004 Recent Research Developments in Nuclear Physics vol 1) respectively, suggests that SSA represents a good approach to describe nuclei achieving the critical point of the $U(5) \rightarrow S U(3)$ shape phase transition.


(Some figures may appear in colour only in the online journal)

## 1. Introduction

Since the liquid drop model was developed [1], the quadrupole shape coordinates were widely used by both phenomenological and microscopic formalisms to describe the basic properties
of nuclear systems. Based on these coordinates, one defines quadrupole boson operators in terms of which model Hamiltonians and transition operators are defined. Since the original spherical harmonic liquid drop model was able to describe only a small amount of data for spherical nuclei, several improvements have been made. Thus, the Bohr-Mottelson model was generalized by Faessler and Greiner [2] in order to describe the small oscillations around a deformed shape which results in obtaining a flexible model, called vibration-rotation model, suitable for the description of deformed nuclei. Later on [3], this picture was extended by including anharmonicities as low-order invariant polynomials in the quadrupole coordinates. With a suitable choice of the parameters involved in the model Hamiltonian, the equipotential energy surface may exhibit several types of minima [4] such as spherical, deformed prolate, deformed oblate, deformed triaxial, etc. To each equilibrium shape, specific properties for excitation energies and electromagnetic transition probabilities show up. Due to this reason, one customarily says that static values of intrinsic coordinates determine a phase for the nuclear system. The boson description with a complex anharmonic Hamiltonian makes use of a large number of structure parameters which are to be fitted. A smaller number of parameters are used by the coherent state model (CSM) [5] which uses a restricted collective space generated through angular momentum projection by three deformed orthogonal functions of coherent type. The model is able to describe in a realistic fashion transitional and well-deformed nuclei of various shapes including states of high and very high angular momenta. Various extensions to include other degrees of freedom such as isospin [6], single particle [7] or octupole [8, 9] degrees of freedom have been formulated [10].

It has been noted that a given nuclear shape may be associated with a certain symmetry. Hence, its properties may be described with the help of the irreducible representation of the respective symmetry group. Thus, the gamma unstable nuclei can be described by the $O(6)$ symmetry [13], the gamma-rigid triaxial rotor by the $D 2$ symmetry [14], the symmetric rotor by the $S U(3)$ symmetry and the spherical vibrator by the $U(5)$ symmetry. Thus, even in the 1950s, the symmetry properties have been greatly appreciated. However, a big push forward was brought by the interacting boson approximation (IBA) [15, 16], which succeeded to describe the basic properties of a large number of nuclei in terms of the symmetries associated with a system of quadrupole (d) and monopole (s) bosons which generate the $U(6)$ algebra of the IBA. The three limiting symmetries $U(5), O(6)$ and $S U(3)$ mentioned above in the context of the collective model are also dynamic symmetries for $U(6)$. Moreover, for each of these symmetries a specific group reduction chain provides the quantum numbers characterizing the states, which are suitable for a certain region of nuclei. Besides the virtue of unifying the group theoretical descriptions of nuclei exhibiting different symmetries, the procedure defines very simple reference pictures for the limiting cases. For nuclei lying close to the region characterized by a certain symmetry, the perturbative corrections are to be included.

In $[17,18]$, it has been proved that on the $U(5)-O(6)$ transition leg there exists a critical point for a second-order phase transition, while the $U(5)-S U(3)$ leg has a first-order phase transition. Actually, the first-order phase transition takes place not only on the mentioned leg of the Casten's triangle, but covers all the interior of the triangle up to the second order [19]. Examples of such nuclei, falling inside the triangle, are the Os isotopes [20].

Recently, Iachello [21, 22] pointed out that these critical points correspond to distinct symmetries, namely $E(5)$ and $X(5)$, respectively. For the critical value of an ordering parameter, energies are given by the zeros of a Bessel function of half-integer and irrational indices, respectively.

The description of low lying states in terms of Bessel functions was used first by Jean and Willet [13], but the interesting feature saying that this is a critical picture in a phase transition and defines a new symmetry was indeed advanced first in [21].

Representatives for the two symmetries have been experimentally identified. To give an example, the relevant data for ${ }^{134} \mathrm{Ba}$ [23] and ${ }^{152} \mathrm{Sm}$ [24] suggest that they are close to the $E(5)$ and $X(5)$ symmetries, respectively. Another candidate for $E(5)$ symmetry is ${ }^{102} \mathrm{Pd}$ [25, 27]. A systematic search for $E(5)$ behavior in nuclei has been reported in [26].

In [30], we advanced the hypothesis that the critical point in a phase transition is state dependent. We tested this with a hybrid model for ${ }^{134} \mathrm{Ba}$ and ${ }^{104} \mathrm{Ru}$. A similar property of the phase transition was investigated in the context of a schematic two-level model in [31, 32]. A rigorous analysis of the characteristics of excited state quantum phase transitions is performed in [33].

The departure from the $\gamma$ unstable picture has been treated by several authors [28] whose contributions are reviewed by Fortunato in [34]. The difficulty in treating the $\gamma$ degree of freedom consists in the fact that this variable is coupled to the rotation variables. A full solution for the Bohr-Mottelson Hamiltonian including an explicit treatment of $\gamma$ deformation variable can be found in [35-39]. Therein, we treated separately also the $\gamma$ unstable and the rotor Hamiltonian. A more complete study of the rotor Hamiltonian and the distinct phases associated with a tilted moving rotor is given in [40].

The treatment of the $\gamma$ variable becomes even more complicated when we add to the liquid drop Hamiltonian a potential depending on $\beta$ and $\gamma$ at a time. To simplify the starting problem related to the inclusion of the $\gamma$ variable, one uses model potentials which are sums of a beta and a $\gamma$ depending potentials. In this way, the nice feature for the beta variable to be decoupled from the remaining four variables, specific to the harmonic liquid drop, is preserved. Further, the potential in $\gamma$ is expanded either around to $\gamma=0$ or around $\gamma=\frac{\pi}{6}$. In the first case, if only the singular term is retained, one obtains the infinite square well (ISW) model described by Bessel functions in gamma. If the $\gamma^{2}$ term is added to this term, then the Laguerre functions are the eigenstates of the approximated gamma depending Hamiltonian, which results in defining the functions characterizing the $X(5)$ approach.

The drawback of these approximation consists in that the resulting $\gamma$ depending functions are not periodic as the starting Hamiltonian is. Moreover, they are orthonormalized on unbound intervals although the underlying equation was derived under the condition of $|\gamma|$ small. The scalar product of the space of the resulting functions is not defined based on the measure $|\sin 3 \gamma| \mathrm{d} \gamma$ as happens in the liquid drop model. Under these circumstances it happens that the approximated Hamiltonian in $\gamma$ looses its hermiticity.

In some earlier publications [41, 42], we proposed a scheme where the gamma variable is described by a solvable Hamiltonian whose eigenstates are spheroidal functions which are periodic. Here, we give details about the calculations and describe some new numerical applications. Moreover, the formalism was completed by treating the $\beta$ variable by a Schrödinger equation associated with the Davidson potential. Alternatively, we considered the equation for a five-dimensional square well potential. We have shown that the new treatment of the gamma variable removes the drawbacks mentioned above and moreover brings a substantial improvement of the numerical analysis.

Here, we keep the description of the gamma variable by spheroidal functions and use a new potential for the beta variable which seems to be more suitable for a realistic description of more complex spectra. We call this approach as sextic and spheroidal approach (SSA). The potential is that of a sextic oscillator plus a centrifugal term which leads to a quasi-exactly solvable model. The resulting formalism will be applied to ten nuclei which were not included in our previous descriptions and moreover are suspected to be good candidate for exhibiting $X(5)$ features having the ratio of excitation energies of the ground band members $4^{+}$and $2^{+}$ close to the value of 2.9. The results of our calculations are compared with those obtained through other methods such as the ISW, D and CSM.

The goals presented in the previous section will be developed according to the following plan. In section 2, the main ingredients of the theoretical models $X(5)$, ISW, D and SSA will be briefly presented. The CSM is separately described in section 3 . Numerical results are given and commented in section 4 , while the final conclusions are drawn in section 5.

## 2. The separation of variables and solutions

In order to describe the critical nuclei of the $U(5)-S U(3)$ shape phase transition, we resort the Bohr-Mottelson Hamiltonian with a potential depending on both the $\beta$ and $\gamma$ variables:

$$
\begin{equation*}
H \psi(\beta, \gamma, \Omega)=E \psi(\beta, \gamma, \Omega) \tag{2.1}
\end{equation*}
$$

where
$H=-\frac{\hbar^{2}}{2 B}\left[\frac{1}{\beta^{4}} \frac{\partial}{\partial \beta} \beta^{4} \frac{\partial}{\partial \beta}+\frac{1}{\beta^{2} \sin 3 \gamma} \frac{\partial}{\partial \gamma} \sin 3 \gamma \frac{\partial}{\partial \gamma}-\frac{1}{4 \beta^{2}} \sum_{k=1}^{3} \frac{\hat{Q}_{k}^{2}}{\sin ^{2}\left(\gamma-\frac{2 \pi}{3} k\right)}\right]+V(\beta, \gamma)$.

Here, $\beta$ and $\gamma$ are the intrinsic deformation variables, $\Omega$ denotes the Euler angles $\theta_{1}, \theta_{2}$ and $\theta_{3}, \hat{Q}_{k}$ are the angular momentum components in the intrinsic reference frame, while $B$ is the so-called mass parameter.

### 2.1. The separation of variables

To achieve the separation of variables in equation (2.1), some approximations are necessary. Choosing the potential energy in the form [13, 34]

$$
\begin{equation*}
V(\beta, \gamma)=V_{1}(\beta)+\frac{V_{2}(\gamma)}{\beta^{2}} \tag{2.3}
\end{equation*}
$$

the $\beta$ variable is separated from the $\gamma$ and the Euler angles $\Omega$, which are still coupled due to the rotational term:

$$
\begin{equation*}
W=\frac{1}{4} \sum_{k=1}^{3} \frac{\hat{Q}_{k}^{2}}{\sin ^{2}\left(\gamma-\frac{2 \pi}{3} k\right)} . \tag{2.4}
\end{equation*}
$$

Furthermore, the $\gamma$ is separated from the Euler angles by using the second-order power expansion of the rotational term around the equilibrium value $\gamma_{0}=0^{0}$ (see equation (B.5) from [42]):
$W \approx \frac{1}{3} \hat{Q}^{2}+\left(\frac{1}{4 \sin ^{2} \gamma}-\frac{1}{3}\right) \hat{Q}_{3}^{2}+\frac{2}{2 \sqrt{3}}\left(\hat{Q}_{2}^{2}-\hat{Q}_{1}^{2}\right) \gamma+\frac{2}{3}\left(\hat{Q}^{2}-\hat{Q}_{3}^{2}\right) \gamma^{2}+\mathcal{O}\left(\gamma^{3}\right)$,
and then averaging the result with the Wigner function $D_{M, K}^{(L)}$ :

$$
\begin{equation*}
\langle W\rangle=\frac{1}{3} L(L+1)+\left(\frac{1}{4 \sin ^{2} \gamma}-\frac{1}{3}\right) K^{2}+\frac{2}{3}\left[L(L+1)-K^{2}\right] \gamma^{2} . \tag{2.6}
\end{equation*}
$$

The term $L(L+1) / 3$ multiplied by $1 / \beta^{2}$ is transferred to the equation for $\beta$,

$$
\begin{equation*}
\left[-\frac{1}{\beta^{4}} \frac{\partial}{\partial \beta} \beta^{4} \frac{\partial}{\partial \beta}+\frac{L(L+1)}{3 \beta^{2}}+v_{1}(\beta)\right] f(\beta)=\varepsilon_{\beta} f(\beta), \tag{2.7}
\end{equation*}
$$

while the sum of remaining terms, denoted with $\tilde{V}(\gamma, L, K)$, is kept in the equation for $\gamma$ :

$$
\begin{equation*}
\left[-\frac{1}{\sin 3 \gamma} \frac{\partial}{\partial \gamma} \sin 3 \gamma \frac{\partial}{\partial \gamma}+\tilde{V}(\gamma, L, K)+v_{2}(\gamma)\right] \eta(\gamma)=\tilde{\varepsilon}_{\gamma} \eta(\gamma) \tag{2.8}
\end{equation*}
$$

In equations (2.7) and (2.8), the following notations were used:
$v_{1}(\beta)=\frac{2 B}{\hbar^{2}} V_{1}(\beta), \quad v_{2}(\gamma)=\frac{2 B}{\hbar^{2}} V_{2}(\gamma), \quad \varepsilon_{\beta}=\frac{2 B}{\hbar^{2}} E_{\beta}, \quad \tilde{\varepsilon}_{\gamma}=\left\langle\beta^{2}\right\rangle \frac{2 B}{\hbar^{2}} E_{\gamma}$.
Equations (2.7) and (2.8) are to be separately solved and finally the full solution of equation (2.1) is obtained by combining the contributions coming from each variable. In what follows, we shall give the necessary details for solving the above-mentioned equations.

