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Application of the sextic oscillator with a centrifugal barrier and the spheroidal equation for some X(5) candidate nuclei

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Abstract

The eigenvalue equation associated with the Bohr-Mottelson Hamiltonian is considered in the intrinsic reference frame and amended by replacing the harmonic oscillator potential in the β variable with a sextic oscillator potential with centrifugal barrier plus a periodic potential for the γ variable. After the separation of variables, the β equation is quasi-exactly solved, while the solutions for the γ equation are just the angular spheroidal functions. An anharmonic transition operator is used to determine the reduced E2 transition probabilities. The formalism is conventionally called the sextic and spheroidal approach (SSA) and applied for several X(5) candidate nuclei: ^{176,178,180,188,190}Os, ¹⁵⁰Nd, ¹⁷⁰W, ¹⁵⁶Dy and ^{166,168}Hf. The SSA predictions are in good agreement with the experimental data of the mentioned nuclei. The comparison of the SSA results with those yielded by other models, such as X(5) (Iachello 2001 Phys. Rev. Lett. 87 052502), infinite square well (Raduta et al 2009 Nucl. Phys. A 819 46) and Davidson like potential (Raduta et al) for the β , otherwise keeping the spheroidal functions for the γ , and the coherent state model (Raduta et al 1981 Phys. Lett. B 99 444, Raduta et al 1982 Nucl. Phys. A 381 253, Raduta et al 1987 Phys. Rev. C 36 2111, Raduta et al 1983 Z. Phys. A 313 69, Raduta et al 1997 Phys. Rev. C 55 1747, Raduta et al 2002 Phys. Rev. C 65 064322, Raduta and Sabac 1983 Ann. Phys., NY 148 1, Raduta 2004 Recent Research Developments in Nuclear Physics vol 1) respectively, suggests that SSA represents a good approach to describe nuclei achieving the critical point of the $U(5) \rightarrow SU(3)$ shape phase transition.

(Some figures may appear in colour only in the online journal)

1. Introduction

Since the liquid drop model was developed [1], the quadrupole shape coordinates were widely used by both phenomenological and microscopic formalisms to describe the basic properties

of nuclear systems. Based on these coordinates, one defines quadrupole boson operators in terms of which model Hamiltonians and transition operators are defined. Since the original spherical harmonic liquid drop model was able to describe only a small amount of data for spherical nuclei, several improvements have been made. Thus, the Bohr-Mottelson model was generalized by Faessler and Greiner [2] in order to describe the small oscillations around a deformed shape which results in obtaining a flexible model, called vibration-rotation model, suitable for the description of deformed nuclei. Later on [3], this picture was extended by including anharmonicities as low-order invariant polynomials in the quadrupole coordinates. With a suitable choice of the parameters involved in the model Hamiltonian, the equipotential energy surface may exhibit several types of minima [4] such as spherical, deformed prolate, deformed oblate, deformed triaxial, etc. To each equilibrium shape, specific properties for excitation energies and electromagnetic transition probabilities show up. Due to this reason, one customarily says that static values of intrinsic coordinates determine a phase for the nuclear system. The boson description with a complex anharmonic Hamiltonian makes use of a large number of structure parameters which are to be fitted. A smaller number of parameters are used by the coherent state model (CSM) [5] which uses a restricted collective space generated through angular momentum projection by three deformed orthogonal functions of coherent type. The model is able to describe in a realistic fashion transitional and well-deformed nuclei of various shapes including states of high and very high angular momenta. Various extensions to include other degrees of freedom such as isospin [6], single particle [7] or octupole [8, 9] degrees of freedom have been formulated [10].

It has been noted that a given nuclear shape may be associated with a certain symmetry. Hence, its properties may be described with the help of the irreducible representation of the respective symmetry group. Thus, the gamma unstable nuclei can be described by the O(6)symmetry [13], the gamma-rigid triaxial rotor by the D2 symmetry [14], the symmetric rotor by the SU(3) symmetry and the spherical vibrator by the U(5) symmetry. Thus, even in the 1950s, the symmetry properties have been greatly appreciated. However, a big push forward was brought by the interacting boson approximation (IBA) [15, 16], which succeeded to describe the basic properties of a large number of nuclei in terms of the symmetries associated with a system of quadrupole (d) and monopole (s) bosons which generate the U(6) algebra of the IBA. The three limiting symmetries U(5), O(6) and SU(3) mentioned above in the context of the collective model are also dynamic symmetries for U(6). Moreover, for each of these symmetries a specific group reduction chain provides the quantum numbers characterizing the states, which are suitable for a certain region of nuclei. Besides the virtue of unifying the group theoretical descriptions of nuclei exhibiting different symmetries, the procedure defines very simple reference pictures for the limiting cases. For nuclei lying close to the region characterized by a certain symmetry, the perturbative corrections are to be included.

In [17, 18], it has been proved that on the U(5)–O(6) transition leg there exists a critical point for a second-order phase transition, while the U(5)–SU(3) leg has a first-order phase transition. Actually, the first-order phase transition takes place not only on the mentioned leg of the Casten's triangle, but covers all the interior of the triangle up to the second order [19]. Examples of such nuclei, falling inside the triangle, are the Os isotopes [20].

Recently, Iachello [21, 22] pointed out that these critical points correspond to distinct symmetries, namely E(5) and X(5), respectively. For the critical value of an ordering parameter, energies are given by the zeros of a Bessel function of half-integer and irrational indices, respectively.

The description of low lying states in terms of Bessel functions was used first by Jean and Willet [13], but the interesting feature saying that this is a critical picture in a phase transition and defines a new symmetry was indeed advanced first in [21].

Representatives for the two symmetries have been experimentally identified. To give an example, the relevant data for 134 Ba [23] and 152 Sm [24] suggest that they are close to the E(5) and X(5) symmetries, respectively. Another candidate for E(5) symmetry is 102 Pd [25, 27]. A systematic search for E(5) behavior in nuclei has been reported in [26].

In [30], we advanced the hypothesis that the critical point in a phase transition is state dependent. We tested this with a hybrid model for ¹³⁴Ba and ¹⁰⁴Ru. A similar property of the phase transition was investigated in the context of a schematic two-level model in [31, 32]. A rigorous analysis of the characteristics of excited state quantum phase transitions is performed in [33].

The departure from the γ unstable picture has been treated by several authors [28] whose contributions are reviewed by Fortunato in [34]. The difficulty in treating the γ degree of freedom consists in the fact that this variable is coupled to the rotation variables. A full solution for the Bohr–Mottelson Hamiltonian including an explicit treatment of γ deformation variable can be found in [35–39]. Therein, we treated separately also the γ unstable and the rotor Hamiltonian. A more complete study of the rotor Hamiltonian and the distinct phases associated with a tilted moving rotor is given in [40].

The treatment of the γ variable becomes even more complicated when we add to the liquid drop Hamiltonian a potential depending on β and γ at a time. To simplify the starting problem related to the inclusion of the γ variable, one uses model potentials which are sums of a beta and a γ depending potentials. In this way, the nice feature for the beta variable to be decoupled from the remaining four variables, specific to the harmonic liquid drop, is preserved. Further, the potential in γ is expanded either around to $\gamma=0$ or around $\gamma=\frac{\pi}{6}$. In the first case, if only the singular term is retained, one obtains the infinite square well (ISW) model described by Bessel functions in gamma. If the γ^2 term is added to this term, then the Laguerre functions are the eigenstates of the approximated gamma depending Hamiltonian, which results in defining the functions characterizing the X(5) approach.

The drawback of these approximation consists in that the resulting γ depending functions are not periodic as the starting Hamiltonian is. Moreover, they are orthonormalized on unbound intervals although the underlying equation was derived under the condition of $|\gamma|$ small. The scalar product of the space of the resulting functions is not defined based on the measure $|\sin 3\gamma| d\gamma$ as happens in the liquid drop model. Under these circumstances it happens that the approximated Hamiltonian in γ looses its hermiticity.

In some earlier publications [41, 42], we proposed a scheme where the gamma variable is described by a solvable Hamiltonian whose eigenstates are spheroidal functions which are periodic. Here, we give details about the calculations and describe some new numerical applications. Moreover, the formalism was completed by treating the β variable by a Schrödinger equation associated with the Davidson potential. Alternatively, we considered the equation for a five-dimensional square well potential. We have shown that the new treatment of the gamma variable removes the drawbacks mentioned above and moreover brings a substantial improvement of the numerical analysis.

Here, we keep the description of the gamma variable by spheroidal functions and use a new potential for the beta variable which seems to be more suitable for a realistic description of more complex spectra. We call this approach as sextic and spheroidal approach (SSA). The potential is that of a sextic oscillator plus a centrifugal term which leads to a quasi-exactly solvable model. The resulting formalism will be applied to ten nuclei which were not included in our previous descriptions and moreover are suspected to be good candidate for exhibiting X(5) features having the ratio of excitation energies of the ground band members 4^+ and 2^+ close to the value of 2.9. The results of our calculations are compared with those obtained through other methods such as the ISW, D and CSM.

The goals presented in the previous section will be developed according to the following plan. In section 2, the main ingredients of the theoretical models X(5), ISW, D and SSA will be briefly presented. The CSM is separately described in section 3. Numerical results are given and commented in section 4, while the final conclusions are drawn in section 5.

2. The separation of variables and solutions

In order to describe the critical nuclei of the U(5)–SU(3) shape phase transition, we resort the Bohr–Mottelson Hamiltonian with a potential depending on both the β and γ variables:

$$H\psi(\beta, \gamma, \Omega) = E\psi(\beta, \gamma, \Omega), \tag{2.1}$$

where

$$H = -\frac{\hbar^2}{2B} \left[\frac{1}{\beta^4} \frac{\partial}{\partial \beta} \beta^4 \frac{\partial}{\partial \beta} + \frac{1}{\beta^2 \sin 3\gamma} \frac{\partial}{\partial \gamma} \sin 3\gamma \frac{\partial}{\partial \gamma} - \frac{1}{4\beta^2} \sum_{k=1}^3 \frac{\hat{Q}_k^2}{\sin^2 \left(\gamma - \frac{2\pi}{3}k\right)} \right] + V(\beta, \gamma). \tag{2.2}$$

Here, β and γ are the intrinsic deformation variables, Ω denotes the Euler angles θ_1 , θ_2 and θ_3 , \hat{Q}_k are the angular momentum components in the intrinsic reference frame, while B is the so-called mass parameter.

2.1. The separation of variables

To achieve the separation of variables in equation (2.1), some approximations are necessary. Choosing the potential energy in the form [13, 34]

$$V(\beta, \gamma) = V_1(\beta) + \frac{V_2(\gamma)}{\beta^2},\tag{2.3}$$

the β variable is separated from the γ and the Euler angles Ω , which are still coupled due to the rotational term:

$$W = \frac{1}{4} \sum_{k=1}^{3} \frac{\hat{Q}_{k}^{2}}{\sin^{2} \left(\gamma - \frac{2\pi}{3}k\right)}.$$
 (2.4)

Furthermore, the γ is separated from the Euler angles by using the second-order power expansion of the rotational term around the equilibrium value $\gamma_0 = 0^0$ (see equation (B.5) from [42]):

$$W \approx \frac{1}{3}\hat{Q}^2 + \left(\frac{1}{4\sin^2\gamma} - \frac{1}{3}\right)\hat{Q}_3^2 + \frac{2}{2\sqrt{3}}(\hat{Q}_2^2 - \hat{Q}_1^2)\gamma + \frac{2}{3}(\hat{Q}^2 - \hat{Q}_3^2)\gamma^2 + \mathcal{O}(\gamma^3), \tag{2.5}$$

and then averaging the result with the Wigner function $D_{MK}^{(L)}$:

$$\langle W \rangle = \frac{1}{3}L(L+1) + \left(\frac{1}{4\sin^2\gamma} - \frac{1}{3}\right)K^2 + \frac{2}{3}[L(L+1) - K^2]\gamma^2.$$
 (2.6)

The term L(L+1)/3 multiplied by $1/\beta^2$ is transferred to the equation for β ,

$$\left[-\frac{1}{\beta^4} \frac{\partial}{\partial \beta} \beta^4 \frac{\partial}{\partial \beta} + \frac{L(L+1)}{3\beta^2} + v_1(\beta) \right] f(\beta) = \varepsilon_{\beta} f(\beta), \tag{2.7}$$

while the sum of remaining terms, denoted with $\tilde{V}(\gamma, L, K)$, is kept in the equation for γ :

$$\left[-\frac{1}{\sin 3\gamma} \frac{\partial}{\partial \gamma} \sin 3\gamma \frac{\partial}{\partial \gamma} + \tilde{V}(\gamma, L, K) + v_2(\gamma) \right] \eta(\gamma) = \tilde{\varepsilon}_{\gamma} \eta(\gamma). \tag{2.8}$$

In equations (2.7) and (2.8), the following notations were used:

$$v_1(\beta) = \frac{2B}{\hbar^2} V_1(\beta), \quad v_2(\gamma) = \frac{2B}{\hbar^2} V_2(\gamma), \quad \varepsilon_\beta = \frac{2B}{\hbar^2} E_\beta, \quad \tilde{\varepsilon}_\gamma = \langle \beta^2 \rangle \frac{2B}{\hbar^2} E_\gamma.$$
 (2.9)

Equations (2.7) and (2.8) are to be separately solved and finally the full solution of equation (2.1) is obtained by combining the contributions coming from each variable. In what follows, we shall give the necessary details for solving the above-mentioned equations.

2.2. Solutions of the β equation

Solutions of the β equation, corresponding to different potentials, were considered by several authors [34, 43]. Here, we mention only three of them, namely the ISW, the Davidson and the sextic potentials. Details about how to solve the β equation for these potentials can be found in [42, 44].

2.2.1. The infinite square well potential. The solution of the β equation with an ISW potential, having the expression

$$v_1(\beta) = \begin{cases} 0, & \beta \leqslant \beta_{\omega} \\ \infty, & \beta > \beta_{\omega} \end{cases}$$
 (2.10)

was first time given in [13] and then in [21, 22] for E(5) and X(5) models. The β wavefunctions are written in terms of the Bessel functions of half-integer [21] and irrational indices [22], respectively. The solution for X(5) is

$$f_{s,L}(\beta) = C_{s,L}\beta^{-\frac{3}{2}}J_{\nu}\left(\frac{x_{s,L}}{\beta_{\omega}}\beta\right), \quad \nu = \sqrt{\frac{L(L+1)}{3} + \frac{9}{4}}, \quad s = 1, 2, 3, \dots$$
 (2.11)

Here, $C_{s,L}$ is the normalization factor, which is determined from the condition

$$\int_0^{\beta_\omega} (f_{s,L}(\beta))^2 \beta^4 \, \mathrm{d}\beta = 1. \tag{2.12}$$

The corresponding eigenvalues are given in terms of the Bessel zeros $x_{s,L}$:

$$E_{\beta}(s,L) = \frac{\hbar^2}{2B} \left(\frac{x_{s,L}}{\beta_{\omega}}\right)^2. \tag{2.13}$$

2.2.2. The Davidson potential. Choosing in equation (2.7) a Davidson potential [29] of the form

$$v_1(\beta) = \beta^2 + \frac{\beta_0^4}{\beta^2},\tag{2.14}$$

solutions are the generalized Laguerre polynomials:

$$f_{n_{\beta},m_{\beta}}(\beta) = \sqrt{\frac{2n_{\beta}!}{\Gamma(n_{\beta} + m_{\beta} + 1)}} L_{n_{\beta}}^{m_{\beta}}(\beta^{2}) \beta^{m_{\beta} - \frac{3}{2}} e^{-\frac{\beta^{2}}{2}}, \quad m_{\beta} = \sqrt{\frac{L(L+1)}{3} + \frac{9}{4} + \beta_{0}^{4}}. \quad (2.15)$$

The wavefunctions, $f_{n_{\beta},m_{\beta}}(\beta)$, are normalized to unity with the integration measure $\beta^4 d\beta$. Energies have the following expression:

$$E_{\beta}(n_{\beta}, L) = \frac{\hbar^2}{2B} \left(2n_{\beta} + 1 + \sqrt{\frac{L(L+1)}{3} + \frac{9}{4} + \beta_0^4} \right), \quad n_{\beta} = 0, 1, 2, \dots, \quad n_{\beta} = s - 1.$$
(2.16)

2.2.3. The sextic oscillator potential with a centrifugal barrier. The solution of the β equation with a sextic potential, for critical nuclei of the $U(5) \rightarrow SU(3)$ shape phase transition, was obtained by taking into consideration the solution of the Schrödinger equation with a sextic potential given in [45] and applied to the E(5) like nuclei in [46] and to the triaxial nuclei in [44].

In order to reduce the β equation to the Schrödinger equation with a sextic potential [45], we rewrite the averaged rotational term, given by equation (2.6), in the following form:

$$\langle W \rangle = [L(L+1) - 2] + \left[2 - \frac{2}{3}L(L+1)\right] + \left(\frac{1}{4\sin^2\gamma} - \frac{1}{3}\right)K^2 + \frac{2}{3}[L(L+1) - K^2]\gamma^2.$$
(2.17)

As already mentioned, the first term of the above equation is added to the β equation, while the other terms remain in the γ equation. Making the substitution $f(\beta) = \beta^{-2} \varphi(\beta)$, we have

$$\left[-\frac{\partial^2}{\partial \beta^2} + \frac{L(L+1)}{\beta^2} + v_1(\beta) \right] \varphi(\beta) = \varepsilon_\beta \varphi(\beta). \tag{2.18}$$

The sextic potential is chosen such that to obtain the description from [44]:

$$v_1^{\pm}(\beta) = (b^2 - 4ac^{\pm})\beta^2 + 2ab\beta^4 + a^2\beta^6 + u_0^{\pm}, \quad c^{\pm} = \frac{L}{2} + \frac{5}{4} + M.$$
 (2.19)

Here, c is a constant which has two different values, one for L even and other for L odd:

$$(M, L): (k, 0); (k - 1, 2); (k - 2, 4); (k - 3, 6) \dots \Rightarrow c = k + \frac{5}{4} \equiv c^{+}(L\text{-even}),$$
 (2.20)

$$(M, L): (k, 1); (k - 1, 3); (k - 2, 5); (k - 3, 7) \dots \Rightarrow c = k + \frac{7}{4} \equiv c^{-}(L \text{-odd}).$$
 (2.21)