### 2.2. Solutions of the $\beta$ equation

Solutions of the $\beta$ equation, corresponding to different potentials, were considered by several authors [34, 43]. Here, we mention only three of them, namely the ISW, the Davidson and the sextic potentials. Details about how to solve the $\beta$ equation for these potentials can be found in [42, 44].
2.2.1. The infinite square well potential. The solution of the $\beta$ equation with an ISW potential, having the expression

$$
v_{1}(\beta)=\left\{\begin{array}{l}
0, \beta \leqslant \beta_{\omega}  \tag{2.10}\\
\infty, \beta>\beta_{\omega}
\end{array}\right.
$$

was first time given in [13] and then in [21, 22] for $E(5)$ and $X(5)$ models. The $\beta$ wavefunctions are written in terms of the Bessel functions of half-integer [21] and irrational indices [22], respectively. The solution for $X(5)$ is
$f_{s, L}(\beta)=C_{s, L} \beta^{-\frac{3}{2}} J_{v}\left(\frac{x_{s, L}}{\beta_{\omega}} \beta\right), \quad v=\sqrt{\frac{L(L+1)}{3}+\frac{9}{4}}, \quad s=1,2,3, \ldots$
Here, $C_{s, L}$ is the normalization factor, which is determined from the condition

$$
\begin{equation*}
\int_{0}^{\beta_{\omega}}\left(f_{s, L}(\beta)\right)^{2} \beta^{4} \mathrm{~d} \beta=1 \tag{2.12}
\end{equation*}
$$

The corresponding eigenvalues are given in terms of the Bessel zeros $x_{s, L}$ :

$$
\begin{equation*}
E_{\beta}(s, L)=\frac{\hbar^{2}}{2 B}\left(\frac{x_{s, L}}{\beta_{\omega}}\right)^{2} \tag{2.13}
\end{equation*}
$$

2.2.2. The Davidson potential. Choosing in equation (2.7) a Davidson potential [29] of the form

$$
\begin{equation*}
v_{1}(\beta)=\beta^{2}+\frac{\beta_{0}^{4}}{\beta^{2}} \tag{2.14}
\end{equation*}
$$

solutions are the generalized Laguerre polynomials:
$f_{n_{\beta}, m_{\beta}}(\beta)=\sqrt{\frac{2 n_{\beta}!}{\Gamma\left(n_{\beta}+m_{\beta}+1\right)}} L_{n_{\beta}}^{m_{\beta}}\left(\beta^{2}\right) \beta^{m_{\beta}-\frac{3}{2}} \mathrm{e}^{-\frac{\beta^{2}}{2}}, \quad m_{\beta}=\sqrt{\frac{L(L+1)}{3}+\frac{9}{4}+\beta_{0}^{4}}$.
The wavefunctions, $f_{n_{\beta}, m_{\beta}}(\beta)$, are normalized to unity with the integration measure $\beta^{4} \mathrm{~d} \beta$. Energies have the following expression:
$E_{\beta}\left(n_{\beta}, L\right)=\frac{\hbar^{2}}{2 B}\left(2 n_{\beta}+1+\sqrt{\frac{L(L+1)}{3}+\frac{9}{4}+\beta_{0}^{4}}\right), n_{\beta}=0,1,2, \ldots, n_{\beta}=s-1$.
2.2.3. The sextic oscillator potential with a centrifugal barrier. The solution of the $\beta$ equation with a sextic potential, for critical nuclei of the $U(5) \rightarrow S U(3)$ shape phase transition, was obtained by taking into consideration the solution of the Schrödinger equation with a sextic potential given in [45] and applied to the $E(5)$ like nuclei in [46] and to the triaxial nuclei in [44].

In order to reduce the $\beta$ equation to the Schrödinger equation with a sextic potential [45], we rewrite the averaged rotational term, given by equation (2.6), in the following form:

$$
\begin{equation*}
\langle W\rangle=[L(L+1)-2]+\left[2-\frac{2}{3} L(L+1)\right]+\left(\frac{1}{4 \sin ^{2} \gamma}-\frac{1}{3}\right) K^{2}+\frac{2}{3}\left[L(L+1)-K^{2}\right] \gamma^{2} . \tag{2.17}
\end{equation*}
$$

As already mentioned, the first term of the above equation is added to the $\beta$ equation, while the other terms remain in the $\gamma$ equation. Making the substitution $f(\beta)=\beta^{-2} \varphi(\beta)$, we have

$$
\begin{equation*}
\left[-\frac{\partial^{2}}{\partial \beta^{2}}+\frac{L(L+1)}{\beta^{2}}+v_{1}(\beta)\right] \varphi(\beta)=\varepsilon_{\beta} \varphi(\beta) \tag{2.18}
\end{equation*}
$$

The sextic potential is chosen such that to obtain the description from [44]:
$v_{1}^{ \pm}(\beta)=\left(b^{2}-4 a c^{ \pm}\right) \beta^{2}+2 a b \beta^{4}+a^{2} \beta^{6}+u_{0}^{ \pm}, \quad c^{ \pm}=\frac{L}{2}+\frac{5}{4}+M$.
Here, $c$ is a constant which has two different values, one for $L$ even and other for $L$ odd:

$$
\begin{align*}
& (M, L):(k, 0) ;(k-1,2) ;(k-2,4) ;(k-3,6) \ldots \Rightarrow c=k+\frac{5}{4} \equiv c^{+}(L \text {-even })  \tag{2.20}\\
& (M, L):(k, 1) ;(k-1,3) ;(k-2,5) ;(k-3,7) \ldots \Rightarrow c=k+\frac{7}{4} \equiv c^{-}(L \text {-odd }) \tag{2.21}
\end{align*}
$$

The constants $u_{0}^{ \pm}$are fixed such that the potential for $L$ odd has the same minimum energy as the potential for $L$ even. The solutions of equation (2.18), with the potential given by equation (2.19), are

$$
\begin{equation*}
\varphi_{n_{\beta}, L}^{(M)}(\beta)=N_{n_{\beta}, L} P_{n_{\beta}, L}^{(M)}\left(\beta^{2}\right) \beta^{L+1} \mathrm{e}^{-\frac{a}{4} \beta^{4}-\frac{b}{2} \beta^{2}}, \quad n_{\beta}=0,1,2, \ldots M \tag{2.22}
\end{equation*}
$$

where $N_{n_{\beta}, L}$ are the normalization factor, while $P_{n_{\beta}, L}^{(M)}\left(\beta^{2}\right)$ are polynomials in $x^{2}$ of $n_{\beta}$ order. The corresponding excitation energy is
$E_{\beta}\left(n_{\beta}, L\right)=\frac{\hbar^{2}}{2 B}\left[b(2 L+3)+\lambda_{n_{\beta}}^{(M)}(L)+u_{0}^{ \pm}\right], \quad n_{\beta}=0, \quad 1,2, \ldots, M$,
where $\lambda_{n_{\beta}}^{(M)}=\varepsilon_{\beta}-u_{0}^{ \pm}-4 b s$ is the eigenvalue of the equation:

$$
\begin{equation*}
\left[-\left(\frac{\partial^{2}}{\partial \beta^{2}}+\frac{4 s-1}{\beta} \frac{\partial}{\partial \beta}\right)+2 b \beta \frac{\partial}{\partial \beta}+2 a \beta^{2}\left(\beta \frac{\partial}{\partial \beta}-2 M\right)\right] P_{n_{\beta}, L}^{(M)}\left(\beta^{2}\right)=\lambda_{n_{\beta}}^{(M)} P_{n_{\beta}, L}^{(M)}\left(\beta^{2}\right) . \tag{2.24}
\end{equation*}
$$

### 2.3. Solutions of the $\gamma$ equation

2.3.1. The $X(5)$ model. Within the $X(5)$ model [22], devoted to the description of the critical point in the phase transition $S U(5) \rightarrow S U(3)$, the potential is a sum of an ISW in the $\beta$ variable and a harmonic oscillator in the $\gamma$ variable. For the rotational term and the other terms of the $\gamma$ equation, the first-order Taylor expansion around $\gamma_{0}=0^{0}$ is considered, which results in obtaining for the $\gamma$ variable the radial equation of a two-dimensional oscillator with the solution

$$
\begin{equation*}
\eta_{n_{\gamma}, K}(\gamma)=C_{n, K} \gamma^{|K / 2|} \mathrm{e}^{-(3 a) \gamma^{2} / 2} L_{n}^{|K|}\left(3 a \gamma^{2}\right), \quad n=\left(\frac{n_{\gamma}-|K|}{2}\right) \tag{2.25}
\end{equation*}
$$

6

Table 1. The fitted values of the parameters involved in the expressions of the energies and transition probabilities of the $X(5)$ model are given for each considered nucleus.

| $X(5)$ | ${ }^{176} \mathrm{Os}$ | ${ }^{178} \mathrm{Os}$ | ${ }^{180} \mathrm{Os}$ | ${ }^{188} \mathrm{Os}$ | ${ }^{190} \mathrm{Os}$ | ${ }^{150} \mathrm{Nd}$ | ${ }^{156} \mathrm{Dy}$ | ${ }^{166} \mathrm{Hf}$ | ${ }^{168} \mathrm{Hf}$ | ${ }^{170} \mathrm{~W}$ |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: |
| $B_{1}(\mathrm{keV})$ | 18.08 | 18.13 | 18.79 | 25.56 | 26.92 | 17.77 | 17.02 | 23.46 | 20.14 | 20.68 |
| $X(\mathrm{keV})$ | 822.28 | 818.68 | 880.10 | 452.71 | 438.53 | 966.50 | 950.46 | 698.15 | 770.26 | 799.14 |
| $t[\text { W.u. }]^{1 / 2}$ | 1.29 | 1.22 | 0.84 | 0.86 | 0.76 | 1.03 | 1.19 | 0.99 | 1.19 | 0.89 |
| $t^{\prime}[\text { W.u. }]^{1 / 2}$ | - | - | - | 0.92 | 1.19 | 0.49 | 0.81 | - | - | - |

where $L_{n}^{|K|}$ are the generalized Laguerre polynomials. The eigenvalue of the $\gamma$ equation has the following expression:

$$
\begin{equation*}
\varepsilon_{\gamma}=\frac{3 a}{\sqrt{\left\langle\beta^{2}\right\rangle}}\left(n_{\gamma}+1\right)-\frac{(K / 2)^{2}}{\left\langle\beta^{2}\right\rangle} \frac{4}{3}, \tag{2.26}
\end{equation*}
$$

where $a$ is a parameter characterizing the oscillator potential in the $\gamma$ variable. The total energy and wavefunction are obtained by combining the results of all variables:
$E\left(s, L, n_{\gamma}, K\right)=E_{0}+B_{1}\left(x_{s, L}\right)^{2}+A n_{\gamma}+C K^{2}$,
$\Psi(\beta, \gamma, \Omega)=\frac{1}{\sqrt{2\left(1+\delta_{K .0}\right)}} f_{s, L}(\beta)\left[\eta_{n_{\gamma}, K}(\gamma) D_{M, K}^{L}(\Omega)+(-1)^{L+K} \eta_{n_{\gamma},-K}(\gamma) D_{M,-K}^{L}(\Omega)\right]$.

If the total energy (2.27) is normalized to the energy of the ground state, then we will have for the ground band and for the first beta band the expression
$E(s, L, 0,0)-E(1,0,0,0)=B_{1}\left(x_{s, L}^{2}-x_{1,0}^{2}\right), \quad s=1, \quad 2 ; \quad L=0,2,4,6, \ldots$,
while for the first $\gamma$ band
$E(s, L, 1,2)-E(1,0,0,0)=B_{1}\left(x_{1, L}^{2}-x_{1,0}^{2}\right)+A+4 C, \quad L=2,3,4,5, \ldots$
One notes that the parameters $A$ and $C$ give contribution only to the $\gamma$ band energies, and that these two parameters can be replaced with only one parameter, for example $X=A+4 C$. The total energy for the ground band and for the first $\beta$ and $\gamma$ bands, normalized to the energy of the ground state, can be written in the following form:

$$
\begin{equation*}
E\left(s, L, n_{\gamma}, K\right)-E(1,0,0,0)=B_{1}\left(x_{s, L}^{2}-x_{1,0}^{2}\right)+\delta_{K, 2} X . \tag{2.31}
\end{equation*}
$$

Further, the parameters $B_{1}$ and $X$ will be fitted by the least-squares procedure for each considered nucleus.
2.3.2. The ISW model. Within the ISW model, employed in this paper, the $\beta$ equation is treated as in the $X(5)$ model, using an ISW, while the $\gamma$ equation is reduced to a spheroidal equation. The ISW model was proposed by one of the authors (AAR) and his collaborators in [41] and subsequently with more details and applications in [42]. Here, only the solutions will be presented. The potential $v_{2}(\gamma)$ was chosen such that a minimum in $\gamma=0^{0}$ is achieved:

$$
\begin{equation*}
v_{2}(\gamma)=u_{1} \cos 3 \gamma+u_{2} \cos ^{2} 3 \gamma \tag{2.32}
\end{equation*}
$$

This potential is renormalized by a contribution coming from the $\gamma$ rotational term and consequently an effective reduced potential for the $\gamma$ variable results

$$
\begin{equation*}
\tilde{v}_{2}(\gamma)=u_{1} \cos 3 \gamma+u_{2} \cos ^{2} 3 \gamma+\frac{9}{4 \sin ^{3} 3 \gamma} \tag{2.33}
\end{equation*}
$$

Table 2. The same as in table 1, but for the ISW model.

| ISW | ${ }^{176}$ Os | ${ }^{178} \mathrm{Os}$ | ${ }^{180} \mathrm{Os}$ | ${ }^{188}$ Os | ${ }^{190} \mathrm{Os}$ | ${ }^{150} \mathrm{Nd}$ | ${ }^{156}$ Dy | ${ }^{166} \mathrm{Hf}$ | ${ }^{168} \mathrm{Hf}$ | ${ }^{170} \mathrm{~W}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $B_{1}(\mathrm{keV})$ | 14.30 | 14.54 | 13.21 | 25.50 | 21.83 | 14.68 | 11.43 | 23.31 | 19.12 | 14.87 |
| $F(\mathrm{keV})$ | 24.24 | 23.19 | 44.66 | 0.69 | 36.73 | 28.88 | 45.99 | 1.69 | 11.30 | 41.12 |
| $u_{1}$ | -159.24 | -168.08 | -36.729 | -25000 | -4999.35 | -152.35 | -12.55 | -10000 | -385.35 | -44.36 |
| $u_{2}$ | 0 | 0 | 0 | 0 | 2560.22 | 0 | 0 | 0 | 0 | 0 |
| $t_{1}[\text { W.u. }]^{1 / 2}$ | -52.91 | 473.53 | 3302.3 | 503.11 | 419.67 | 538.99 | 591.54 | 1881.39 | 1197.94 | 1827.11 |
| $t_{2}[\text { W.u. }]^{1 / 2}$ | -4305.14 | -1323.6 | 14304.2 | -241.19 | -48.09 | -387.08 | -468.57 | 8242.45 | 2702.98 | 6436.57 |

Table 3. The same as in table 1, but for the D model.

| D | ${ }^{176} \mathrm{Os}$ | ${ }^{178} \mathrm{Os}$ | ${ }^{180} \mathrm{Os}$ | ${ }^{188} \mathrm{Os}$ | ${ }^{190} \mathrm{Os}$ | ${ }^{150} \mathrm{Nd}$ | ${ }^{156}$ Dy | ${ }^{166} \mathrm{Hf}$ | ${ }^{168} \mathrm{Hf}$ | ${ }^{170} \mathrm{~W}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $E(\mathrm{keV})$ | 316.34 | 317.31 | 334.32 | 559.76 | 462.44 | 369.50 | 324.08 | 532.22 | 463.88 | 379.93 |
| $F(\mathrm{keV})$ | 38.41 | 37.33 | 39.01 | 28.48 | 42.45 | 26.48 | 33.11 | 11.87 | 25.87 | 37.72 |
| $\beta_{0}$ | 1.64 | 1.56 | 1.61 | 1.98 | 1.64 | 1.71 | 1.45 | 1.79 | 2.02 | 1.63 |
| $u_{1}$ | -55.48 | -57.20 | $-52.40$ | -7.70 | -4098.61 | -168.78 | -58.16 | -320.01 | -130.49 | -54.50 |
| $u_{2}$ | 0 | 0 | 0 | 0 | 2167.18 | 0 | 0 | 0 | 0 | 0 |
| $t_{1}[\text { W.u. }]^{1 / 2}$ | 197.92 | 264.47 | 758.41 | 126.88 | 126.70 | 154.70 | 191.28 | 448.76 | 329.01 | 411.63 |
| $t_{2}[\text { W.u. }]^{1 / 2}$ | -25.31 | 78.30 | 931.21 | -17.09 | -3.92 | -25.31 | -13.46 | 430.42 | 193.06 | 363.66 |

Table 4. The same as in table 1, but for the SSA model.

| SSA | ${ }^{176} \mathrm{Os}$ | ${ }^{178} \mathrm{Os}$ | ${ }^{180} \mathrm{Os}$ | ${ }^{188} \mathrm{Os}$ | ${ }^{190} \mathrm{Os}$ | ${ }^{150} \mathrm{Nd}$ | ${ }^{156}$ Dy | ${ }^{166} \mathrm{Hf}$ | ${ }^{168} \mathrm{Hf}$ | ${ }^{170} \mathrm{~W}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $E(\mathrm{keV})$ | 0.99 | 0.46 | 1.46 | 2.53 | 5.29 | 0.75 | 0.91 | 1.82 | 0.54 | 0.31 |
| $F(\mathrm{keV})$ | 2.67 | 3.12 | 1.69 | 11.31 | 5.55 | 3.87 | 1.93 | 15.97 | 1.99 | 2.84 |
| $a$ | 951.49 | 4466.56 | 600.70 | 644.98 | 111.79 | 2636.48 | 1248.40 | 1205.13 | 7897.62 | 13197.99 |
| $b$ | 126 | 279 | 50 | 27 | 15.8 | 88 | 87 | 46 | 32 | 341 |
| $u_{1}$ | -5607.45 | -4048.06 | -15000 | -215.19 | -452.74 | -3877.84 | $-10000$ | -224.90 | -9980.01 | -4585.44 |
| $u_{2}$ | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| $t_{1}[\text { W.u. }]^{1 / 2}$ | 376.70 | 2260.6 | 8541.32 | 1033.43 | 675.12 | 1754.26 | 1882.91 | 4759.23 | 3463.05 | 8901.59 |
| $t^{t_{2}[W . u .]^{1 / 2}}$ | -32619.3 | -22343.8 | 117781 | -1022.41 | 32.73 | -6698.41 | -4846.17 | 46113.3 | 15247.9 | 200989 |

Table 5. The same as in table 1, but for the CSM model.

| CSM | ${ }^{176} \mathrm{Os}$ | ${ }^{178} \mathrm{Os}$ | ${ }^{180} \mathrm{Os}$ | ${ }^{188}$ Os | ${ }^{190} \mathrm{Os}$ | ${ }^{150} \mathrm{Nd}$ | ${ }^{156}$ Dy | ${ }^{166} \mathrm{Hf}$ | ${ }^{168} \mathrm{Hf}$ | ${ }^{170} \mathrm{~W}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $A_{1}(\mathrm{keV})$ | 17.03 | 17.26 | 16.51 | 10.25 | 9.063 | 19.219 | 15.45 | 14.87 | 16.04 | 16.19 |
| $A_{2}(\mathrm{keV})$ | 4.33 | 4.32 | 5.19 | 14.40 | 15.68 | 3.467 | 5.2 | 7.13 | 6.40 | 6.018 |
| $A_{3}(\mathrm{keV})$ | -395.96 | -240.13 | -7.39 | 101.362 | 6.84 | -658.299 | -559.913 | -5.04 | -61.47 | -186.946 |
| $A_{4}(\mathrm{keV})$ | -275.24 | -158.87 | 13.83 | 0.0 | 0.0 | -491.884 | -398.775 | 0.0 | -36.48 | -124.55 |
| $A_{5}$ (keV) | -4.93 | 30.76 | 80.01 | 0.0 | 0.0 | -438.394 | -32.15 | 0.0 | 0.0 | 0.0 |
| $d$ | 2.33 | 2.36 | 2.26 | 2.35 | 2.05 | 2.42 | 2.1 | 2.08 | 2.43 | 2.14 |
| $q_{1}[\text { W.u. }]^{1 / 2}$ | 0.411 | 0.246 | 0.86 | 0.409 | 0.229 | 0.527 | 1.112 | 0.158 | 0.211 | -0.217 |
| $q_{2}[\text { W.u. }]^{1 / 2}$ | -3.698 | -3.862 | 6.99 | 0.785 | 1.213 | -4.916 | -9.474 | -5.075 | -3.936 | -5.602 |
| $q_{3}[\text { W.u. }]^{1 / 2}$ | 0.0 | 0.0 | 0.0 | -5.222 | -9.395 | 6.344 | 19.576 | 0.0 | 0.0 | 0.0 |

Table 6. The energy spectra of the ground and first $\beta$ and $\gamma$ bands of the ${ }^{176} \mathrm{Os}$ nucleus yielded by the $X(5)$, ISW, D, SSA and CSM models are compared with the corresponding experimental data taken from [53]. The energies are given in keV units. The approach which describe best the experimental data is mentioned in a box.

| ${ }^{176} \mathrm{Os}$ | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2_{g}^{+}$ | 135 | 126 | 115 | 125 | 125 | 135 |
| $4_{8}^{+}$ | 396 | 367 | 340 | 386 | 377 | 394 |
| $6_{8}^{+}$ | 743 | 686 | 647 | 746 | 723 | 742 |
| $8_{8}^{+}$ | 1158 | 1072 | 1026 | 1176 | 1143 | 1159 |
| $10_{g}^{+}$ | 1634 | 1520 | 1473 | 1661 | 1624 | 1631 |
| $12_{g}^{+}$ | 2168 | 2028 | 1986 | 2192 | 2157 | 2152 |
| $14^{+}$ | 2755 | 2593 | 2564 | 2764 | 2736 | 2718 |
| $16_{g}^{+}$ | 3382 | 3216 | 3205 | 3374 | 3354 | 3326 |
| $18_{g}^{+}$ | 4019 | 3894 | 3909 | 4017 | 4008 | 3973 |
| $20_{g}^{+}$ | 4683 | 4628 | 4673 | 4693 | 4695 | 4660 |
| $22_{g}^{+}$ | 5399 | 5417 | 5499 | 5399 | 5412 | 5385 |
| $24_{g}^{+}$ | 6147 | 6261 | 6385 | 6134 | 6157 | 6147 |
| $0_{\beta}^{+}$ | 601 | 714 | 565 | 633 | 498 | 601 |
| $2_{\beta}^{+}$ | 742 | 942 | 760 | 757 | 723 | 742 |
| $4_{\beta}^{+}$ | 1026 | 1351 | 1118 | 1019 | 1075 | 1032 |
| $6_{\beta}^{+}$ | 1432 | 1865 | 1578 | 1378 | 1511 | 1432 |
| $8_{\beta}^{+}$ |  | 2458 | 2121 | 1808 | 2011 | 1914 |
| $10_{\beta}^{+}$ |  | 3121 | 2738 | 2293 | 2565 | 2411 |
| $2{ }_{\gamma}^{+}$ | 864 | 949 | 951 | 926 | 943 | 989 |
| $3_{\gamma}^{+}$ | 1038 | 1058 | 1056 | 1045 | 1053 | 1081 |
| $4_{\gamma}^{+}$ | 1224 | 1189 | 1184 | 1196 | 1195 | 1201 |
| $5_{\gamma}^{+}$ | 1410 | 1340 | 1333 | 1371 | 1345 | 1342 |
| $6_{\gamma}^{+}$ |  | 1509 | 1503 | 1568 | 1542 | 1511 |
| $7{ }_{\gamma}^{+}$ |  | 1694 | 1691 | 1784 | 1719 | 1689 |
| $8_{\gamma}^{+}$ |  | 1895 | 1898 | 2016 | 1962 | 1900 |
| $9_{\gamma}^{+}$ |  | 2111 | 2124 | 2264 | 2161 | 2106 |
| $10_{\gamma}^{+}$ |  | 2343 | 2367 | 2525 | 2444 | 2354 |
| $\underline{\text { rms (keV) }}$ |  | 156 | 119 | 25 | 41 | 39 |

whose minima are shifted with respect to the $v_{2}(\gamma)$ minima. This can be viewed as the reduced potential of

$$
\begin{equation*}
\tilde{V}_{2}=\frac{\hbar^{2}}{2 B} \tilde{v}_{2} \tag{2.34}
\end{equation*}
$$

Performing a second-order expansion in $\sin 3 \gamma$ of $v_{2}(\gamma)$ and of the terms originating from the rotational term, i.e. $\frac{9}{4 \sin ^{3} 3 \gamma}$, and then making the change of variable $x=\cos 3 \gamma$ in equation (2.8), we obtain the equation for the spheroidal functions [42]:

$$
\begin{equation*}
\left[\left(1-x^{2}\right) \frac{\partial^{2}}{\partial x^{2}}-2 x \frac{\partial}{\partial x}+\lambda_{m_{\gamma}, n_{\gamma}}-c^{2} x^{2}-\frac{m_{\gamma}^{2}}{1-x^{2}}\right] S_{m_{\gamma}, n_{\gamma}}(x)=0 \tag{2.35}
\end{equation*}
$$

Table 7. The same as in table 6, but for ${ }^{178}$ Os. The experimental data are taken from [57].

| ${ }^{178}$ Os | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: |
| $2_{g}^{+}$ | 132 | 127 | 116 | 131 | 130 | 132 |
| $4_{q}^{+}$ | 399 | 368 | 342 | 402 | 388 | 389 |
| $6_{g}^{+}$ | 762 | 688 | 650 | 769 | 739 | 736 |
| $8_{g}^{+}$ | 1194 | 1075 | 1031 | 1203 | 1163 | 1152 |
| $10_{q}^{+}$ | 1682 | 1525 | 1479 | 1689 | 1647 | 1625 |
| $12_{q}^{+}$ | 2220 | 2033 | 1994 | 2220 | 2181 | 2147 |
| $14_{q}^{+}$ | 2805 | 2600 | 2572 | 2789 | 2758 | 2715 |
| $16_{q}^{+}$ | 3429 | 3224 | 3214 | 3395 | 3374 | 3325 |
| $18_{q}^{+}$ | 4020 | 3905 | 3918 | 4033 | 4025 | 3975 |
| $20_{q}^{+}$ | 4663 | 4641 | 4684 | 4701 | 4706 | 4664 |
| $22_{q}^{+}$ | 5382 | 5432 | 5510 | 5399 | 5415 | 5391 |
| $24_{g}^{+}$ | 6155 | 6278 | 6397 | 6125 | 6150 | 6155 |
| $0_{\beta}^{+}$ | 651 | 716 | 574 | 635 | 493 | 651 |
| $2_{\beta}^{+}$ | 771 | 944 | 771 | 766 | 730 | 771 |
| $4_{\beta}^{+}$ | 1023 | 1355 | 1133 | 1037 | 1092 | 1029 |
| $6_{\beta}^{+}$ | 1396 | 1870 | 1598 | 1403 | 1535 | 1396 |
| $8_{\beta}^{+}$ |  | 2464 | 2144 | 1838 | 2041 | 1850 |
| $10_{\beta}^{+}$ |  | 3129 | 2766 | 2324 | 2599 | 2374 |
| $2_{\gamma}^{+}$ | 864 | 945 | 947 | 916 | 936 | 999 |
| $3_{\gamma}^{+}$ | 1032 | 1055 | 1052 | 1041 | 1048 | 1091 |
| $4_{\gamma}^{+}$ | 1213 | 1187 | 1181 | 1195 | 1195 | 1211 |
| $5_{\gamma}^{+}$ | 1416 | 1338 | 1331 | 1375 | 1346 | 1350 |
| $6_{\gamma}^{+}$ |  | 1507 | 1501 | 1575 | 1546 | 1519 |
| $7_{\gamma}^{+}$ |  | 1692 | 1690 | 1793 | 1725 | 1696 |
| $8_{\gamma}^{+}$ |  | 1894 | 1898 | 2027 | 1971 | 1907 |
| $9_{\gamma}^{+}$ |  | 2111 | 2123 | 2275 | 2170 | 2113 |
| $10_{\gamma}^{+}$ |  | 2343 | 2367 | 2537 | 2455 | 2361 |
| rms (keV) |  | 170 | 141 | 22 | 61 | 54 |

where

$$
\begin{align*}
& \lambda_{m_{\gamma}, n_{\gamma}}=\frac{1}{9}\left[\tilde{\varepsilon}_{\gamma}-\frac{u_{1}}{2}-\frac{11}{27} D+\frac{1}{3} L(L+1)\right] \\
& c^{2}=\frac{1}{9}\left(\frac{u_{1}}{2}+u_{2}-\frac{2}{27} D\right) \\
& m_{\gamma}=\frac{K}{2}, \quad D=L(L+1)-K^{2}-2 \tag{2.36}
\end{align*}
$$

From equation (2.36), we can determine the eigenvalue of the $\gamma$ equation:
$E_{\gamma}\left(n_{\gamma}, m_{\gamma}, L, K\right)=\frac{1}{\left\langle\beta^{2}\right\rangle} \frac{\hbar^{2}}{2 B}\left(9 \lambda_{m_{\gamma}, n_{\gamma}}(c)+\frac{u_{1}}{2}+\frac{11}{27} D-\frac{L(L+1)}{3}\right)$.
In equation (2.37), the term $u_{1} / 2$ is washed out when the total energy is normalized to the ground state energy, which results in obtaining the $\gamma$ eigenvalue depending on the sum of the $\gamma$ potential parameters, due to the term $c^{2}$. Hence, in some cases we can set one parameter to be equal to zero, for example $u_{2}$, and consequently fit only $u_{1}$. The $\gamma$ functions are normalized to unity with the integration measure $|\sin 3 \gamma| \mathrm{d} \gamma$ as the Bohr-Mottelson model requires

$$
\begin{equation*}
\frac{3\left(2 n_{\gamma}+1\right)\left(n_{\gamma}-m_{\gamma}\right)!}{2\left(n_{\gamma}+m_{\gamma}\right)!} \int_{0}^{\frac{\pi}{3}}\left|S_{m_{\gamma}, n_{\gamma}}(\cos 3 \gamma)\right|^{2}|\sin 3 \gamma| \mathrm{d} \gamma=1 \tag{2.38}
\end{equation*}
$$