The constants u_0^{\pm} are fixed such that the potential for L odd has the same minimum energy as the potential for L even. The solutions of equation (2.18), with the potential given by equation (2.19), are

$$\varphi_{n_{\beta},L}^{(M)}(\beta) = N_{n_{\beta},L} P_{n_{\beta},L}^{(M)}(\beta^2) \beta^{L+1} e^{-\frac{a}{4}\beta^4 - \frac{b}{2}\beta^2}, \quad n_{\beta} = 0, 1, 2, \dots M,$$
 (2.22)

where $N_{n_{\beta},L}$ are the normalization factor, while $P_{n_{\beta},L}^{(M)}(\beta^2)$ are polynomials in x^2 of n_{β} order. The corresponding excitation energy is

$$E_{\beta}(n_{\beta}, L) = \frac{\hbar^2}{2B} \left[b(2L+3) + \lambda_{n_{\beta}}^{(M)}(L) + u_0^{\pm} \right], \quad n_{\beta} = 0, \quad 1, 2, \dots, M,$$
 (2.23)

where $\lambda_{n_s}^{(M)} = \varepsilon_{\beta} - u_0^{\pm} - 4bs$ is the eigenvalue of the equation:

$$\left[-\left(\frac{\partial^2}{\partial \beta^2} + \frac{4s - 1}{\beta} \frac{\partial}{\partial \beta} \right) + 2b\beta \frac{\partial}{\partial \beta} + 2a\beta^2 \left(\beta \frac{\partial}{\partial \beta} - 2M \right) \right] P_{n_{\beta},L}^{(M)}(\beta^2) = \lambda_{n_{\beta}}^{(M)} P_{n_{\beta},L}^{(M)}(\beta^2). \tag{2.24}$$

2.3. Solutions of the γ equation

2.3.1. The X(5) model. Within the X(5) model [22], devoted to the description of the critical point in the phase transition $SU(5) \rightarrow SU(3)$, the potential is a sum of an ISW in the β variable and a harmonic oscillator in the γ variable. For the rotational term and the other terms of the γ equation, the first-order Taylor expansion around $\gamma_0 = 0^0$ is considered, which results in obtaining for the γ variable the radial equation of a two-dimensional oscillator with the solution

$$\eta_{n_{\gamma},K}(\gamma) = C_{n,K} \gamma^{|K/2|} e^{-(3a)\gamma^2/2} L_n^{|K|} (3a\gamma^2), \quad n = \left(\frac{n_{\gamma} - |K|}{2}\right),$$
(2.25)

Table 1. The fitted values of the parameters involved in the expressions of the energies and transition probabilities of the X(5) model are given for each considered nucleus.

<i>X</i> (5)	¹⁷⁶ Os	¹⁷⁸ Os	¹⁸⁰ Os	¹⁸⁸ Os	¹⁹⁰ Os	¹⁵⁰ Nd	¹⁵⁶ Dy	¹⁶⁶ Hf	¹⁶⁸ Hf	¹⁷⁰ W
B_1 (keV)	18.08	18.13	18.79	25.56	26.92	17.77	17.02	23.46	20.14	20.68
X (keV)	822.28	818.68	880.10	452.71	438.53	966.50	950.46	698.15	770.26	799.14
$t [W.u.]^{1/2}$	1.29	1.22	0.84	0.86	0.76	1.03	1.19	0.99	1.19	0.89
$t' [W.u.]^{1/2}$	_	_	_	0.92	1.19	0.49	0.81	_	_	_

where $L_n^{|K|}$ are the generalized Laguerre polynomials. The eigenvalue of the γ equation has the following expression:

$$\varepsilon_{\gamma} = \frac{3a}{\sqrt{\langle \beta^2 \rangle}} (n_{\gamma} + 1) - \frac{(K/2)^2}{\langle \beta^2 \rangle} \frac{4}{3},\tag{2.26}$$

where a is a parameter characterizing the oscillator potential in the γ variable. The total energy and wavefunction are obtained by combining the results of all variables:

$$E(s, L, n_{\gamma}, K) = E_0 + B_1(x_{s,L})^2 + An_{\gamma} + CK^2,$$
(2.27)

$$\Psi(\beta, \gamma, \Omega) = \frac{1}{\sqrt{2(1 + \delta_{K.0})}} f_{s,L}(\beta) \left[\eta_{n_{\gamma},K}(\gamma) D_{M,K}^{L}(\Omega) + (-1)^{L+K} \eta_{n_{\gamma},-K}(\gamma) D_{M,-K}^{L}(\Omega) \right].$$
(2.28)

If the total energy (2.27) is normalized to the energy of the ground state, then we will have for the ground band and for the first beta band the expression

$$E(s, L, 0, 0) - E(1, 0, 0, 0) = B_1(x_{s,L}^2 - x_{1,0}^2), \quad s = 1, \quad 2; \quad L = 0, 2, 4, 6, \dots,$$
 (2.29)

while for the first γ band

$$E(s, L, 1, 2) - E(1, 0, 0, 0) = B_1(x_{1,L}^2 - x_{1,0}^2) + A + 4C, L = 2, 3, 4, 5, \dots$$
 (2.30)

One notes that the parameters A and C give contribution only to the γ band energies, and that these two parameters can be replaced with only one parameter, for example X = A + 4C. The total energy for the ground band and for the first β and γ bands, normalized to the energy of the ground state, can be written in the following form:

$$E(s, L, n_{\nu}, K) - E(1, 0, 0, 0) = B_1(x_{s, L}^2 - x_{1, 0}^2) + \delta_{K, 2}X.$$
(2.31)

Further, the parameters B_1 and X will be fitted by the least-squares procedure for each considered nucleus.

2.3.2. The ISW model. Within the ISW model, employed in this paper, the β equation is treated as in the X(5) model, using an ISW, while the γ equation is reduced to a spheroidal equation. The ISW model was proposed by one of the authors (AAR) and his collaborators in [41] and subsequently with more details and applications in [42]. Here, only the solutions will be presented. The potential $v_2(\gamma)$ was chosen such that a minimum in $\gamma = 0^0$ is achieved:

$$v_2(\gamma) = u_1 \cos 3\gamma + u_2 \cos^2 3\gamma. \tag{2.32}$$

This potential is renormalized by a contribution coming from the γ rotational term and consequently an effective reduced potential for the γ variable results

$$\tilde{v}_2(\gamma) = u_1 \cos 3\gamma + u_2 \cos^2 3\gamma + \frac{9}{4 \sin^3 3\gamma},$$
(2.33)

Table 2. The same as in table 1, but for the ISW model.

ISW	¹⁷⁶ Os	¹⁷⁸ Os	¹⁸⁰ Os	¹⁸⁸ Os	¹⁹⁰ Os	¹⁵⁰ Nd	¹⁵⁶ Dy	¹⁶⁶ Hf	¹⁶⁸ Hf	¹⁷⁰ W
B_1 (keV)	14.30	14.54	13.21	25.50	21.83	14.68	11.43	23.31	19.12	14.87
F (keV)	24.24	23.19	44.66	0.69	36.73	28.88	45.99	1.69	11.30	41.12
u_1	-159.24	-168.08	-36.729	-25000	-4999.35	-152.35	-12.55	-10000	-385.35	-44.36
u_2	0	0	0	0	2560.22	0	0	0	0	0
$t_1 [W.u.]^{1/2}$	-52.91	473.53	3302.3	503.11	419.67	538.99	591.54	1881.39	1197.94	1827.11
$t_2 [W.u.]^{1/2}$	-4305.14	-1323.6	14 304.2	-241.19	-48.09	-387.08	-468.57	8242.45	2702.98	6436.57

Table 3. The same as in table 1, but for the D model.

D	¹⁷⁶ Os	¹⁷⁸ Os	¹⁸⁰ Os	¹⁸⁸ Os	¹⁹⁰ Os	¹⁵⁰ Nd	¹⁵⁶ Dy	¹⁶⁶ Hf	¹⁶⁸ Hf	¹⁷⁰ W
E (keV)	316.34	317.31	334.32	559.76	462.44	369.50	324.08	532.22	463.88	379.93
F(keV)	38.41	37.33	39.01	28.48	42.45	26.48	33.11	11.87	25.87	37.72
β_0	1.64	1.56	1.61	1.98	1.64	1.71	1.45	1.79	2.02	1.63
u_1	-55.48	-57.20	-52.40	-7.70	-4098.61	-168.78	-58.16	-320.01	-130.49	-54.50
u_2	0	0	0	0	2167.18	0	0	0	0	0
$t_1 [W.u.]^{1/2}$	197.92	264.47	758.41	126.88	126.70	154.70	191.28	448.76	329.01	411.63
$t_2 [W.u.]^{1/2}$	-25.31	78.30	931.21	-17.09	-3.92	-25.31	-13.46	430.42	193.06	363.66

Table 4. The same as in table 1, but for the SSA model.

SSA	¹⁷⁶ Os	¹⁷⁸ Os	¹⁸⁰ Os	¹⁸⁸ Os	¹⁹⁰ Os	¹⁵⁰ Nd	¹⁵⁶ Dy	¹⁶⁶ Hf	¹⁶⁸ Hf	¹⁷⁰ W
E (keV)	0.99	0.46	1.46	2.53	5.29	0.75	0.91	1.82	0.54	0.31
F(keV)	2.67	3.12	1.69	11.31	5.55	3.87	1.93	15.97	1.99	2.84
a	951.49	4466.56	600.70	644.98	111.79	2636.48	1248.40	1205.13	7897.62	13197.99
b	126	279	50	27	15.8	88	87	46	32	341
u_1	-5607.45	-4048.06	-15000	-215.19	-452.74	-3877.84	-10000	-224.90	-9980.01	-4585.44
u_2	0	0	0	0	0	0	0	0	0	0
$t_1 [W.u.]^{1/2}$	376.70	2260.6	8541.32	1033.43	675.12	1754.26	1882.91	4759.23	3463.05	8901.59
$t_2 [W.u.]^{1/2}$	-32619.3	-22343.8	117781	-1022.41	32.73	-6698.41	-4846.17	46 113.3	15 247.9	200 989

Table 5. The same as in table 1, but for the CSM model.