Table 8. The same as in table 6, but for ${ }^{180}$ Os. The experimental data are taken from [58].

| ${ }^{180} \mathrm{Os}$ | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2_{g}^{+}$ | 132 | 131 | 124 | 133 | 125 | 147 |
| $4_{8}^{+}$ | 409 | 381 | 374 | 412 | 384 | 423 |
| $6_{8}^{+}$ | 795 | 713 | 723 | 792 | 748 | 792 |
| $8{ }_{8}^{+}$ | 1257 | 1115 | 1163 | 1244 | 1196 | 1234 |
| $10_{g}^{+}$ | 1768 | 1580 | 1688 | 1752 | 1716 | 1735 |
| $12_{g}^{+}$ | 2309 | 2108 | 2297 | 2308 | 2299 | 2291 |
| $\underline{14}{ }_{8}^{+}$ | 2875 | 2695 | 2987 | 2906 | 2937 | 2897 |
| $0_{\beta}^{+}$ | 736 | 742 | 522 | 669 | 555 | 736 |
| $2_{\beta}^{+}$ | 831 | 979 | 720 | 802 | 774 | 831 |
| $4_{\beta}^{+}$ | 1053 | 1404 | 1093 | 1080 | 1137 | 1051 |
| $6_{\beta}^{+}$ | 1379 | 1938 | 1584 | 1460 | 1596 | 1379 |
| $8_{\beta}^{+}$ |  | 2554 | 2175 | 1912 | 2133 | 1799 |
| $10_{\beta}^{+}$ |  | 3243 | 2858 | 2421 | 2734 | 2299 |
| $2+$ | 870 | 1011 | 975 | 935 | 985 | 969 |
| $3_{\gamma}^{+}$ | 1023 | 1125 | 1090 | 1062 | 1100 | 1068 |
| $4_{\gamma}^{+}$ | 1197 | 1262 | 1233 | 1221 | 1245 | 1198 |
| $5_{\gamma}^{+}$ | 1406 | 1418 | 1402 | 1406 | 1402 | 1348 |
| $6_{\gamma}^{+}$ | 1627 | 1593 | 1596 | 1614 | 1609 | 1529 |
| $7{ }_{\gamma}^{+}$ | 1881 | 1786 | 1813 | 1841 | 1797 | 1718 |
| $8_{\gamma}^{+}$ |  | 1995 | 2054 | 2084 | 2057 | 1944 |
| $9_{\gamma}^{+}$ | 2411 | 2220 | 2318 | 2344 | 2270 | 2164 |
| $10_{\gamma}^{+}$ |  | 2460 | 2604 | 2617 | 2577 | 2429 |
| rms (keV) |  | 194 | 96 | 38 | 92 | 35 |

The total energy is obtained by summing the contributions coming from the $\beta(2.13)$ and the $\gamma$ (2.37) equations:
$E\left(s, n_{\gamma}, m_{\gamma}, L, K\right)=B_{1} x_{s, L}^{2}+F\left[9 \lambda_{m_{\gamma}, n_{\gamma}}(c)+\frac{u_{1}}{2}+\frac{11}{27} D-\frac{L(L+1)}{3}\right]$,
where the following notations were introduced:

$$
\begin{equation*}
B_{1}=\frac{1}{\beta_{\omega}^{2}} \frac{\hbar^{2}}{2 B}, \quad F=\frac{1}{\left\langle\beta^{2}\right\rangle} \frac{\hbar^{2}}{2 B} \tag{2.40}
\end{equation*}
$$

The total wavefunction is
$\Psi(\beta, \gamma, \Omega)=C_{s, L} C_{n_{\gamma}, m_{\gamma}} C_{L, K} \beta^{-\frac{3}{2}} J_{\nu}\left(\frac{x_{s, L}}{\beta_{\omega}} \beta\right) S_{m_{\gamma}, n_{\gamma}}(\cos 3 \gamma)\left[D_{M, K}^{L}(\Omega)+(-1)^{L} D_{M,-K}^{L}(\Omega)\right]$,
where with $C_{n_{\nu}, m_{\nu}}$ was denoted the normalization factor of the $\gamma$ function, while $C_{L, K}$ is the normalization factor of the Wigner function:

$$
\begin{equation*}
C_{L, K}=\sqrt{\frac{2 L+1}{16 \pi^{2}\left(1+\delta_{K, 0}\right)}} \tag{2.42}
\end{equation*}
$$

2.3.3. The D model. The D model was proposed by the present authors and collaborators in [42] and differs from the ISW model by that the ISW potential for the $\beta$ variable is replaced

Table 9. The same as in table 6, but for ${ }^{188}$ Os. The experimental data are taken from [59].

| ${ }^{188} \mathrm{Os}$ | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2_{g}^{+}$ | 155 | 179 | 179 | 151 | 152 | 150 |
| $4_{8}^{+}$ | 478 | 519 | 519 | 479 | 476 | 468 |
| $6_{g}^{+}$ | 940 | 970 | 970 | 945 | 935 | 934 |
| $8_{g}^{+}$ | 1515 | 1516 | 1516 | 1512 | 1501 | 1535 |
| $10_{g}^{+}$ | 2170 | 2149 | 2150 | 2156 | 2154 | 2264 |
| $\underline{12}{ }_{g}^{+}$ | 2856 | 2867 | 2868 | 2860 | 2877 | 3116 |
| $0_{\beta}^{+}$ | 1086 | 1009 | 1007 | 1120 | 1063 | 1164 |
| $2_{\beta}^{+}$ | 1305 | 1331 | 1328 | 1270 | 1330 | 1305 |
| $4_{\beta}^{+}$ |  | 1910 | 1907 | 1599 | 1808 | 1621 |
| $6_{\beta}^{+}$ |  | 2636 | 2632 | 2064 | 2421 | 2096 |
| $8_{\beta}^{+}$ |  | 3474 | 3470 | 2632 | 3132 | 2717 |
| $10_{\beta}^{+}$ |  | 4412 | 4407 | 3276 | 3920 | 3475 |
| $2_{\gamma}^{+}$ | 633 | 631 | 631 | 627 | 641 | 665 |
| $3{ }_{\gamma}^{+}$ | 790 | 786 | 785 | 773 | 791 | 790 |
| $4_{\gamma}^{+}$ | 966 | 972 | 971 | 959 | 969 | 956 |
| $5_{\gamma}^{+}$ | 1181 | 1185 | 1185 | 1180 | 1172 | 1157 |
| $6_{\gamma}^{+}$ | 1425 | 1423 | 1423 | 1432 | 1434 | 1399 |
| $7{ }_{\gamma}^{+}$ | 1686 | 1685 | 1684 | 1709 | 1674 | 1669 |
| $8_{\gamma}^{+}$ |  | 1969 | 1969 | 2009 | 2008 | 1983 |
| $9_{\gamma}^{+}$ |  | 2275 | 2275 | 2329 | 2273 | 2318 |
| $10_{\gamma}^{+}$ |  | 2602 | 2603 | 2666 | 2670 | 2701 |
| rms (keV) |  | 27 | 27 | 16 | 13 | 36 |

with the Davidson potential (2.14). Hence, the total energy of the system is obtained by adding the energy of the $\beta$ equation with the Davidson potential given by equation (2.16) and the energy of the $\gamma$ equation (2.37):

$$
\begin{gather*}
E\left(n_{\beta}, n_{\gamma}, m_{\gamma}, L, K\right)=E\left(2 n_{\beta}+1+\sqrt{\frac{L(L+1)}{3}+\frac{9}{4}+\beta_{0}^{4}}\right) \\
+F\left[9 \lambda_{m_{\beta}, n_{\gamma}}(c)+\frac{u_{1}}{2}+\frac{11}{27} D-\frac{L(L+1)}{3}\right] \tag{2.43}
\end{gather*}
$$

where $E=\hbar^{2} / 2 B$. The total wavefunction has the expression
$\Psi(\beta, \gamma, \Omega)=C_{n_{\beta}, L} C_{n_{\gamma}, m_{\gamma}} C_{L, K} f_{n_{\beta}, L}(\beta) S_{m_{\gamma}, n_{\gamma}}(\cos 3 \gamma)\left[D_{M, K}^{L}(\Omega)+(-1)^{L} D_{M,-K}^{L}(\Omega)\right]$,
where $C_{n_{\beta}, L}$ is the normalization factor of $f_{n_{\beta}, L}(\beta)$ given by equation (2.15).
2.3.4. The present approach. In the present approach, called conventionally the SSA, a sextic potential (2.19) for the $\beta$ variable is considered, while for the $\gamma$ variable a periodic potential (2.32) with a minimum at $\gamma_{0}=0^{0}$. The $\beta$ equation is quasi-exactly solved, having the solutions given by equations (2.22) and (2.23), while the $\gamma$ equation is reduced to the spheroidal equation (2.35) with

$$
\begin{align*}
& \lambda_{m_{\gamma}, n_{\gamma}}=\frac{1}{9}\left[\tilde{\varepsilon}_{\gamma}-\frac{u_{1}}{2}-\frac{11}{27} D+\frac{1}{3} L(L+1)\right]+\frac{2 L(L+1)}{27} \\
& c^{2}=\frac{1}{9}\left(\frac{u_{1}}{2}+u_{2}-\frac{2}{27} D\right) \\
& m_{\gamma}=\frac{K}{2}, \quad D=L(L+1)-K^{2}-2 \tag{2.45}
\end{align*}
$$

Table 10. The same as in table 6, but for ${ }^{190}$ Os. The experimental data are taken from [60].

| ${ }^{190} \mathrm{Os}$ | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: |
| $2_{\beta}^{+}$ | 187 | 188 | 182 | 178 | 172 | 180 |
| $4_{\gamma}^{+}$ | 548 | 547 | 541 | 551 | 531 | 531 |
| $6_{\beta}^{+}$ | 1050 | 1022 | 1034 | 1062 | 1034 | 1031 |
| $8_{\xi}^{+}$ | 1666 | 1597 | 1647 | 1672 | 1653 | 1670 |
| $10_{g}^{+}$ | 2357 | 2264 | 2373 | 2359 | 2367 | 2441 |
| $0_{\beta}^{+}$ | 912 | 1063 | 862 | 925 | 860 | 912 |
| $2_{\beta}^{+}$ | 1115 | 1402 | 1166 | 1103 | 1168 | 1072 |
| $4_{\beta}^{+}$ |  | 2012 | 1729 | 1476 | 1682 | 1417 |
| $6_{\beta}^{+}$ |  | 2777 | 2457 | 1987 | 2331 | 1925 |
| $8_{\beta}^{+}$ |  | 3659 | 3319 | 2596 | 3083 | 2582 |
| $10_{\beta}^{+}$ |  | 4647 | 4305 | 3283 | 3921 | 3380 |
| $2_{\nu}^{+}$ | 558 | 627 | 594 | 583 | 593 | 618 |
| $3_{\gamma}^{+}$ | 756 | 789 | 756 | 750 | 754 | 756 |
| $4_{\nu}^{+}$ | 955 | 985 | 955 | 957 | 954 | 939 |
| $5_{\gamma}^{+}$ | 1204 | 1210 | 1187 | 1199 | 1172 | 1156 |
| $6_{\gamma}^{+}$ | 1474 | 1461 | 1451 | 1469 | 1459 | 1419 |
| $7_{\nu}^{+}$ |  | 1736 | 1745 | 1764 | 1718 | 1708 |
| $8_{\gamma}^{+}$ | 2090 | 2035 | 2067 | 2081 | 2080 | 2045 |
| $9_{\gamma}^{+}$ |  | 2358 | 2419 | 2417 | 2370 | 2401 |
| $10_{\gamma}^{+}$ | 2772 | 2702 | 2799 | 2770 | 2798 | 2810 |
| rms (keV) |  | 98 | 26 | 10 | 27 | 36 |

In equation (2.45), the term $2 L(L+1) / 3$ multiplied with $1 / 9$ comes from the rotational term (2.17). The expression for the total energy of the system is obtained by using equations (2.23) and (2.45):

$$
\begin{align*}
& E\left(n_{\beta}, n_{\gamma}, m_{\gamma}, L, K\right)=E\left[b(2 L+3)+\lambda_{n_{\beta}}^{(M)}+u_{0}^{ \pm}\right] \\
& \quad+F\left[9 \lambda_{m_{\beta}, n_{\gamma}}(c)+\frac{u_{1}}{2}+\frac{11}{27} D-L(L+1)\right] . \tag{2.46}
\end{align*}
$$

The corresponding wavefunction is
$\Psi(\beta, \gamma, \Omega)=N_{n_{\beta}, L} C_{n_{\gamma}, m_{\gamma}} C_{L, K} \beta^{-2} \varphi_{n_{\beta}, L}(\beta) S_{m_{\gamma}, n_{\gamma}}(\cos 3 \gamma)\left[D_{M, K}^{L}(\Omega)+(-1)^{L} D_{M,-K}^{L}(\Omega)\right]$,
where $\varphi_{n_{\beta}, L}(\beta)$ is given by equation (2.22).

### 2.4. E2 transition probabilities

The reduced $E 2$ transition probabilities are determined by

$$
\begin{equation*}
\left.B\left(E 2 ; L_{i} \rightarrow L_{f}\right)=\left|\left\langle L_{i}\right|\right| T_{2}^{(E 2)}| | L_{f}\right\rangle\left.\right|^{2} \tag{2.48}
\end{equation*}
$$

where the Rose convention [47] was used. For the ISW, D and SSA models, in equation (2.48), an anharmonic transition operator is used:

$$
\begin{align*}
T_{2 \mu}^{(E 2)}=t_{1} \beta & {\left[\cos \gamma D_{\mu 0}^{2}(\Omega)+\frac{\sin \gamma}{\sqrt{2}}\left(D_{\mu 2}^{2}(\Omega)+D_{\mu,-2}^{2}(\Omega)\right)\right] } \\
& \quad+t_{2} \sqrt{\frac{2}{7}} \beta^{2}\left[-\cos 2 \gamma D_{\mu 0}^{2}(\Omega)+\frac{\sin 2 \gamma}{\sqrt{2}}\left(D_{\mu 2}^{2}(\Omega)+D_{\mu,-2}^{2}(\Omega)\right)\right] \tag{2.49}
\end{align*}
$$

Table 11. The same as in table 6, but for ${ }^{150} \mathrm{Nd}$. The experimental data are taken from [61].

| ${ }^{150} \mathrm{Nd}$ | Exp. | $X(5)$ | ISW | D | SSA | $\boxed{C S M}$ |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: |
| $2_{g}^{+}$ | 130 | 124 | 121 | 124 | 111 | 130 |
| $4_{g}^{+}$ | 381 | 361 | 358 | 384 | 348 | 386 |
| $6_{g}^{+}$ | 720 | 675 | 682 | 738 | 683 | 734 |
| $8_{g}^{+}$ | 1130 | 1054 | 1084 | 1158 | 1098 | 1149 |
| $10_{g}^{+}$ | 1599 | 1494 | 1560 | 1625 | 1580 | 1618 |
| $12_{g}^{+}$ | 2119 | 1993 | 2106 | 2129 | 2118 | 2133 |
| $14_{g}^{+}$ | 2683 | 2549 | 2722 | 2664 | 2707 | 2688 |
| $0_{\beta}^{+}$ | 675 | 702 | 580 | 739 | 630 | 675 |
| $2_{\beta}^{+}$ | 851 | 926 | 783 | 863 | 822 | 852 |
| $4_{\beta}^{+}$ | 1138 | 1328 | 1157 | 1123 | 1158 | 1167 |
| $6_{\beta}^{+}$ | 1541 | 1833 | 1639 | 1477 | 1590 | 1541 |
| $8_{\beta}^{+}$ |  | 2415 | 2209 | 1897 | 2095 | 1931 |
| $10_{\beta}^{+}$ |  | 3067 | 2859 | 2364 | 2661 | 2319 |
| $2_{\gamma}^{+}$ | 1062 | 1091 | 1087 | 1076 | 1091 | 1101 |
| $3_{\gamma}^{+}$ | 1201 | 1198 | 1197 | 1195 | 1197 | 1191 |
| $4_{\gamma}^{+}$ | 1353 | 1327 | 1333 | 1345 | 1328 | 1310 |
| $5_{\gamma}^{+}$ |  | 1476 | 1491 | 1518 | 1474 | 1448 |
| $6_{\gamma}^{+}$ |  | 1641 | 1671 | 1713 | 1663 | 1615 |
| $7_{\gamma}^{+}$ |  | 1823 | 1872 | 1924 | 1838 | 1790 |
| $8_{\gamma}^{+}$ |  | 2020 | 2093 | 2151 | 2079 | 1998 |
| $9_{\gamma}^{+}$ |  | 2233 | 2334 | 2390 | 2276 | 2201 |
| $10_{\gamma}^{+}$ |  | 2461 | 2594 | 2641 | 2561 | 2445 |
| $\mathrm{rms}(\mathrm{keV})$ |  | 114 | 48 | 28 | 29 | 20 |

The parameters $t_{1}$ and $t_{2}$ will be determined by the least-squares method. For the $X(5)$ model, in the limit of $\gamma$-small, only the harmonic part of the transition operator (2.49) is used:

$$
\begin{equation*}
T_{2 \mu, X(5)}^{(E 2)}=t \beta D_{\mu 0}^{2}(\Omega)+t \beta \frac{\gamma}{\sqrt{2}}\left(D_{\mu 2}^{2}(\Omega)+D_{\mu, 2}^{2}(\Omega)\right) \tag{2.50}
\end{equation*}
$$

The first term of equation (2.50) gives contributions only to $\Delta K=0$ transitions, while the second term to $\Delta K=2$ transitions. For $\Delta K=0$ transitions, the matrix element of the $\gamma$ variable is reduced to the orthogonality condition, while for $\Delta K=2$ the $\gamma$ matrix element can be considered as an intrinsic transition matrix element. Finally, the reduced transition probabilities will depend on two parameters [48]. Here, we will denote these two parameters with $t$ and $t^{\prime}$ for $\Delta K=0$ and $\Delta K=2$ transitions, respectively.

## 3. The coherent state model

The CSM defines [5] first a restricted collective space whose vectors are model states of the ground, $\beta$ and $\gamma$ bands. In choosing these states, we were guided by some experimental information which results in formulating a set of criteria to be fulfilled by the searched states.

All these restrictions required are fulfilled by the following set of three deformed quadrupole boson states:

$$
\begin{equation*}
\psi_{g}=\mathrm{e}^{\left[d\left(b_{0}^{\dagger}-b_{0}\right)\right]}|0\rangle \equiv T|0\rangle, \quad \psi_{\gamma}=\Omega_{\gamma, 2}^{\dagger} \psi_{g}, \quad \psi_{\beta}=\Omega_{\beta}^{\dagger} \psi_{g}, \tag{3.1}
\end{equation*}
$$

Table 12. The same as in table 6, but for ${ }^{156} \mathrm{Dy}$. The experimental data are taken from [63].

| ${ }^{156}$ Dy | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: |
| $2_{g}^{+}$ | 138 | 119 | 114 | 140 | 131 | 168 |
| $4_{g}^{+}$ | 404 | 345 | 344 | 422 | 391 | 457 |
| $6_{g}^{+}$ | 770 | 646 | 668 | 796 | 745 | 829 |
| $8_{g}^{+}$ | 1216 | 1009 | 1079 | 1230 | 1175 | 1267 |
| $10_{\gamma}^{+}$ | 1725 | 1431 | 1572 | 1712 | 1667 | 1761 |
| $12_{+}^{+}$ | 2286 | 1908 | 2145 | 2232 | 2151 | 2307 |
| $14_{g}^{+}$ | 2888 | 2440 | 2796 | 2787 | 2807 | 2899 |
| $0_{\beta}^{+}$ | 676 | 672 | 451 | 648 | 461 | 676 |
| $2_{\beta}^{+}$ | 829 | 886 | 629 | 788 | 703 | 829 |
| $4_{\beta}^{+}$ | 1088 | 1272 | 966 | 1070 | 1068 | 1102 |
| $6_{\beta}^{+}$ | 1437 | 1755 | 1413 | 1444 | 1515 | 1452 |
| $8_{\beta}^{+}$ | 1859 | 2313 | 1955 | 1878 | 2026 | 1859 |
| $10_{\beta}^{+}$ | 2316 | 2937 | 2584 | 2360 | 2593 | 2312 |
| $2_{\gamma}^{+}$ | 891 | 1069 | 898 | 839 | 928 | 921 |
| $3_{\gamma}^{+}$ | 1022 | 1172 | 1004 | 970 | 1041 | 1024 |
| $4_{\gamma}^{+}$ | 1168 | 1296 | 1136 | 1129 | 1188 | 1159 |
| $5_{\gamma}^{+}$ | 1336 | 1438 | 1292 | 1312 | 1339 | 1312 |
| $6_{\gamma}^{+}$ | 1525 | 1596 | 1472 | 1514 | 1542 | 1497 |
| $7_{\gamma}^{+}$ | 1729 | 1771 | 1674 | 1732 | 1720 | 1686 |
| $8_{\gamma}^{+}$ | 1959 | 1960 | 1899 | 1964 | 1972 | 1913 |
| $9_{\gamma}^{+}$ | 2192 | 2163 | 2145 | 2210 | 2171 | 2131 |
| $10_{\gamma}^{+}$ | 2448 | 2381 | 2413 | 2467 | 2464 | 2395 |
| $11_{\gamma}^{+}$ | 2712 | 2613 | 2702 | 2735 | 2680 | 2636 |
| $12_{\gamma}^{+}$ | 2997 | 2859 | 3013 | 3013 | 2949 | 2934 |
| $13_{\gamma}^{+}$ | 3274 | 3118 | 3345 | 3301 | 3240 | 3153 |
| $14_{\gamma}^{+}$ |  | 3391 | 3698 | 3600 | 3606 | 3526 |
| $15_{\gamma}^{+}$ | 3861 | 3677 | 4071 | 3908 | 3847 | 3805 |
| rms (keV) |  | 232 | 114 | 35 | 90 | 41 |
|  |  |  |  |  |  |  |

where the excitation operators for $\beta$ and $\gamma$ bands are defined by
$\Omega_{\gamma, 2}^{\dagger}=\left(b^{\dagger} b^{\dagger}\right)_{2,2}+d \sqrt{\frac{2}{7}} b_{2,2}^{\dagger}, \quad \Omega_{\beta}^{\dagger}=\left(b^{\dagger} b^{\dagger} b^{\dagger}\right)_{0}+\frac{3 d}{\sqrt{14}}\left(b^{\dagger} b^{\dagger}\right)_{0}-\frac{d^{3}}{\sqrt{70}}$.
Here, $d$ is a real parameter simulating the nuclear deformation. From the three deformed states, one generates through projection three sets of mutually orthogonal states

$$
\begin{equation*}
\varphi_{J M}^{i}=N_{J}^{i} P_{M 0}^{J} \psi_{i}, \quad i=g, \beta, \gamma \tag{3.3}
\end{equation*}
$$

where $P_{M K}^{J}$ denotes the projection operator:

$$
\begin{equation*}
P_{M K}^{J}=\frac{2 J+1}{8 \pi^{2}} \int D_{M K}^{J^{*}} \hat{R}(\Omega) \mathrm{d} \Omega \tag{3.4}
\end{equation*}
$$

$N_{J}^{i}$ the normalization factors and $D_{M K}^{J}$ the rotation matrix elements. The rotation operator corresponding to the Euler angles $\Omega$ is denoted by $\hat{R}(\Omega)$. It was proved that the deformed and projected states contain the salient features of the major collective bands. Since we attempt to set up a very simple model, we rely on the experimental feature saying that the $\beta$ band is largely decoupled from the ground as well as from the $\gamma$ bands and choose a model Hamiltonian whose

Table 13. The same as in table 6, but for ${ }^{166} \mathrm{Hf}$. The experimental data are taken from [64].

| ${ }^{166} \mathrm{Hf}$ | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2_{g}^{+}$ | 159 | 164 | 164 | 152 | 149 | 177 |
| $4{ }_{8}^{+}$ | 470 | 476 | 476 | 471 | 458 | 488 |
| $6_{8}^{+}$ | 897 | 891 | 890 | 906 | 883 | 897 |
| $8{ }_{8}^{+}$ | 1406 | 1391 | 1392 | 1415 | 1392 | 1385 |
| $10_{g}^{+}$ | 1972 | 1973 | 1975 | 1973 | 1966 | 1943 |
| $\underline{12+}$ | 2566 | 2631 | 2635 | 2566 | 2588 | 2568 |
| $0_{\beta}^{+}$ | 1065 | 926 | 921 | 1064 | 1000 | 1098 |
| $2_{\beta}^{+}$ | 1219 | 1222 | 1215 | 1216 | 1286 | 1219 |
| $4_{\beta}^{+}$ |  | 1753 | 1745 | 1536 | 1761 | 1490 |
| $6_{\beta}^{+}$ |  | 2419 | 2410 | 1970 | 2344 | 1870 |
| $8_{\beta}^{+}$ |  | 3189 | 3178 | 2479 | 3002 | 2342 |
| $10_{\beta}^{+}$ |  | 4049 | 4038 | 3038 | 3713 | 2893 |
| $2_{\gamma}^{+}$ | 810 | 862 | 862 | 854 | 864 | 899 |
| $3_{\gamma}^{+}$ | 1007 | 1004 | 1003 | 997 | 1007 | 1011 |
| $4_{\gamma}^{+}$ |  | 1174 | 1174 | 1177 | 1178 | 1160 |
| $5_{\gamma}^{+}$ | 1419 | 1370 | 1370 | 1385 | 1364 | 1330 |
| $6_{\gamma}^{+}$ |  | 1589 | 1588 | 1617 | 1611 | 1535 |
| $7{ }_{\gamma}^{+}$ |  | 1829 | 1829 | 1867 | 1822 | 1748 |
| $8_{\gamma}^{+}$ |  | 2090 | 2090 | 2133 | 2132 | 2002 |
| $9_{\gamma}^{+}$ |  | 2370 | 2372 | 2411 | 2357 | 2251 |
| $10_{\gamma}^{+}$ |  | 2671 | 2673 | 2701 | 2720 | 2550 |
| rms (keV) |  | 51 | 53 | 18 | 38 | 39 |

matrix elements between beta states and states belonging either to the ground or to the gamma band are all equal to zero. The simplest Hamiltonian obeying this restriction is

$$
\begin{equation*}
H=A_{1}\left(22 \hat{N}+5 \Omega_{\beta^{\prime}}^{\dagger} \Omega_{\beta^{\prime}}\right)+A_{2} \hat{J}^{2}+A_{3} \Omega_{\beta}^{\dagger} \Omega_{\beta} \tag{3.5}
\end{equation*}
$$

where $\hat{N}$ is the boson number, $\hat{J}^{2}$-angular momentum squared and $\Omega_{\beta^{\prime}}^{\dagger}$ denotes

$$
\begin{equation*}
\Omega_{\beta^{\prime}}^{\dagger}=\left(b^{\dagger} b^{\dagger}\right)_{00}-\frac{d^{2}}{\sqrt{5}} . \tag{3.6}
\end{equation*}
$$

Higher order terms in boson operators can be added to the Hamiltonian $H$ without altering the decoupling condition for the beta band. An example of this kind is the correction

$$
\begin{equation*}
\Delta H=A_{4}\left(\Omega_{\beta}^{\dagger} \Omega_{\beta^{\prime}}^{2}+\text { h.c. }\right)+A_{5} \Omega_{\beta^{\prime}}^{\dagger 2} \Omega_{\beta^{\prime}}^{2} \tag{3.7}
\end{equation*}
$$

The energies for beta band and for the gamma band states of odd angular momentum are described as average values of $H$ (3.5), or $H+\Delta H$ on $\varphi_{J M}^{\beta}$ and $\varphi_{J M}^{\gamma}$ (J-odd), respectively. As for the energies for the ground band and those of gamma band states with even angular momentum, they are obtained by diagonalizing a $2 \times 2$ matrix for each $J$.