CSM	¹⁷⁶ Os	¹⁷⁸ Os	¹⁸⁰ Os	¹⁸⁸ Os	¹⁹⁰ Os	¹⁵⁰ Nd	¹⁵⁶ Dy	¹⁶⁶ Hf	¹⁶⁸ Hf	¹⁷⁰ W
A_1 (keV)	17.03	17.26	16.51	10.25	9.063	19.219	15.45	14.87	16.04	16.19
A_2 (keV)	4.33	4.32	5.19	14.40	15.68	3.467	5.2	7.13	6.40	6.018
A_3 (keV)	-395.96	-240.13	-7.39	101.362	6.84	-658.299	-559.913	-5.04	-61.47	-186.946
A_4 (keV)	-275.24	-158.87	13.83	0.0	0.0	-491.884	-398.775	0.0	-36.48	-124.55
A_5 (keV)	-4.93	30.76	80.01	0.0	0.0	-438.394	-32.15	0.0	0.0	0.0
d	2.33	2.36	2.26	2.35	2.05	2.42	2.1	2.08	2.43	2.14
$q_1 [W.u.]^{1/2}$	0.411	0.246	0.86	0.409	0.229	0.527	1.112	0.158	0.211	-0.217
$q_2 [W.u.]^{1/2}$	-3.698	-3.862	6.99	0.785	1.213	-4.916	-9.474	-5.075	-3.936	-5.602
$q_3 [W.u.]^{1/2}$	0.0	0.0	0.0	-5.222	-9.395	6.344	19.576	0.0	0.0	0.0

Table 6. The energy spectra of the ground and first β and γ bands of the ¹⁷⁶Os nucleus yielded by the X(5), ISW, D, SSA and CSM models are compared with the corresponding experimental data taken from [53]. The energies are given in keV units. The approach which describe best the experimental data is mentioned in a box.

¹⁷⁶ Os	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$ \begin{array}{c} 2_{g}^{+} \\ 4_{g}^{+} \\ 6_{g}^{+} \\ 8_{g}^{+} \\ 10_{\sigma}^{+} \end{array} $	135	126	115	125	125	135
4+	396	367	340	386	377	394
6^{+}_{ρ}	743	686	647	746	723	742
8+	1158	1072	1026	1176	1143	1159
10^{+}_{o}	1634	1520	1473	1661	1624	1631
$12^{\circ +}_{g}$	2168	2028	1986	2192	2157	2152
14_{o}^{+}	2755	2593	2564	2764	2736	2718
16_{g}^{+}	3382	3216	3205	3374	3354	3326
18^{+}_{g}	4019	3894	3909	4017	4008	3973
20°+	4683	4628	4673	4693	4695	4660
22_{g}^{+}	5399	5417	5499	5399	5412	5385
24_{g}^{+}	6147	6261	6385	6134	6157	6147
0^+_{β}	601	714	565	633	498	601
2_{β}^{+}	742	942	760	757	723	742
4_{β}^{+}	1026	1351	1118	1019	1075	1032
6^+_{β}	1432	1865	1578	1378	1511	1432
8_{β}^{+}		2458	2121	1808	2011	1914
$\begin{array}{c} 0^+_{\beta} \\ 2^+_{\beta} \\ 4^+_{\beta} \\ 6^+_{\beta} \\ 8^+_{\beta} \\ 10^+_{\beta} \end{array}$		3121	2738	2293	2565	2411
$\frac{1}{2_{\nu}^{+}}$	864	949	951	926	943	989
3^+_{ν}	1038	1058	1056	1045	1053	1081
4^+_{ν}	1224	1189	1184	1196	1195	1201
5^{+}_{ν}	1410	1340	1333	1371	1345	1342
6^+_{ν}		1509	1503	1568	1542	1511
7^+_{ν}		1694	1691	1784	1719	1689
8^+_{ν}		1895	1898	2016	1962	1900
9^+_{ν}		2111	2124	2264	2161	2106
$\begin{array}{c} 2^{+}_{\gamma} \\ 3^{+}_{\gamma} \\ 4^{+}_{\gamma} \\ 5^{+}_{\gamma} \\ 6^{+}_{\gamma} \\ 7^{+}_{\gamma} \\ 8^{+}_{\gamma} \\ 9^{+}_{\gamma} \\ 10^{+}_{\gamma} \end{array}$		2343	2367	2525	2444	2354
rms (keV)		156	119	25	41	39

whose minima are shifted with respect to the $v_2(\gamma)$ minima. This can be viewed as the reduced potential of

$$\tilde{V}_2 = \frac{\hbar^2}{2B} \tilde{v}_2. \tag{2.34}$$

Performing a second-order expansion in $\sin 3\gamma$ of $v_2(\gamma)$ and of the terms originating from the rotational term, i.e. $\frac{9}{4\sin^3 3\gamma}$, and then making the change of variable $x=\cos 3\gamma$ in equation (2.8), we obtain the equation for the spheroidal functions [42]:

$$\[\left[(1 - x^2) \frac{\partial^2}{\partial x^2} - 2x \frac{\partial}{\partial x} + \lambda_{m_{\gamma}, n_{\gamma}} - c^2 x^2 - \frac{m_{\gamma}^2}{1 - x^2} \right] S_{m_{\gamma}, n_{\gamma}}(x) = 0, \qquad (2.35)$$

Table 7. The same as in table 6, but for ¹⁷⁸Os. The experimental data are taken from [57].

¹⁷⁸ Os	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
2+	132	127	116	131	130	132
4+	399	368	342	402	388	389
6^{+}_{σ}	762	688	650	769	739	736
2_{g}^{+} 4_{g}^{+} 6_{g}^{+} 8_{g}^{+} 10_{g}^{+}	1194	1075	1031	1203	1163	1152
10^{+}_{g}	1682	1525	1479	1689	1647	1625
12 ^g +	2220	2033	1994	2220	2181	2147
14_{g}^{+}	2805	2600	2572	2789	2758	2715
16_{g}^{+}	3429	3224	3214	3395	3374	3325
18 ⁺	4020	3905	3918	4033	4025	3975
20°+	4663	4641	4684	4701	4706	4664
22_{g}^{+}	5382	5432	5510	5399	5415	5391
24_{g}^{+}	6155	6278	6397	6125	6150	6155
0^+_{β}	651	716	574	635	493	651
2^{+}_{β}	771	944	771	766	730	771
4_{β}^{+}	1023	1355	1133	1037	1092	1029
6^+_{β}	1396	1870	1598	1403	1535	1396
8_{β}^{+}		2464	2144	1838	2041	1850
$\begin{array}{c} 0^+_{\beta} \\ 2^+_{\beta} \\ 4^+_{\beta} \\ 6^+_{\beta} \\ 8^+_{\beta} \\ 10^+_{\beta} \end{array}$		3129	2766	2324	2599	2374
2^+_{ν}	864	945	947	916	936	999
3^+_{ν}	1032	1055	1052	1041	1048	1091
4^+_{ν}	1213	1187	1181	1195	1195	1211
5^{+}_{ν}	1416	1338	1331	1375	1346	1350
6^+_{ν}		1507	1501	1575	1546	1519
7^+_{ν}		1692	1690	1793	1725	1696
8^+_{ν}		1894	1898	2027	1971	1907
9^+_{γ}		2111	2123	2275	2170	2113
2+ 3+ 4+ 5+ 6+ 7+ 8+ 9+ 10+		2343	2367	2537	2455	2361
rms (keV)		170	141	22	61	54

where

$$\lambda_{m_{\gamma},n_{\gamma}} = \frac{1}{9} \left[\tilde{\varepsilon}_{\gamma} - \frac{u_{1}}{2} - \frac{11}{27}D + \frac{1}{3}L(L+1) \right],$$

$$c^{2} = \frac{1}{9} \left(\frac{u_{1}}{2} + u_{2} - \frac{2}{27}D \right),$$

$$m_{\gamma} = \frac{K}{2}, \quad D = L(L+1) - K^{2} - 2.$$
From equation (2.36), we can determine the eigenvalue of the γ equation:

$$E_{\gamma}(n_{\gamma}, m_{\gamma}, L, K) = \frac{1}{\langle \beta^{2} \rangle} \frac{\hbar^{2}}{2B} \left(9\lambda_{m_{\gamma}, n_{\gamma}}(c) + \frac{u_{1}}{2} + \frac{11}{27}D - \frac{L(L+1)}{3} \right). \tag{2.37}$$

In equation (2.37), the term $u_1/2$ is washed out when the total energy is normalized to the ground state energy, which results in obtaining the γ eigenvalue depending on the sum of the γ potential parameters, due to the term c^2 . Hence, in some cases we can set one parameter to be equal to zero, for example u_2 , and consequently fit only u_1 . The γ functions are normalized to unity with the integration measure $|\sin 3\gamma| d\gamma$ as the Bohr–Mottelson model requires

$$\frac{3(2n_{\gamma}+1)(n_{\gamma}-m_{\gamma})!}{2(n_{\gamma}+m_{\gamma})!} \int_{0}^{\frac{\pi}{3}} |S_{m_{\gamma},n_{\gamma}}(\cos 3\gamma)|^{2} |\sin 3\gamma| \, \mathrm{d}\gamma = 1.$$
 (2.38)

¹⁸⁰Os X(5)**ISW** D SSA CSM Exp.

Table 8. The same as in table 6, but for ¹⁸⁰Os. The experimental data are taken from [58].

The total energy is obtained by summing the contributions coming from the β (2.13) and the γ (2.37) equations:

$$E(s, n_{\gamma}, m_{\gamma}, L, K) = B_1 x_{s,L}^2 + F \left[9\lambda_{m_{\gamma}, n_{\gamma}}(c) + \frac{u_1}{2} + \frac{11}{27}D - \frac{L(L+1)}{3} \right], \tag{2.39}$$

where the following notations were introduced:

rms (keV)

$$B_1 = \frac{1}{\beta_\omega^2} \frac{\hbar^2}{2B}, \quad F = \frac{1}{\langle \beta^2 \rangle} \frac{\hbar^2}{2B}.$$
 (2.40)

The total wavefunction is

$$\Psi(\beta, \gamma, \Omega) = C_{s,L} C_{n_{\gamma}, m_{\gamma}} C_{L,K} \beta^{-\frac{3}{2}} J_{\nu} \left(\frac{x_{s,L}}{\beta_{\omega}} \beta \right) S_{m_{\gamma}, n_{\gamma}} (\cos 3\gamma) \left[D_{M,K}^{L}(\Omega) + (-1)^{L} D_{M,-K}^{L}(\Omega) \right],$$
(2.41)

where with $C_{n_{\gamma},m_{\gamma}}$ was denoted the normalization factor of the γ function, while $C_{L,K}$ is the normalization factor of the Wigner function:

$$C_{L,K} = \sqrt{\frac{2L+1}{16\pi^2(1+\delta_{K,0})}}. (2.42)$$

2.3.3. The D model. The D model was proposed by the present authors and collaborators in [42] and differs from the ISW model by that the ISW potential for the β variable is replaced

¹⁸⁸ Os	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
2 ⁺	155	179	179	151	152	150
4+	478	519	519	479	476	468
6_a^{+}	940	970	970	945	935	934
84	1515	1516	1516	1512	1501	1535
10^{+}_{a}	2170	2149	2150	2156	2154	2264
$\begin{array}{c} 2_g^+ \\ 4_g^+ \\ 6_g^+ \\ 8_g^+ \\ 10_g^+ \\ 12_g^+ \end{array}$	2856	2867	2868	2860	2877	3116
0^+_{β}	1086	1009	1007	1120	1063	1164
2_{β}^{+}	1305	1331	1328	1270	1330	1305
4_{β}^{+}		1910	1907	1599	1808	1621
6^+_{β}		2636	2632	2064	2421	2096
8^{+}_{β}		3474	3470	2632	3132	2717
$\begin{array}{c} 0^{+}_{\beta} \\ 2^{+}_{\beta} \\ 4^{+}_{\beta} \\ 6^{+}_{\beta} \\ 8^{+}_{\beta} \\ 10^{+}_{\beta} \end{array}$		4412	4407	3276	3920	3475
$\frac{1}{2^{+}_{\nu}}$	633	631	631	627	641	665
3 ⁺ +	790	786	785	773	791	790
4+	966	972	971	959	969	956
5_{ν}^{+}	1181	1185	1185	1180	1172	1157
6+	1425	1423	1423	1432	1434	1399
7_{ν}^{+}	1686	1685	1684	1709	1674	1669
8 ^{'+}		1969	1969	2009	2008	1983
9+		2275	2275	2329	2273	2318
$\begin{array}{c} 2^{+} \\ 3^{+} \\ 3^{+} \\ 4^{+} \\ 5^{+} \\ 6^{+} \\ 7^{+} \\ 8^{+} \\ 9^{+} \\ 10^{+} \\ \end{array}$		2602	2603	2666	2670	2701
rms (keV)		27	27	16	13	36

Table 9. The same as in table 6, but for ¹⁸⁸Os. The experimental data are taken from [59].

with the Davidson potential (2.14). Hence, the total energy of the system is obtained by adding the energy of the β equation with the Davidson potential given by equation (2.16) and the energy of the γ equation (2.37):

$$E(n_{\beta}, n_{\gamma}, m_{\gamma}, L, K) = E\left(2n_{\beta} + 1 + \sqrt{\frac{L(L+1)}{3} + \frac{9}{4} + \beta_0^4}\right) + F\left[9\lambda_{m_{\beta}, n_{\gamma}}(c) + \frac{u_1}{2} + \frac{11}{27}D - \frac{L(L+1)}{3}\right],$$
(2.43)

where $E = \hbar^2/2B$. The total wavefunction has the expression

$$\Psi(\beta, \gamma, \Omega) = C_{n_{\beta}, L} C_{n_{\gamma}, m_{\gamma}} C_{L,K} f_{n_{\beta}, L}(\beta) S_{m_{\gamma}, n_{\gamma}}(\cos 3\gamma) \left[D_{M,K}^{L}(\Omega) + (-1)^{L} D_{M,-K}^{L}(\Omega) \right], \quad (2.44)$$
where $C_{n_{\beta}, L}$ is the normalization factor of $f_{n_{\beta}, L}(\beta)$ given by equation (2.15).

2.3.4. The present approach. In the present approach, called conventionally the SSA, a sextic potential (2.19) for the β variable is considered, while for the γ variable a periodic potential (2.32) with a minimum at $\gamma_0 = 0^0$. The β equation is quasi-exactly solved, having the solutions given by equations (2.22) and (2.23), while the γ equation is reduced to the spheroidal equation (2.35) with

$$\lambda_{m_{\gamma},n_{\gamma}} = \frac{1}{9} \left[\tilde{\varepsilon}_{\gamma} - \frac{u_{1}}{2} - \frac{11}{27}D + \frac{1}{3}L(L+1) \right] + \frac{2L(L+1)}{27},$$

$$c^{2} = \frac{1}{9} \left(\frac{u_{1}}{2} + u_{2} - \frac{2}{27}D \right),$$

$$m_{\gamma} = \frac{K}{2}, \quad D = L(L+1) - K^{2} - 2.$$
(2.45)

¹⁹⁰Os X(5)ISW D SSA **CSM** Exp. $\begin{array}{c} 2^{+}_{,4}4^{+}_{g}+\frac{1}{9}\\ 4^{+}_{g}+\frac{1}{9}\\ 0^{+}_{,4}+\frac{1}{9}\\ 2^{+}_{,4}+\frac{1}{9}\\ 4^{+}_{,4}+\frac{1}{9}\\ 4^{+}_{,4}+$ rms (keV)

Table 10. The same as in table 6, but for ¹⁹⁰Os. The experimental data are taken from [60].

In equation (2.45), the term 2L(L+1)/3 multiplied with 1/9 comes from the rotational term (2.17). The expression for the total energy of the system is obtained by using equations (2.23) and (2.45):

$$E(n_{\beta}, n_{\gamma}, m_{\gamma}, L, K) = E[b(2L+3) + \lambda_{n_{\beta}}^{(M)} + u_{0}^{\pm}] + F\left[9\lambda_{m_{\beta}, n_{\gamma}}(c) + \frac{u_{1}}{2} + \frac{11}{27}D - L(L+1)\right].$$
(2.46)

The corresponding wavefunction is

$$\Psi(\beta, \gamma, \Omega) = N_{n_{\beta}, L} C_{n_{\gamma}, m_{\gamma}} C_{L, K} \beta^{-2} \varphi_{n_{\beta}, L}(\beta) S_{m_{\gamma}, n_{\gamma}}(\cos 3\gamma) \left[D_{M, K}^{L}(\Omega) + (-1)^{L} D_{M, -K}^{L}(\Omega) \right], \tag{2.47}$$

where $\varphi_{n_{\beta},L}(\beta)$ is given by equation (2.22).

2.4. E2 transition probabilities

The reduced E2 transition probabilities are determined by

$$B(E2; L_i \to L_f) = |\langle L_i || T_2^{(E2)} || L_f \rangle|^2,$$
 (2.48)

where the Rose convention [47] was used. For the ISW, D and SSA models, in equation (2.48), an anharmonic transition operator is used:

$$T_{2\mu}^{(E2)} = t_1 \beta \left[\cos \gamma D_{\mu 0}^2(\Omega) + \frac{\sin \gamma}{\sqrt{2}} (D_{\mu 2}^2(\Omega) + D_{\mu, -2}^2(\Omega)) \right] + t_2 \sqrt{\frac{2}{7}} \beta^2 \left[-\cos 2\gamma D_{\mu 0}^2(\Omega) + \frac{\sin 2\gamma}{\sqrt{2}} (D_{\mu 2}^2(\Omega) + D_{\mu, -2}^2(\Omega)) \right].$$
(2.49)

¹⁵⁰Nd X(5)**ISW** D SSA CSM Exp. $\begin{array}{c} 2^{+}_{,4} + 6^{+}_{,6} + 6^{+}_{,6} \\ 6^{+}_{,6} + 6^{+}_{,6} + 6^{+}_{,6} \\ 10^{+}_{,6} + 6^{+}_{,6} + 6^{+}_{,6} \\ 0^{+}_{,6} + 6^{+}_{,6} + 6^{+}_{,6} \\ 2^{+}_{,7} + 2^{+}_{,7} + 6^{+}_{,7} + 7^{+}_{,7} \\ 10^{+}_{,7} + 2^{+}_{,7} + 2^{+}_{,7} \end{array}$ rms (keV)

Table 11. The same as in table 6, but for ¹⁵⁰Nd. The experimental data are taken from [61].