The quadrupole transition operator is considered to be a sum of a linear term in bosons and one which is quadratic in the quadrupole bosons:

$$
\begin{equation*}
Q_{2 \mu}=q_{1}\left(b_{2 \mu}^{\dagger}+(-)^{\mu} b_{2,-\mu}\right)+q_{2}\left(\left(b^{\dagger} b^{\dagger}\right)_{2 \mu}+(b b)_{2 \mu}\right)+q_{3}\left(b^{\dagger} b\right)_{2 \mu} . \tag{3.8}
\end{equation*}
$$

Note that if $q_{3}=2 q_{2}$, the quadrupole transition operator can be obtained from the quadrupole transition operator expressed in terms of the collective quadruple coordinates $\alpha_{2 \mu}$ :

$$
\begin{equation*}
Q_{2 \mu}=Q_{1} \alpha_{2 \mu}+Q_{1}^{\prime}\left(\alpha_{2} \alpha_{2}\right)_{2 \mu} \tag{3.9}
\end{equation*}
$$

Table 14. The same as in table 6, but for ${ }^{168} \mathrm{Hf}$. The experimental data are taken from [65].

| ${ }^{168} \mathrm{Hf}$ | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2_{8}^{+}$ | 124 | 141 | 140 | 120 | 108 | 128 |
| $4_{8}^{+}$ | 386 | 409 | 409 | 382 | 351 | 389 |
| $6_{8}^{+}$ | 757 | 765 | 769 | 757 | 710 | 756 |
| $8{ }_{8}^{+}$ | 1214 | 1195 | 1207 | 1215 | 1172 | 1206 |
| $10_{g}^{+}$ | 1736 | 1694 | 1720 | 1736 | 1723 | 1730 |
| $\underline{12+}$ | 2306 | 2259 | 2303 | 2307 | 2354 | 2320 |
| $0_{\beta}^{+}$ | 942 | 795 | 755 | 928 | 878 | 942 |
| $2_{\beta}^{+}$ | 1059 | 1049 | 1002 | 1048 | 1039 | 1049 |
| $4_{\beta}^{+}$ | 1285 | 1505 | 1449 | 1310 | 1368 | 1285 |
| $6_{\beta}^{+}$ |  | 2077 | 2015 | 1684 | 1823 | 1630 |
| $8_{\beta}^{+}$ |  | 2738 | 2672 | 2143 | 2380 | 2068 |
| $10_{\beta}^{+}$ |  | 3477 | 3412 | 2664 | 3024 | 2587 |
| $2_{\gamma}^{+}$ | 876 | 911 | 906 | 902 | 928 | 939 |
| $3_{\gamma}^{+}$ | 1031 | 1033 | 1028 | 1020 | 1042 | 1035 |
| $4_{\gamma}^{+}$ | 1161 | 1179 | 1178 | 1172 | 1171 | 1161 |
| $5_{\gamma}^{+}$ | 1386 | 1347 | 1350 | 1353 | 1334 | 1311 |
| $6_{\gamma}^{+}$ | 1551 | 1535 | 1543 | 1558 | 1530 | 1492 |
| $7{ }_{\gamma}^{+}$ |  | 1741 | 1755 | 1786 | 1733 | 1687 |
| $8_{\gamma}^{+}$ |  | 1965 | 1988 | 2033 | 1992 | 1916 |
| $9_{\gamma}^{+}$ |  | 2206 | 2239 | 2297 | 2226 | 2148 |
| $10_{\gamma}^{+}$ |  | 2464 | 2508 | 2576 | 2543 | 2421 |
| rms (keV) |  | 75 | 70 | 15 | 43 | 31 |

The anharmonic term in the above expression can be obtained by expanding the deformed mean field around the spherical equilibrium shape $[50,51]$ of the nuclear surface. For the near vibrational regime, the interband matrix elements of the $q_{3}$ term vanish within the CSM [5]. Moreover, a transition operator depending on two free parameters seems to be suitable for describing the $E 2$ transition probabilities in several regions of the nuclide's chart [52].

Using the Rose convention [47], the reduced probability for the $E 2$ transition $J_{i}^{+} \rightarrow J_{f}^{+}$ can be expressed as

$$
\begin{equation*}
B\left(E 2 ; J_{i}^{+} \rightarrow J_{f}^{+}\right)=\left(\left\langle J_{i}^{+}\left\|Q_{2}\right\| J_{f}^{+}\right\rangle\right)^{2} \tag{3.10}
\end{equation*}
$$

Three specific features of the CSM are worth mentioning.
(a) The model states are generated through projection from a coherent state and two excitations of that through simple polynomial boson operators. Thus, it is expected that the projected states may account for the semiclassical behavior of the nuclear system staying in a state of high spin.
(b) The states are an infinite series of bosons and thus highly deformed states can be described.
(c) The model Hamiltonian is not commuting with the boson number operator and because of this property a basis generated from a coherent state is expected to be most suitable.
The CSM has been successfully applied to several nuclei exhibiting various equilibrium shapes which, according to the IBA classification, exhibit the $S O(6), S U(5)$ and $S U(3)$ symmetries. Several improvements of the CSM have been proposed by considering additional degrees of freedom such as isospin [6], quasi-particle [7] or collective octupole coordinates

Table 15. The same as in table 6, but for ${ }^{170} \mathrm{~W}$. The experimental data are taken from [66].

| ${ }^{170} \mathrm{~W}$ | Exp. | $X(5)$ | ISW | D | SSA | $\boxed{ }$ CSM |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: |
| $2_{g}^{+}$ | 157 | 145 | 133 | 145 | 151 | 171 |
| $4_{g}^{+}$ | 462 | 420 | 398 | 447 | 446 | 475 |
| $6_{g}^{+}$ | 876 | 785 | 767 | 858 | 844 | 873 |
| $8_{g}^{+}$ | 1363 | 1226 | 1228 | 1347 | 1323 | 1346 |
| $10_{g}^{+}$ | 1902 | 1739 | 1777 | 1894 | 1869 | 1882 |
| $12_{g}^{+}$ | 2464 | 2319 | 2411 | 2489 | 2471 | 2477 |
| $14_{g}^{+}$ | 3118 | 2965 | 3128 | 3126 | 3124 | 3128 |
| $16_{g}^{+}$ | 3816 | 3677 | 3927 | 3801 | 3821 | 3831 |
| $0_{\beta}^{+}$ |  | 816 | 587 | 760 | 507 | 823 |
| $2_{\beta}^{+}$ | 953 | 1077 | 804 | 905 | 790 | 953 |
| $4_{\beta}^{+}$ | 1202 | 1545 | 1208 | 1207 | 1204 | 1215 |
| $6_{\beta}^{+}$ | 1578 | 2132 | 1736 | 1618 | 1706 | 1578 |
| $8_{\beta}^{+}$ |  | 2810 | 2367 | 2107 | 2277 | 2020 |
| $10_{\beta}^{+}$ |  | 3568 | 3093 | 2654 | 2905 | 2531 |
| $2_{\gamma}^{+}$ | 937 | 944 | 945 | 928 | 936 | 965 |
| $3_{\gamma}^{+}$ | 1074 | 1068 | 1068 | 1066 | 1064 | 1074 |
| $4_{\gamma}^{+}$ | 1220 | 1219 | 1219 | 1238 | 1231 | 1217 |
| $5_{\gamma}^{+}$ |  | 1391 | 1397 | 1438 | 1400 | 1381 |
| $6_{\gamma}^{+}$ |  | 1584 | 1600 | 1662 | 1630 | 1578 |
| $7_{\gamma}^{+}$ |  | 1796 | 1828 | 1906 | 1828 | 1783 |
| $8_{\gamma}^{+}$ |  | 2025 | 2080 | 2168 | 2109 | 2027 |
| $9_{\gamma}^{+}$ |  | 2273 | 2355 | 2446 | 2329 | 2264 |
| $10_{\gamma}^{+}$ | 2538 | 2653 | 2739 | 2655 | 2550 |  |
| $\mathrm{rms}(\mathrm{keV})$ |  | 200 | 90 | 21 | 58 | 13 |

[8, 9]. The CSM has also been used to describe some nonaxial nuclei [49], and the results were compared with those obtained with the rotationvibration model [2]. A review of the CSM achievements is found in [10]. The terms involved in the model Hamiltonians used in by the CSM [5] and its generalized version [6] have microscopic counterparts as shown in [11] and [12], respectively.

## 4. Numerical results

### 4.1. Parameters

The parameters which define the energies and the $E 2$ transition probabilities of the models $X(5)$, ISW, D, SSA and CSM were fitted by the least-squares method for ten nuclei: ${ }^{176,178,180,188,190} \mathrm{Os},{ }^{150} \mathrm{Nd},{ }^{156} \mathrm{Dy},{ }^{166,168} \mathrm{Hf}$ and ${ }^{170} \mathrm{~W}$. In the least-squares procedure, all experimental energies were considered. The resulting values are those given in tables $1-5$. For the first three and the last three nuclei from table 1, the parameter $t^{\prime}$ cannot be determined since the corresponding term from the transition operator does not contribute to the intraband decays.

Some parameters vary by a large amount from one isotope to another but the relative variation is small. For example in the case of Os isotopes, the parameters could be interpolated

Table 16. The reduced $E 2$ transition probabilities determined with the $X(5)$, ISW, D, SSA and CSM models for the ${ }^{176} \mathrm{Os}$ nucleus are compared with the corresponding experimental data taken from [54].

| $\mathrm{B}(\mathrm{E} 2)$ (W.u.) | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | ---: | ---: | ---: | :--- | ---: | ---: |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $144_{-5}^{+5}$ | 167 | 127 | 145 | 136 | 144 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $243_{-5}^{+5}$ | 264 | 224 | 228 | 227 | 253 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $305_{-11}^{+11}$ | 330 | 305 | 292 | 297 | 328 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $321_{-14}^{+15}$ | 379 | 377 | 360 | 366 | 393 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | $441_{-63}^{+88}$ | 419 | 438 | 433 | 435 | 452 |
| $12_{g}^{+} \rightarrow 10_{g}^{+}$ | $517_{-146}^{+336}$ | 450 | 490 | 510 | 504 | 517 |

Table 17. The same as in table 16, but for ${ }^{178}$ Os. The experimental data are taken from [54-56]

| $\mathrm{B}(\mathrm{E} 2)$ (W.u.) | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | 138 | 147 | 137 | 146 | 141 | 138 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | 226 | 232 | 225 | 226 | 226 | 227 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | 290 | 291 | 287 | 280 | 283 | 282 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | 327 | 334 | 337 | 332 | 334 | 327 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | 384 | 369 | 378 | 384 | 382 | 368 |

Table 18. The same as in table 16, but for ${ }^{180} \mathrm{Os}$. The experimental data are taken from [58].

| B(E2)(W.u.) | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $120_{-30}^{+30}$ | 70 | 152 | 148 | 151 | 150 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $193_{-25}^{+25}$ | 111 | 167 | 177 | 172 | 149 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $160_{-40}^{+40}$ | 139 | 132 | 139 | 135 | 120 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $63_{-13}^{+13}$ | 160 | 95 | 83 | 90 | 96 |

by smooth curves. One parameter falls aside namely those of ${ }^{188} \mathrm{Os}$, which seems to achieve the critical point of the shape transition, i.e. exhibits an $X(5)$ behavior.

We note that the parameter $F$ involves the average value $\left\langle\beta^{2}\right\rangle$ which, in principle, is an angular-momentum-dependent quantity. Therefore, the differential equation in $\gamma$ should be iteratively solved, at each step the inserted average value being calculated with the wavefunction provided in the previous step. When the convergence of the process is met, one keeps the average value for the chosen angular momentum. Here, $\left\langle\beta^{2}\right\rangle$ was kept constant. Whether this hypothesis is valid or not can be posterity checked. To this goal, we represented in figure 1, the average $\left\langle\beta^{2}\right\rangle$ for each of the models ISW, D and SSA. We note that the average value only slightly depends on $J$ and that is especially true for ISW and SSA. If the limit of $\left\langle\beta^{2}\right\rangle$, when the convergence of the iterations mentioned above is reached, depends on $J$ like the averages shown in figure 1 , one could say that keeping $\left\langle\beta^{2}\right\rangle$ constant one ignores a slight decrease of energy with angular momentum.

With the parameters listed above the potentials in the variables $\beta$ and $\gamma$ and the wavefunctions describing the low lying states from the ground, beta and gamma bands respectively, are represented for four nuclei, as shown in figures 2-5. Analyzing these figures, several features can be noted. The $\beta$ potential has a deformed minimum located at a deformation which differs from one nucleus to another. The wavefunctions in $\beta$ for $0_{g}^{+}$and $2_{\gamma}^{+}$are almost


Figure 1. The average values of $\beta^{2}$ versus the angular momentum calculated within the ISW (panel a)), the D (panel b)) and the SSA (panel c)) models.
identical and have only one maximum and no node, while the band for $0_{\beta}^{+}$has one node, one maximum and one minimum. The maximum of the $|\phi|^{2}$ distribution for the three states represented in the quoted figures is achieved at a point which is close to the potential minimum. If $|\phi|^{2}$ is multiplied with the integration measure over $\beta$, then the probability distribution has a maximum closer to the potential minimum. The state $0_{\beta}^{+}$is characterized by two maxima for the probability distribution of the beta variable. This feature reflects the specific structure of the excitation operator of this state, from the ground state, i.e. $n_{\beta}=1$. The behavior of the wavefunctions in the variable $\gamma$ is mainly determined by the discontinuity for $\gamma=0$ and $\gamma=\frac{\pi}{3}$. The potential has two minima, one well pronounced near the first wall and one very flat close to the $\gamma=\frac{\pi}{3}$ discontinuity. Due to this structure, the wavefunction describing a state in the ground band has two maxima located above the mentioned minima. The state $2_{\gamma}^{+}$ heading the gamma band has an additional maximum.

### 4.2. Energies

The spectra of the chosen nuclei, determined by the models $X(5)$, ISW, D, SSA and CSM, are compared with the corresponding experimental data in tables $6-15$. The quality of agreement between the results of our calculations and the corresponding experimental data is given by


Figure 2. The solutions for the equation in $\beta$, corresponding to various angular momenta and the potential from the left-bottom panel, are plotted, in panel left-up, as a function of $\beta$. Similarly, on the right column the wavefunctions for $\gamma$ for different angular momenta and the effective potential shown in the right-bottom panel are plotted as a function of $\gamma$. The results correspond to ${ }^{150} \mathrm{Nd}$.


Figure 3. The same as in figure 2, but for ${ }^{166} \mathrm{Hf}$.
the rms values of the deviations. Thus, comparing the rms values corresponding to different models we conclude that for ${ }^{180} \mathrm{Os},{ }^{150} \mathrm{Nd}$ and ${ }^{170} \mathrm{~W}$ the best description of the spectra is that given by the CSM approach, energies of ${ }^{188}$ Os calculated with the SSA are closest to the experimental ones, while for the remaining nuclei the D formalism provides the most realist picture.

Using the experimental data listed in tables 6-15, one can calculate the ratio of the excitation energies for the states $4_{g}^{+}$and $2_{g}^{+}$, denoted by $R_{4_{g}^{+} / 2_{g}^{+}}$. The results are $2.93\left({ }^{176} \mathrm{Os}\right.$, ${ }^{190} \mathrm{Os},{ }^{150} \mathrm{Nd}$, ${ }^{156} \mathrm{Dy}$ ), $2.94\left({ }^{170} \mathrm{~W}\right)$, $2.96\left({ }^{166} \mathrm{Hf}\right), 3.02\left({ }^{178} \mathrm{Os}\right), 3.08\left({ }^{188} \mathrm{Os}\right), 3.10\left({ }^{180} \mathrm{Os}\right)$ and $3.11\left({ }^{168} \mathrm{Hf}\right)$. We note that all nuclei are characterized by a ratio $R_{4_{8}^{+} / 2_{8}^{+}}$which is close to the


Figure 4. The same as in figure 2, but for ${ }^{188} \mathrm{Os}$.


Figure 5. The same as in figure 2, but for ${ }^{190} \mathrm{Os}$.
value of 2.9 assigned to the critical point of the transition $S U(5) \rightarrow S U(3)$, which is described by the solvable model called $X(5)$. Despite this, the $X(5)$ approach provides a description which is worse than those obtained with the other models proposed here.

### 4.3. Reduced transition probabilities

As mentioned before, the parameters involved in the transition operators employed by different models have been fixed by fitting through a least-squares procedure the existent data. With the fitted parameter, the results for the reduced $E 2$ transition probabilities are presented in tables 16-25 where one also gives for comparison the available experimental data. For the lightest three isotopes of Os as well as for ${ }^{166,168} \mathrm{Hf}$ and ${ }^{170} \mathrm{~W}$, the available experimental data

Table 19. The same as in table 16, but for ${ }^{188}$ Os. The experimental data are taken from [59].

| $\mathrm{B}(\mathrm{E} 2)$ (W.u.) | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | :--- | :---: | :---: | :---: | :---: | :---: |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $79_{-2}^{+2}$ | 74 | 72 | 79 | 82 | 42 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $133_{-8}^{+8}$ | 118 | 115 | 121 | 123 | 87 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $138_{-8}^{+8}$ | 147 | 144 | 147 | 145 | 125 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $161_{-11}^{+11}$ | 169 | 166 | 174 | 162 | 161 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | $188_{-25}^{+25}$ | 187 | 184 | 203 | 178 | 195 |
| $0_{\beta}^{+} \rightarrow 2_{g}^{+}$ | $0.95_{-0.08}^{+0.08}$ | 47 | 48 | 33 | 21 | 0.95 |
| $0_{\beta}^{+} \rightarrow 2_{\gamma}^{+}$ | $4.3_{-0.5}^{+0.5}$ | 5.2 | 5.2 | 1.9 | 1.5 | 44 |
| $4_{\gamma}^{+} \rightarrow 2_{\gamma}^{+}$ | $47_{-10}^{+10}$ | 47 | 50 | 52 | 56 | 14 |
| $4_{\gamma}^{+} \rightarrow 3_{\gamma}^{+}$ | $320_{-120}^{+120}$ | 112 | 117 | 120 | 132 | 43 |
| $6_{\gamma}^{+} \rightarrow 4_{\gamma}^{+}$ | $70_{-30}^{+30}$ | 107 | 111 | 114 | 118 | 31 |
| $2_{\gamma}^{+} \rightarrow 0_{g}^{+}$ | $5_{-0.6}^{+0.6}$ | 8.4 | 10.9 | 10.8 | 9.9 | 5 |
| $2_{\gamma}^{+} \rightarrow 2_{g}^{+}$ | $16_{-2}^{+2}$ | 13 | 17 | 16 | 14 | 10.4 |
| $2_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | $34_{-5}^{+5}$ | 0.65 | 0.85 | 0.80 | 0.73 | 1.4 |
| $4_{\gamma}^{+} \rightarrow 2_{g}^{+}$ | $1.29_{-0.19}^{+0.19}$ | 5.7 | 7.1 | 6.7 | 6.1 | 1.7 |
| $4_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | $19_{-3}^{+3}$ | 18 | 23 | 20 | 19 | 10.7 |
| $4_{\gamma}^{+} \rightarrow 6_{g}^{+}$ | $16_{-7}^{+7}$ | 2 | 2 | 2 | 2 | 5 |
| $6_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | $0.21_{-0.11}^{+0.11}$ | 5.3 | 6.4 | 5.8 | 5.3 | 0.9 |
| $6_{\gamma}^{+} \rightarrow 6_{g}^{+}$ | $>9.4$ | 21 | 25 | 23 | 20 | 8.3 |

Table 20. The same as in table 16, but for ${ }^{190}$ Os. The experimental data are taken from [60].

| $\mathrm{B}(\mathrm{E} 2)(\mathrm{W} . \mathrm{u})$. | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | :--- | :---: | :---: | :---: | :---: | :---: |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $72_{-2}^{+2}$ | 58 | 57 | 56 | 61 | 45 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $105_{-6}^{+6}$ | 91 | 91 | 88 | 94 | 83 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $113_{-10}^{+10}$ | 115 | 113 | 112 | 112 | 112 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $137_{-20}^{+20}$ | 131 | 130 | 138 | 126 | 137 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | $120_{-30}^{+30}$ | 145 | 143 | 165 | 139 | 160 |
| $0_{\beta}^{+} \rightarrow 2_{g}^{+}$ | $2.2_{-0.5}^{+0.5}$ | 36 | 36 | 30 | 19 | 2.2 |
| $0_{\beta}^{+} \rightarrow 2_{\gamma}^{+}$ | $23_{-7}^{+7}$ | 8.9 | 9 | 8 | 5 | 148 |
| $4_{\gamma}^{+} \rightarrow 2_{\gamma}^{+}$ | $53_{-5}^{+5}$ | 36 | 38 | 37 | 41 | 20.4 |
| $4_{\gamma}^{+} \rightarrow 3_{\gamma}^{+}$ | $65_{-13}^{+13}$ | 87 | 90 | 87 | 98 | 84 |
| $6_{\gamma}^{+} \rightarrow 4_{\gamma}^{+}$ | $65_{-13}^{+13}$ | 83 | 85 | 84 | 89 | 49 |
| $8_{\gamma}^{+} \rightarrow 6_{\gamma}^{+}$ | $61_{-16}^{+16}$ | 112 | 113 | 119 | 115 | 72 |
| $2_{\gamma}^{+} \rightarrow 0_{g}^{+}$ | $5.9_{-0.6}^{+0.6}$ | 14.2 | 15.6 | 15.9 | 16.2 | 14 |
| $2_{\gamma}^{+} \rightarrow 2_{g}^{+}$ | $33_{-4}^{+4}$ | 21 | 24 | 24 | 24 | 33 |
| $4_{\gamma}^{+} \rightarrow 2_{g}^{+}$ | $0.68_{-0.06}^{+0.06}$ | 9.7 | 10.3 | 10.4 | 10.3 | 4.3 |
| $4_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | $30_{-4}^{+4}$ | 31 | 33 | 33 | 32 | 31 |
| $6_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | $<0.8$ | 9 | 10 | 10 | 9 | 1.7 |
| $6_{\gamma}^{+} \rightarrow 6_{g}^{+}$ | $31_{-8}^{+8}$ | 36 | 38 | 40 | 36 | 26 |

refer to the states of ground band. Agreements with the experimental data showed up by the five theoretical models are comparable in quality.

For ${ }^{156} \mathrm{Dy}$, besides the intraband transitions in the ground band, a few interband transitions from the gamma to the ground band are experimentally known. As seen from table 22, agreement between calculations with the experimental data is quite good.

Table 21. The same as in table 16, but for ${ }^{150} \mathrm{Nd}$. The experimental data are taken from [62].

| $\mathrm{B}(\mathrm{E} 2)(\mathrm{W} . \mathrm{u})$. | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | :--- | :--- | :---: | :---: | :---: | :---: |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $115_{-2}^{+2}$ | 107 | 104 | 92 | 116 | 81 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $182_{-2}^{+2}$ | 169 | 168 | 144 | 177 | 160 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $210_{-2}^{+2}$ | 212 | 210 | 183 | 211 | 222 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $278_{-25}^{+25}$ | 243 | 243 | 224 | 240 | 278 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | $204_{-12}^{+12}$ | 269 | 269 | 268 | 266 | 330 |
| $2_{\beta}^{+} \rightarrow 0_{\beta}^{+}$ | $114_{-23}^{+23}$ | 85 | 83 | 130 | 86 | 116 |
| $4_{\beta}^{+} \rightarrow 2_{\beta}^{+}$ | $170_{-51}^{+51}$ | 128 | 125 | 194 | 144 | 165 |
| $0_{\beta}^{+} \rightarrow 2_{g}^{+}$ | $39_{-2}^{+2}$ | 67 | 73 | 51 | 37 | 41.2 |
| $2_{\beta}^{+} \rightarrow 0_{g}^{+}$ | $1.2_{-0.2}^{+0.2}$ | 2.1 | 2.9 | 3.1 | 1.6 | 5.2 |
| $2_{\beta}^{+} \rightarrow 2_{g}^{+}$ | $9_{-2}^{+2}$ | 10 | 10 | 9 | 6 | 9 |
| $2_{\beta}^{+} \rightarrow 4_{g}^{+}$ | $17_{-3}^{+3}$ | 39 | 42 | 40 | 26 | 26 |
| $4_{\beta}^{+} \rightarrow 2_{g}^{+}$ | $0.12_{-0.02}^{+0.02}$ | 1.07 | 1.61 | 1.64 | 0.57 | 5.6 |
| $4_{\beta}^{+} \rightarrow 4_{g}^{+}$ | $7_{-1}^{+1}$ | 6 | 8 | 8 | 5 | 7.2 |
| $4_{\beta}^{+} \rightarrow 6_{g}^{+}$ | $70_{-13}^{+13}$ | 30 | 33 | 46 | 26 | 26 |
| $2_{\gamma}^{+} \rightarrow 0_{g}^{+}$ | $3_{-0.8}^{+0.8}$ | 2.4 | 8 | 9.8 | 5.1 | 16.3 |
| $2_{\gamma}^{+} \rightarrow 2_{g}^{+}$ | $5.4_{-1.7}^{+1.7}$ | 3.6 | 11.9 | 14.3 | 7.3 | 5.4 |
| $2_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | $2.6_{-2.0}^{+2.0}$ | 0.2 | 0.6 | 0.7 | 0.4 | 0.74 |
| $4_{\gamma}^{+} \rightarrow 2_{g}^{+}$ | $0.9_{-0.3}^{+0.3}$ | 1.6 | 5 | 6.1 | 3 | 28.6 |
| $4_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | $3.9_{-1.2}^{+1.2}$ | 5.3 | 15.5 | 18.9 | 9 | 9.6 |

Table 22. The same as in table 16, but for ${ }^{156} \mathrm{Dy}$. The experimental data are taken from [63].

| B(E2)(W.u.) | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | :--- | :--- | :--- | :--- | :--- | ---: |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $149.3_{-2.5}^{+2.5}$ | 142 | 138 | 111 | 137 | 66 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $261_{-17}^{+17}$ | 225 | 223 | 179 | 219 | 149 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $200_{-15}^{+15}$ | 282 | 279 | 235 | 271 | 221 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $289_{-14}^{+14}$ | 323 | 323 | 295 | 316 | 289 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | $366_{-25}^{+25}$ | 358 | 358 | 359 | 357 | 354 |
| $12_{g}^{+} \rightarrow 10_{g}^{+}$ | $382_{-22}^{+22}$ | 385 | 386 | 425 | 395 | 418 |
| $2_{\gamma}^{+} \rightarrow 0_{g}^{+}$ | $7.2_{-0.4}^{+0.4}$ | 6.6 | 9.9 | 23.3 | 11.6 | 7.2 |
| $2_{\gamma}^{+} \rightarrow 2_{g}^{+}$ | $9.4_{-1.0}^{+1.0}$ | 9.8 | 14.6 | 35.1 | 17.4 | 9.4 |
| $2_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | $12.6_{-1.9}^{+1.9}$ | 0.5 | 0.7 | 1.8 | 0.9 | 19.5 |

In [62], measured data in ${ }^{150} \mathrm{Nd}$ for intraband transitions ground to ground and beta to beta as well interband transitions to ground band have been reported. These data are described reasonably well by the five approaches as shown in table 21 . One remarks the good agreement obtained with the CSM approach. The largest discrepancies with the experimental data are obtained for the transitions $4_{\beta}^{+} \rightarrow 2_{g}^{+}$and $4_{\gamma}^{+} \rightarrow 2_{g}^{+}$which are overestimated by the theoretical results.