The parameters t_1 and t_2 will be determined by the least-squares method. For the X(5) model, in the limit of γ -small, only the harmonic part of the transition operator (2.49) is used:

$$T_{2\mu,X(5)}^{(E2)} = t\beta D_{\mu 0}^2(\Omega) + t\beta \frac{\gamma}{\sqrt{2}} (D_{\mu 2}^2(\Omega) + D_{\mu,2}^2(\Omega)). \tag{2.50}$$

The first term of equation (2.50) gives contributions only to $\Delta K = 0$ transitions, while the second term to $\Delta K = 2$ transitions. For $\Delta K = 0$ transitions, the matrix element of the γ variable is reduced to the orthogonality condition, while for $\Delta K = 2$ the γ matrix element can be considered as an intrinsic transition matrix element. Finally, the reduced transition probabilities will depend on two parameters [48]. Here, we will denote these two parameters with t and t' for $\Delta K = 0$ and $\Delta K = 2$ transitions, respectively.

3. The coherent state model

The CSM defines [5] first a restricted collective space whose vectors are model states of the ground, β and γ bands. In choosing these states, we were guided by some experimental information which results in formulating a set of criteria to be fulfilled by the searched states.

All these restrictions required are fulfilled by the following set of three deformed quadrupole boson states:

$$\psi_g = e^{[d(b_0^{\dagger} - b_0)]} |0\rangle \equiv T|0\rangle, \quad \psi_{\gamma} = \Omega_{\gamma, 2}^{\dagger} \psi_g, \quad \psi_{\beta} = \Omega_{\beta}^{\dagger} \psi_g, \tag{3.1}$$

Table 12. The same as in table 6, but for ¹⁵⁶Dy. The experimental data are taken from [63].

¹⁵⁶ Dy	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
2_{σ}^{+}	138	119	114	140	131	168
4+	404	345	344	422	391	457
6^{+}_{σ}	770	646	668	796	745	829
2_{g}^{+} 4_{g}^{+} 6_{g}^{+} 8_{g}^{+} 10_{g}^{+}	1216	1009	1079	1230	1175	1267
10^{+}_{g}	1725	1431	1572	1712	1667	1761
12°+	2286	1908	2145	2232	2151	2307
14 [‡] _g	2888	2440	2796	2787	2807	2899
$0^+_{eta} \ 2^+_{eta}$	676	672	451	648	461	676
2^+_{β}	829	886	629	788	703	829
4^+_{β}	1088	1272	966	1070	1068	1102
6^+_{β}	1437	1755	1413	1444	1515	1452
8^+_{β}	1859	2313	1955	1878	2026	1859
4^{+}_{β} 6^{+}_{β} 8^{+}_{β} 10^{+}_{β}	2316	2937	2584	2360	2593	2312
2+ 3+ 4+ 5+ 6+ 7+ 8+ 9+ 10+ 11+	891	1069	898	839	928	921
3_{ν}^{+}	1022	1172	1004	970	1041	1024
4_{ν}^{+}	1168	1296	1136	1129	1188	1159
5^{+}_{ν}	1336	1438	1292	1312	1339	1312
6^+_{ν}	1525	1596	1472	1514	1542	1497
7^+_{ν}	1729	1771	1674	1732	1720	1686
8^+_{ν}	1959	1960	1899	1964	1972	1913
9^+_{ν}	2192	2163	2145	2210	2171	2131
10^{+}_{γ}	2448	2381	2413	2467	2464	2395
11^{+}_{γ}	2712	2613	2702	2735	2680	2636
12 ⁺ _γ	2997	2859	3013	3013	2949	2934
13^{+}_{γ}	3274	3118	3345	3301	3240	3153
14^{+}_{ν}		3391	3698	3600	3606	3526
15^{+}_{γ}	3861	3677	4071	3908	3847	3805
rms (keV)		232	114	35	90	41

where the excitation operators for β and γ bands are defined by

$$\Omega_{\gamma,2}^{\dagger} = (b^{\dagger}b^{\dagger})_{2,2} + d\sqrt{\frac{2}{7}}b_{2,2}^{\dagger}, \quad \Omega_{\beta}^{\dagger} = (b^{\dagger}b^{\dagger}b^{\dagger})_{0} + \frac{3d}{\sqrt{14}}(b^{\dagger}b^{\dagger})_{0} - \frac{d^{3}}{\sqrt{70}}.$$
 (3.2)

Here, d is a real parameter simulating the nuclear deformation. From the three deformed states, one generates through projection three sets of mutually orthogonal states

$$\varphi_{JM}^i = N_J^i P_{M0}^J \psi_i, \quad i = g, \beta, \gamma, \tag{3.3}$$

where P_{MK}^{J} denotes the projection operator:

$$P_{MK}^{J} = \frac{2J+1}{8\pi^2} \int D_{MK}^{J^*} \hat{R}(\Omega) \, d\Omega, \tag{3.4}$$

 N_J^i the normalization factors and D_{MK}^I the rotation matrix elements. The rotation operator corresponding to the Euler angles Ω is denoted by $\hat{R}(\Omega)$. It was proved that the deformed and projected states contain the salient features of the major collective bands. Since we attempt to set up a very simple model, we rely on the experimental feature saying that the β band is largely decoupled from the ground as well as from the γ bands and choose a model Hamiltonian whose

¹⁶⁶Hf X(5)ISW D SSA **CSM** Exp. $\begin{array}{c} 2_{g}^{+} \\ 4_{g}^{+} \\ 6_{g}^{+} \\ 6_{g}^{+} \\ 10_{g}^{+} \\ 6_{g}^{+} \\ 10_{g}^{+} \\ 6_{g}^{+} \\ 7_{g}^{+} \\ 6_{g}^{+} \\ 7_{g}^{+} \\ 7_{g}^{+} \\ 7_{g}^{+} \\ 7_{g}^{+} \\ 6_{g}^{+} \\ 7_{g}^{+} \\ 7_{g}^{+}$ rms (keV)

Table 13. The same as in table 6, but for ¹⁶⁶Hf. The experimental data are taken from [64].

matrix elements between beta states and states belonging either to the ground or to the gamma band are all equal to zero. The simplest Hamiltonian obeying this restriction is

$$H = A_1 (22\hat{N} + 5\Omega_{\beta'}^{\dagger} \Omega_{\beta'}) + A_2 \hat{J}^2 + A_3 \Omega_{\beta}^{\dagger} \Omega_{\beta}, \tag{3.5}$$

where \hat{N} is the boson number, \hat{J}^2 -angular momentum squared and $\Omega_{\beta'}^{\dagger}$ denotes

$$\Omega_{\beta'}^{\dagger} = (b^{\dagger}b^{\dagger})_{00} - \frac{d^2}{\sqrt{5}}.$$
 (3.6)

Higher order terms in boson operators can be added to the Hamiltonian H without altering the decoupling condition for the beta band. An example of this kind is the correction

$$\Delta H = A_4(\Omega_{\beta}^{\dagger} \Omega_{\beta'}^2 + \text{h.c.}) + A_5 \Omega_{\beta'}^{\dagger 2} \Omega_{\beta'}^2. \tag{3.7}$$

The energies for beta band and for the gamma band states of odd angular momentum are described as average values of H (3.5), or $H + \Delta H$ on φ_{JM}^{β} and φ_{JM}^{γ} (J-odd), respectively. As for the energies for the ground band and those of gamma band states with even angular momentum, they are obtained by diagonalizing a 2×2 matrix for each J.

The quadrupole transition operator is considered to be a sum of a linear term in bosons and one which is quadratic in the quadrupole bosons:

$$Q_{2\mu} = q_1(b_{2\mu}^{\dagger} + (-)^{\mu}b_{2,-\mu}) + q_2((b^{\dagger}b^{\dagger})_{2\mu} + (bb)_{2\mu}) + q_3(b^{\dagger}b)_{2\mu}.$$
(3.8)

Note that if $q_3 = 2q_2$, the quadrupole transition operator can be obtained from the quadrupole transition operator expressed in terms of the collective quadruple coordinates $\alpha_{2\mu}$:

$$Q_{2\mu} = Q_1 \alpha_{2\mu} + Q_1'(\alpha_2 \alpha_2)_{2\mu}. \tag{3.9}$$

Table 14. The same as in table 6, but for ¹⁶⁸Hf. The experimental data are taken from [65].

¹⁶⁸ Hf	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
2_{σ}^{+}	124	141	140	120	108	128
4 +	386	409	409	382	351	389
6^{+}_{o}	757	765	769	757	710	756
8+	1214	1195	1207	1215	1172	1206
10^{+}_{g}	1736	1694	1720	1736	1723	1730
$\begin{array}{c} 2_g^+ \\ 4_g^+ \\ 6_g^+ \\ 8_g^+ \\ 10_g^+ \\ 12_g^+ \end{array}$	2306	2259	2303	2307	2354	2320
0^+_{β}	942	795	755	928	878	942
2_{β}^{+}	1059	1049	1002	1048	1039	1049
4_{β}^{+}	1285	1505	1449	1310	1368	1285
6^+_{β}		2077	2015	1684	1823	1630
8^{+}_{β}		2738	2672	2143	2380	2068
$\begin{array}{c} 0^+_{\beta} \\ 2^+_{\beta} \\ 4^+_{\beta} \\ 6^+_{\beta} \\ 8^+_{\beta} \\ 10^+_{\beta} \end{array}$		3477	3412	2664	3024	2587
2_{ν}^{+}	876	911	906	902	928	939
3 ⁺ _{\(\nu\)}	1031	1033	1028	1020	1042	1035
$4_{\nu}^{'+}$	1161	1179	1178	1172	1171	1161
5_{ν}^{+}	1386	1347	1350	1353	1334	1311
6^+_{ν}	1551	1535	1543	1558	1530	1492
7_{ν}^{+}		1741	1755	1786	1733	1687
$8_{\nu}^{'+}$		1965	1988	2033	1992	1916
9_{ν}^{+}		2206	2239	2297	2226	2148
$\begin{array}{c} 2^{+}_{\gamma} \\ 3^{+}_{\gamma} \\ 4^{+}_{\gamma} \\ 5^{+}_{\gamma} \\ 6^{+}_{\gamma} \\ 7^{+}_{\gamma} \\ 8^{+}_{\gamma} \\ 9^{+}_{\gamma} \\ 10^{+}_{\gamma} \end{array}$		2464	2508	2576	2543	2421
rms (keV)		75	70	15	43	31

The anharmonic term in the above expression can be obtained by expanding the deformed mean field around the spherical equilibrium shape [50, 51] of the nuclear surface. For the near vibrational regime, the interband matrix elements of the q_3 term vanish within the CSM [5]. Moreover, a transition operator depending on two free parameters seems to be suitable for describing the E2 transition probabilities in several regions of the nuclide's chart [52].

Using the Rose convention [47], the reduced probability for the E2 transition $J_i^+ \to J_f^+$ can be expressed as

$$B(E2; J_i^+ \to J_f^+) = (\langle J_i^+ || Q_2 || J_f^+ \rangle)^2.$$
 (3.10)

Three specific features of the CSM are worth mentioning.

- (a) The model states are generated through projection from a coherent state and two excitations of that through simple polynomial boson operators. Thus, it is expected that the projected states may account for the semiclassical behavior of the nuclear system staying in a state of high spin.
- (b) The states are an infinite series of bosons and thus highly deformed states can be described.
- (c) The model Hamiltonian is not commuting with the boson number operator and because of this property a basis generated from a coherent state is expected to be most suitable.

The CSM has been successfully applied to several nuclei exhibiting various equilibrium shapes which, according to the IBA classification, exhibit the SO(6), SU(5) and SU(3) symmetries. Several improvements of the CSM have been proposed by considering additional degrees of freedom such as isospin [6], quasi-particle [7] or collective octupole coordinates

Table 15. The same as in table 6, but for ¹⁷⁰W. The experimental data are taken from [66].

$\begin{array}{c ccccccccccccccccccccccccccccccccccc$							
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$^{170}\mathbf{W}$	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	2_o^+	157	145	133	145	151	171
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	4+	462	420	398	447	446	475
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	6_{σ}^{+}	876	785	767	858	844	873
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	8^+_{σ}	1363	1226	1228	1347	1323	1346
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	10 ⁺	1902	1739	1777	1894	1869	1882
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	12^{+}_{g}	2464	2319	2411	2489	2471	2477
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	14_{ρ}^{+}	3118	2965	3128	3126	3124	3128
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	16_g^{+}	3816	3677	3927	3801	3821	3831
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	0_{β}^{+}		816	587	760	507	823
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	2_{β}^{+}	953	1077	804	905	790	953
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$4_{\scriptscriptstyle B}^{\scriptscriptstyle +}$	1202	1545	1208	1207	1204	1215
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	6_{β}^{+}	1578	2132	1736	1618	1706	1578
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	8_{β}^{+}		2810	2367	2107	2277	2020
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	10^{+}_{β}		3568	3093	2654	2905	2531
		937	944		928	936	965
	3^+_{ν}	1074	1068	1068	1066	1064	1074
	4^+_{ν}	1220	1219	1219	1238	1231	1217
	5^+_{ν}		1391	1397	1438	1400	1381
	6^+_{ν}		1584	1600	1662	1630	1578
	7^{+}_{ν}		1796	1828	1906	1828	1783
	8^+_{ν}		2025	2080	2168	2109	2027
	9^+_{ν}		2273	2355	2446	2329	2264
rms (keV) 200 90 21 58 13	10^{+}_{γ}		2538	2653	2739	2655	2550
	rms (keV)		200	90	21	58	13

[8, 9]. The CSM has also been used to describe some nonaxial nuclei [49], and the results were compared with those obtained with the rotationvibration model [2]. A review of the CSM achievements is found in [10]. The terms involved in the model Hamiltonians used in by the CSM [5] and its generalized version [6] have microscopic counterparts as shown in [11] and [12], respectively.

4. Numerical results

4.1. Parameters

The parameters which define the energies and the E2 transition probabilities of the models X(5), ISW, D, SSA and CSM were fitted by the least-squares method for ten nuclei: 176,178,180,188,190 Os, 150 Nd, 156 Dy, 166,168 Hf and 170 W. In the least-squares procedure, all experimental energies were considered. The resulting values are those given in tables 1-5. For the first three and the last three nuclei from table 1, the parameter t' cannot be determined since the corresponding term from the transition operator does not contribute to the intraband decays.

Some parameters vary by a large amount from one isotope to another but the relative variation is small. For example in the case of Os isotopes, the parameters could be interpolated

Table 16. The reduced E2 transition probabilities determined with the X(5), ISW, D, SSA and CSM models for the ¹⁷⁶Os nucleus are compared with the corresponding experimental data taken from [54].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_g^+ \rightarrow 0_g^+$	144^{+5}_{-5}	167	127	145	136	144
$4_g^+ \rightarrow 2_g^+$	243^{+5}_{-5}	264	224	228	227	253
$6_g^+ \rightarrow 4_g^+$	305^{+11}_{-11}	330	305	292	297	328
$8_g^+ \rightarrow 6_g^+$	321^{+15}_{-14}	379	377	360	366	393
$10_g^+ \rightarrow 8_g^+$	441^{+88}_{-63}	419	438	433	435	452
$12_g^+ \to 10_g^+$	517^{+336}_{-146}	450	490	510	504	517

Table 17. The same as in table 16, but for ¹⁷⁸Os. The experimental data are taken from [54–56]

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	138	147	137	146	141	138
$4_g^+ \rightarrow 2_g^+$	226	232	225	226	226	227
$6^{+}_{\sigma} \rightarrow 4^{+}_{\sigma}$	290	291	287	280	283	282
$8^{+}_{\scriptscriptstyle \rho} \rightarrow 6^{+}_{\scriptscriptstyle \rho}$	327	334	337	332	334	327
$1\mathring{0}_g^+ \rightarrow \mathring{8}_g^+$	384	369	378	384	382	368

Table 18. The same as in table 16, but for ¹⁸⁰Os. The experimental data are taken from [58].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_{\varrho}^{+} \rightarrow 0_{\varrho}^{+}$	120^{+30}_{-30}	70	152	148	151	150
$\begin{array}{c} 2_g^+ \rightarrow 0_g^+ \\ 4_g^+ \rightarrow 2_g^+ \\ 6_g^+ \rightarrow 4_g^+ \end{array}$	193^{+25}_{-25}	111	167	177	172	149
$6_g^+ o 4_g^+$	160^{+40}_{-40}	139	132	139	135	120
$8_g^+ \rightarrow 6_g^+$	63^{+13}_{-13}	160	95	83	90	96

by smooth curves. One parameter falls aside namely those of 188 Os, which seems to achieve the critical point of the shape transition, i.e. exhibits an X(5) behavior.

We note that the parameter F involves the average value $\langle \beta^2 \rangle$ which, in principle, is an angular-momentum-dependent quantity. Therefore, the differential equation in γ should be iteratively solved, at each step the inserted average value being calculated with the wavefunction provided in the previous step. When the convergence of the process is met, one keeps the average value for the chosen angular momentum. Here, $\langle \beta^2 \rangle$ was kept constant. Whether this hypothesis is valid or not can be posterity checked. To this goal, we represented in figure 1, the average $\langle \beta^2 \rangle$ for each of the models ISW, D and SSA. We note that the average value only slightly depends on J and that is especially true for ISW and SSA. If the limit of $\langle \beta^2 \rangle$, when the convergence of the iterations mentioned above is reached, depends on J like the averages shown in figure 1, one could say that keeping $\langle \beta^2 \rangle$ constant one ignores a slight decrease of energy with angular momentum.

With the parameters listed above the potentials in the variables β and γ and the wavefunctions describing the low lying states from the ground, beta and gamma bands respectively, are represented for four nuclei, as shown in figures 2–5. Analyzing these figures, several features can be noted. The β potential has a deformed minimum located at a deformation which differs from one nucleus to another. The wavefunctions in β for 0^+_g and 2^+_{γ} are almost

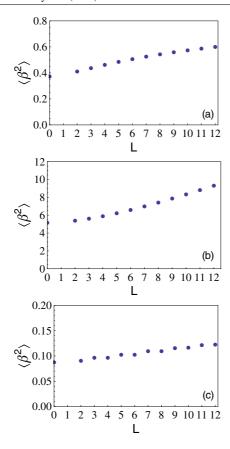


Figure 1. The average values of β^2 versus the angular momentum calculated within the ISW (panel a)), the D (panel b)) and the SSA (panel c)) models.

identical and have only one maximum and no node, while the band for 0^+_β has one node, one maximum and one minimum. The maximum of the $|\phi|^2$ distribution for the three states represented in the quoted figures is achieved at a point which is close to the potential minimum. If $|\phi|^2$ is multiplied with the integration measure over β , then the probability distribution has a maximum closer to the potential minimum. The state 0^+_β is characterized by two maxima for the probability distribution of the beta variable. This feature reflects the specific structure of the excitation operator of this state, from the ground state, i.e. $n_\beta=1$. The behavior of the wavefunctions in the variable γ is mainly determined by the discontinuity for $\gamma=0$ and $\gamma=\frac{\pi}{3}$. The potential has two minima, one well pronounced near the first wall and one very flat close to the $\gamma=\frac{\pi}{3}$ discontinuity. Due to this structure, the wavefunction describing a state in the ground band has two maxima located above the mentioned minima. The state 2^+_γ heading the gamma band has an additional maximum.