As for ${ }^{188,190} \mathrm{Os}$, the available data are about the intraband transitions ground to ground and gamma to gamma bands as well about the interband transition beta to ground and gamma to ground. They are compared with the results of our calculations in tables 19 and 20. Again, the agreement qualities obtained with the five sets of calculations are comparable with each other. The predictions for the decay probabilities of the transitions $4_{\gamma}^{+} \rightarrow 2_{g}^{+}$and $6_{\gamma}^{+} \rightarrow 4_{g}^{+}$

Table 23. The same as in table 16, but for ${ }^{166} \mathrm{Hf}$. The experimental data are taken from [64].

| B(E2)(W.u.) | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | :--- | :--- | ---: | :--- | ---: | ---: |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $128_{-7}^{+7}$ | 98 | 153 | 154 | 155 | 128 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $202_{-7}^{+7}$ | 155 | 212 | 216 | 215 | 203 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $221_{-13}^{+13}$ | 194 | 225 | 232 | 226 | 245 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $280_{-30}^{+30}$ | 223 | 225 | 230 | 225 | 280 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | $250_{-110}^{+400}$ | 246 | 220 | 219 | 218 | 311 |
| $12_{g}^{+} \rightarrow 10_{g}^{+}$ | $155_{-70}^{+50}$ | 265 | 213 | 199 | 209 | 351 |

Table 24. The same as in table 16, but for ${ }^{168} \mathrm{Hf}$. The experimental data are taken from [65].

| B(E2)(W.u.) | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $154_{-7}^{+7}$ | 141 | 165 | 176 | 175 | 154 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $244_{-12}^{+12}$ | 223 | 250 | 257 | 255 | 249 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $285_{-18}^{+18}$ | 279 | 294 | 292 | 291 | 304 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $350_{-50}^{+50}$ | 320 | 322 | 318 | 316 | 350 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | $370_{-60}^{+60}$ | 354 | 342 | 338 | 338 | 391 |
| $12_{g}^{+} \rightarrow 10_{g}^{+}$ | $320_{-120}^{+120}$ | 381 | 356 | 354 | 357 | 438 |

Table 25. The same as in table 16, but for ${ }^{170} \mathrm{~W}$. The experimental data are taken from [66].

| B(E2)(W.u.) | Exp. | $X(5)$ | ISW | D | SSA | CSM |
| :--- | :--- | ---: | :--- | :--- | :--- | :--- |
| $2_{g}^{+} \rightarrow 0_{g}^{+}$ | $124_{-3}^{+3}$ | 79 | 133 | 126 | 129 | 124 |
| $4_{g}^{+} \rightarrow 2_{g}^{+}$ | $179_{-18}^{+18}$ | 125 | 179 | 177 | 179 | 168 |
| $6_{g}^{+} \rightarrow 4_{g}^{+}$ | $189_{-14}^{+14}$ | 157 | 184 | 189 | 187 | 182 |
| $8_{g}^{+} \rightarrow 6_{g}^{+}$ | $190_{-50}^{+50}$ | 180 | 180 | 187 | 183 | 190 |
| $10_{g}^{+} \rightarrow 8_{g}^{+}$ | $170_{-40}^{+40}$ | 199 | 173 | 175 | 174 | 197 |
| $12_{g}^{+} \rightarrow 10_{g}^{+}$ | $160_{-30}^{+30}$ | 214 | 167 | 158 | 162 | 214 |

are larger than the corresponding experimental data. Also, the result for $0_{\beta}^{+} \rightarrow 2_{\gamma}^{+}$obtained within the CSM is about 6.5 larger than the corresponding experimental value. For some cases, the value of the $t_{2}$ obtained through the least-squares procedure is very large. The reason is as follows.

Within the SSA, the $t_{2}$ term of the transition operator contributes mainly to the interband transitions, while its matrix elements between states of a given band are very small. However, for the mentioned cases there are only few experimental data for interband transitions, most of the data referring to the intraband transitions. Consequently, the least-squares procedure uses small matrix elements of the intraband transitions which results in obtaining a huge number for $t_{2}$. An equally good description of these cases would be obtained by ignoring the $t_{2}$ term. We kept however this term just for the sake of having an unitary approach.

The results for the $E 2$ transitions raise the question: Why the models $X$ (5), ISW, D, SSA predict close results, although the states involved are described by different wavefunctions in the variables $\beta$ and $\gamma$ ? It seems that these differences are washed out by the fitting procedure adopted for the strengths of the transition operator. Moreover, the factor function depending
on the Euler angles are common in the mentioned four approaches, thus giving the dominant contribution to the reduced transition probability.

One signature for the triaxiality of the nuclear shape is the equality:

$$
\begin{equation*}
E_{2_{1}^{+}}+E_{2_{2}^{+}}=E_{3_{1}^{+}} . \tag{4.1}
\end{equation*}
$$

The departure from this rule, $\Delta E=\left|E_{2_{1}^{+}}+E_{2_{2}^{+}}-E_{3_{1}^{+}}\right|$, is equal to 2 and 11 keV for ${ }^{188} \mathrm{Os}$ and ${ }^{190} \mathrm{Os}$, respectively. The magnitude of these deviations was the argument for treating the two isotopes as triaxial nuclei [44]. On the other hand, the ratio $E_{4 / 2}$ amounts to 2.93 and 3.08 for ${ }^{188} \mathrm{Os}$ and ${ }^{190} \mathrm{Os}$, respectively, which are quite close to the specific value of $X(5)$ nuclei. Given these facts, we asked ourselves whether these nuclei are axially symmetric or behave like a triaxial rigid rotor. In order to answer this question, we compared the rms values of deviations for both energies and $B(E 2)$ values provided by the SMA and SSA approaches, respectively. Concerning the excitation energies in the three major bands, the rms values of prediction deviations from the corresponding experimental data yielded by the SMA for ${ }^{188} \mathrm{Os}$ and ${ }^{190} \mathrm{Os}$ are 24 and 32 keV , respectively, while the SSA results for these values being 13 and 27 keV , respectively. Therefore, regarding the excitation energies the two isotopes behave more like axially deformed nuclei. However, comparing the results for the reduced transition probabilities, it comes out that the triaxial rigid rotor behavior is favored. Indeed, the rms values for the SMA approach applied to ${ }^{188} \mathrm{Os}$ and ${ }^{190} \mathrm{Os}$ are 13 and $16 \mathrm{~W} . u$., respectively, while those corresponding to the SSA are 16 and 17 W.u., respectively. Remarkable is the fact that the differences of the rms values characterizing the two approaches, SMA and SSA, are quite small. Therefore, one could conclude that the two investigations, from [44] and from here, indicate that the two nuclei might be equally well described by both approaches.

## 5. Conclusions

Here, we summarize the main results obtained by this work. We selected ten nuclei characterized by a ratio $R_{4_{8}^{+} / 2_{8}^{+}}$close to 2.9 which is specific to the so-called $X(5)$ nuclei. Spectra of these nuclei are described by a new approach which treats the beta variable by the Schrödinger equation associated with a sextic oscillator plus a centrifugal potential. For the variable $\gamma$, one finds a differential equation which is satisfied by the spheroidal function. The excitation energies are obtained by summing the eigenvalues provided by the differential equations for the $\beta$ and $\gamma$ variables, respectively, while the corresponding functions are used to calculate the $E 2$ transition probabilities. The results are compared with the corresponding experimental data as well as with those obtained through other formalisms called $X(5)$, ISW, D and CSM which were previously used by the present authors to describe the spectroscopic properties of other $X(5)$ like nuclei.

Note that while the formalisms $X(5)$, ISW, D and SSA treat the energies and transition probabilities using the intrinsic coordinates and the rotation matrix function, the CSM is a quadrupole boson approach and therefore the mentioned observables are calculated with the collective coordinates which are specific to the laboratory frame.

A comparison of the rms values yielded by the five approaches shows that the D, CSM and SSA approaches produce the best agreement with the experimental energies. Concerning the $E 2$ transitions, one may show that all five sets of results quantitatively describe the experimental situation in a comparable manner with a slight advantage for SSA and CSM. Since the formalisms ISW, D, and SSA differ from each other by the way the variable beta is treated, otherwise the $\gamma$ equation being the same, the transition probabilities produced by the three approaches exhibit similar agreement with the experimental data. The SSA method produces very good agreement with the experimental energies for ${ }^{188} \mathrm{Os},{ }^{150} \mathrm{Nd}$ and ${ }^{168} \mathrm{Hf}$.

Table 5 shows that these nuclei have the largest deformations and moreover for the first two nuclei the ratio $R_{4^{+} / 2^{+}}$has the values 3.08 and 3.11, respectively, which deviate most from the $X(5)$ value. The quoted ratio for ${ }^{150} \mathrm{Nd}$ is 2.93 , which is close to the $X(5)$ value but its deformation is the largest one.

The sextic potential for the $\beta$ ensures a more realistic description of the excited states where the excitation of the beta degree of freedom is important. This is best seen in the excellent agreement of the calculated excitation energies in the beta and gamma bands with the corresponding experimental data.

The final conclusion is that the SSA, proposed in this paper, proves to be a suitable tool for a realistic description of the $X(5)$ like nuclei.

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## References

[1] Bohr A 1952 Mat.-Fys. Medd. K. Dan. Vidensk. Selsk. 161 Bohr A and Mottelson B 1953 Mat.-Fys. Medd. K. Dan. Vidensk. Selsk. 271
[2] Faessler A and Greiner W 1962 Z. Phys. 168425 Faessler A and Greiner W 1962 Z. Phys. 170105 Faessler A and Greiner W 1964 Z. Phys. 177190 Faessler A, Greiner W and Sheline R 1965 Nucl. Phys. 7033
[3] Gneuss G, Mosel U and Greiner W 1969 Phys. Lett. B 30397
[4] Hess P, Maruhn J and Greiner W 1981 Phys. Rev. C 232335 Hess P, Maruhn J and Greiner W 1981 J. Phys. G: Nucl. Phys. 7737
[5] Raduta A A, Ceausescu V, Gheorghe A and Dreizler R M 1981 Phys. Lett. B 99444 Raduta A A, Ceausescu V, Gheorghe A and Dreizler R M 1982 Nucl. Phys. A 381253
[6] Raduta A A, Ceausescu V and Faessler A 1987 Phys. Rev. C 362111
[7] Raduta A A, Lima C and Faessler A 1983 Z. Phys. A 31369
[8] Raduta A A, Raduta A1 H and Faessler A 1997 Phys. Rev. C 551747 Raduta A A, Ionescu D and Faessler A 2002 Phys. Rev. C 65064322
[9] Raduta A A and Sabac C 1983 Ann. Phys., NY 1481
[10] Raduta A A 2004 Recent Research Developments in Nuclear Physics vol 1 pp 1-70 ISBN:81-7895-124-X (managing editor: S G Pandalai, Transworld Research Network)
[11] Ceausescu V and Raduta A A 1976 Ann. Phys., NY 10094
[12] Raduta A A, Simkovic F and Faessler A 2000 J. Phys. G: Nucl. Part. Phys. 26793
[13] Wilets L and Jean M 1956 Phys. Rev. 102788
[14] Davydov A S and Filippov G F 1958 Nucl. Phys. 8237
[15] Arima A and Iachello F 1976 Ann. Phys., NY 99253 Arima A and Iachello F 1979 Ann. Phys., NY 123468
[16] Iachello F and Arima A 1987 The Interacting Boson Model (Cambridge: Cambridge University Press)
[17] Ginocchio J H and Kirson M W 1980 Phys. Rev. Lett. 441744
[18] Dieperink A E L, Scholten O and Iachello F 1980 Phys. Rev. Lett. 441767
[19] McCutchan E A, Zamfir N V and Casten R F 2004 Phys. Rev. C 69064306
[20] McCutchan E A and Zamfir N V 2005 Phys. Rev. C 71054306
[21] Iachello F 2000 Phys. Rev. Lett. 853580
[22] Iachello F 2001 Phys. Rev. Lett. 87052502
[23] Casten R F and Zamfir N V 2000 Phys. Rev. Lett. 853584
[24] Casten R F and Zamfir N V 2001 Phys. Rev. Lett. 87052503
[25] Zamfir N V et al 2002 Phys. Rev. C 65044325
[26] Clark R M et al 2004 Phys. Rev. C 69064322
[27] Zhang D-1 and Liu Y-x 2002 Phys. Rev. C 65057301
[28] Bonatsos D, Lenis D, Minkov N, Petrellis D, Raychev P P and Terziev P A 2003 arXiv:nucl-th/0312121v1 Bonatsos D, Lenis D, Minkov N, Petrellis D, Raychev P P and Terziev P A 2004 Phys. Lett. B 58440
[29] Davidson P M 1932 Proc. R. Soc. 135459
[30] Raduta A A, Gheorghe A and Faessler A 2005 J. Phys. G: Nucl. Part. Phys. 31337
[31] Leyvraz F and Heiss W D 2005 Phys. Rev. Lett. 95050402
[32] Gilmore R and Feng D H 1978 Nucl. Phys. A 301189
[33] Caprio M A, Cejnar P and Iachello F 2008 Ann. Phys., NY 3231106
[34] Fortunato L 2005 Eur. J. Phys. A 26 1-30
[35] Gheorghe A, Raduta A A and Ceausescu V 1978 Nucl. Phys. A 296228
[36] Raduta A A, Gheorghe A and Ceausescu V 1978 Nucl. Phys. A 311118
[37] Chacon E, Moshinski M and Sharp R T 1976 J. Math. Phys. 17668
[38] Corrigan T M, Margetan F J and Williams S A 1976 Phys. Rev. C 142279
[39] Margetan T M and Williams S A 1982 Phys. Rev. C 251602
[40] Gheorghe A, Raduta A A and Ceausescu V 1998 Nucl. Phys. A 637201
[41] Gheorghe A, Raduta A A and Faessler A 2007 Phys. Lett. B 648171
[42] Raduta A A, Gheorghe A C, Buganu P and Faessler A 2009 Nucl. Phys. A 81946
[43] Cejnar P, Jolie J and Casten R F 2010 Rev. Mod. Phys. 822155
[44] Raduta A A and Buganu P 2011 Phys. Rev. C 83034313
[45] Ushveridze A G 1994 Quasi-Exactly Solvable Models in Quantum Mechanics (Bristol: IOP)
[46] Lévai G and Arias J M 2004 Phys. Rev. C 69014304
[47] Rose M E 1957 Elementary Theory of Angular Momentum (New York: Wiley)
[48] Bijker R, Casten R F, Zamfir N V and McCutchan E A 2003 Phys. Rev. C 68064304
[49] Meyer U, Raduta A A and Faessler A 1998 Nucl. Phys. A 637321
[50] Raduta A A, Ceausescu V and Dreizler R M 1976 Nucl. Phys. A 27211
[51] Greiner W and Maruhn J Nuclear Models (Berlin: Springer) p 132 ISBN 3-540-59180-X
[52] Raduta A A, Budaca R and Faessler A 2012 Ann. Phys., NY 327 671-704
[53] Basunia M S 2006 Nucl. Data Sheets 107791
[54] Melon B 2011 Investigation of the X(5)-structure in ${ }^{176}$ Os using absolute transition probabilities InauguralDissertation zur Erlangung des Doktorgrades der Mathematisch-Naturwissenschaftlochen Facultät der Universität zu Köln
[55] Kibédi T, Dracoulis G D, Byrne A P and Davidson P M 1994 Nucl. Phys. A 567183
[56] Möller O et al 2005 Phys. Rev. C 72034306
[57] Achterberg E, Capurro O A and Marti G V 2009 Nucl. Data Sheets 1101473
[58] Wu S-C and Niu H 2003 Nucl. Data Sheets 100483
[59] Singh B 2002 Nucl. Data Sheets 95387
[60] Singh B 2003 Nucl. Data Sheets 99275
[61] Dermateosian E and Tuli J K 1995 Nucl. Data Sheets 75827
[62] Krücken R et al 2002 Phys. Rev. Lett. 88232501
[63] Reich C W 2003 Nucl. Data Sheets 99753
[64] Baglin C M 2008 Nucl. Data Sheets 1091103
[65] Baglin C M 2010 Nucl. Data Sheets 1111807
[66] Baglin C M 2002 Nucl. Data Sheets 96611