4.2. Energies

The spectra of the chosen nuclei, determined by the models X(5), ISW, D, SSA and CSM, are compared with the corresponding experimental data in tables 6–15. The quality of agreement between the results of our calculations and the corresponding experimental data is given by

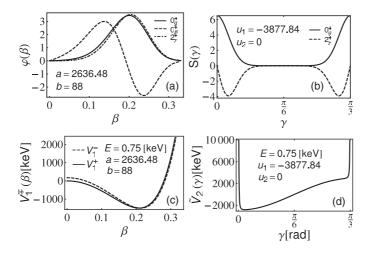


Figure 2. The solutions for the equation in β , corresponding to various angular momenta and the potential from the left-bottom panel, are plotted, in panel left-up, as a function of β . Similarly, on the right column the wavefunctions for γ for different angular momenta and the effective potential shown in the right-bottom panel are plotted as a function of γ . The results correspond to 150 Nd.

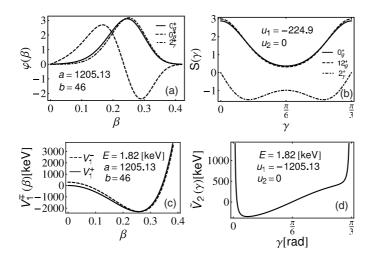


Figure 3. The same as in figure 2, but for ¹⁶⁶Hf.

the rms values of the deviations. Thus, comparing the rms values corresponding to different models we conclude that for 180 Os, 150 Nd and 170 W the best description of the spectra is that given by the CSM approach, energies of 188 Os calculated with the SSA are closest to the experimental ones, while for the remaining nuclei the D formalism provides the most realist picture.

Using the experimental data listed in tables 6–15, one can calculate the ratio of the excitation energies for the states 4_g^+ and 2_g^+ , denoted by $R_{4_g^+/2_g^+}$. The results are 2.93 (176 Os, 190 Os, 150 Nd, 156 Dy), 2.94 (170 W), 2.96 (166 Hf), 3.02 (178 Os), 3.08 (188 Os), 3.10 (180 Os) and 3.11 (168 Hf). We note that all nuclei are characterized by a ratio $R_{4_g^+/2_g^+}$ which is close to the

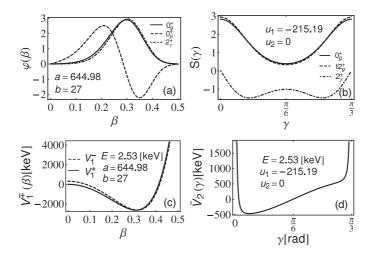


Figure 4. The same as in figure 2, but for ¹⁸⁸Os.

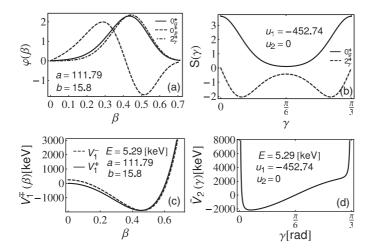


Figure 5. The same as in figure 2, but for ¹⁹⁰Os.

value of 2.9 assigned to the critical point of the transition $SU(5) \rightarrow SU(3)$, which is described by the solvable model called X(5). Despite this, the X(5) approach provides a description which is worse than those obtained with the other models proposed here.

4.3. Reduced transition probabilities

As mentioned before, the parameters involved in the transition operators employed by different models have been fixed by fitting through a least-squares procedure the existent data. With the fitted parameter, the results for the reduced E2 transition probabilities are presented in tables 16-25 where one also gives for comparison the available experimental data. For the lightest three isotopes of Os as well as for 166,168 Hf and 170 W, the available experimental data

Table 19. The same as in table 16, but for ¹⁸⁸Os. The experimental data are taken from [59].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_g^+ \to 0_g^+$	79 ⁺² ₋₂	74	72	79	82	42
$4_g^+ \rightarrow 2_g^+$	133_{-8}^{+8}	118	115	121	123	87
$6_g^+ \rightarrow 4_g^+$	138^{+8}_{-8}	147	144	147	145	125
$8_g^+ \rightarrow 6_g^+$	161^{+11}_{-11}	169	166	174	162	161
$10_g^+ \rightarrow 8_g^+$	188^{+25}_{-25}	187	184	203	178	195
$0^+_eta ightarrow 2^+_g$	$0.95^{+0.08}_{-0.08}$	47	48	33	21	0.95
$0^+_{\beta} ightarrow 2^+_{\gamma}$	$4.3^{+0.5}_{-0.5}$	5.2	5.2	1.9	1.5	44
$4^+_{\nu} ightarrow 2^+_{ u}$	47^{+10}_{-10}	47	50	52	56	14
$4^+_{\gamma} \rightarrow 3^+_{\gamma}$	320^{+120}_{-120}	112	117	120	132	43
$6_{\gamma}^{+} \rightarrow 4_{\gamma}^{+}$	70^{+30}_{-30}	107	111	114	118	31
$2^+_{\gamma} ightarrow 0^+_{g}$	$5^{+0.6}_{-0.6}$	8.4	10.9	10.8	9.9	5
$2_{\gamma}^{+} \rightarrow 2_{g}^{+}$	16^{+2}_{-2}	13	17	16	14	10.4
$2_{\gamma}^{+} \rightarrow 4_{g}^{+}$	34^{+5}_{-5}	0.65	0.85	0.80	0.73	1.4
$4_{\gamma}^{+} \rightarrow 2_{g}^{+}$	$1.29^{+0.19}_{-0.19}$	5.7	7.1	6.7	6.1	1.7
$4_{\gamma}^{+} \rightarrow 4_{g}^{+}$	19^{+3}_{-3}	18	23	20	19	10.7
$4_{\gamma}^{\prime} \rightarrow 6_{g}^{+}$	16^{+7}_{-7}	2	2	2	2	5
$6^{+}_{\nu} \rightarrow 4^{+}_{a}$	$0.21^{+0.11}_{-0.11}$	5.3	6.4	5.8	5.3	0.9
$6_{\gamma}^{\prime} \rightarrow 6_{g}^{\circ}$	>9.4	21	25	23	20	8.3

Table 20. The same as in table 16, but for ¹⁹⁰Os. The experimental data are taken from [60].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_g^+ \rightarrow 0_g^+$	72^{+2}_{-2}	58	57	56	61	45
$4_a^+ \rightarrow 2_a^+$	105_{-6}^{+6}	91	91	88	94	83
$6^+_{\sigma} \rightarrow 4^+_{\sigma}$	113^{+10}_{-10}	115	113	112	112	112
$6_g^{\sharp} \to 4_g^{\sharp} 8_g^{+} \to 6_g^{+}$	137^{+20}_{-20}	131	130	138	126	137
$10_g^+ \rightarrow 8_g^+$	120^{+30}_{-30}	145	143	165	139	160
$0^+_{eta} ightarrow 2^+_{g}$	$2.2^{+0.5}_{-0.5}$	36	36	30	19	2.2
$0^{\stackrel{r}{+}}_{eta} ightarrow 2^{\stackrel{\circ}{+}}_{\gamma}$	23^{+7}_{-7}	8.9	9	8	5	148
$\overline{4^+_{\nu} ightarrow 2^+_{ u}}$	53+5	36	38	37	41	20.4
$4_{\gamma}^{+} \rightarrow 3_{\gamma}^{+}$	65^{+13}_{-13}	87	90	87	98	84
$6^{+}_{\nu} \to 4^{+}_{\nu}$	65^{+13}_{-13}	83	85	84	89	49
$8_{\gamma}^{+} \rightarrow 6_{\gamma}^{+}$	61_{-16}^{-13}	112	113	119	115	72
$2^+_{\nu} \to 0^+_{\sigma}$	$5.9^{+0.6}_{-0.6}$	14.2	15.6	15.9	16.2	14
$2^+_{\nu} \rightarrow 2^+_{\varrho}$	33^{+4}_{-4}	21	24	24	24	33
$4_{\gamma}^{+} ightarrow 2_{g}^{+}$	$0.68^{+0.06}_{-0.06}$	9.7	10.3	10.4	10.3	4.3
$4_{\gamma}^{+} \rightarrow 4_{g}^{+}$	30^{+4}_{-4}	31	33	33	32	31
$6^+_{\nu} \rightarrow 4^+_{\varrho}$	< 0.8	9	10	10	9	1.7
$6_{\gamma}^{'+} ightarrow 6_{g}^{'+}$	31^{+8}_{-8}	36	38	40	36	26

refer to the states of ground band. Agreements with the experimental data showed up by the five theoretical models are comparable in quality.

For ¹⁵⁶Dy, besides the intraband transitions in the ground band, a few interband transitions from the gamma to the ground band are experimentally known. As seen from table 22, agreement between calculations with the experimental data is quite good.

Table 21. The same as in table 16, but for ¹⁵⁰Nd. The experimental data are taken from [62].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_g^+ \to 0_g^+$	115 ⁺² ₋₂	107	104	92	116	81
$4_g^+ \rightarrow 2_g^+$	182^{+2}_{-2}	169	168	144	177	160
$6_g^+ \rightarrow 4_g^+$	210^{+2}_{-2}	212	210	183	211	222
$\begin{array}{l} 2_g^+ \to 0_g^+ \\ 4_g^+ \to 2_g^+ \\ 6_g^+ \to 4_g^+ \\ 8_g^+ \to 6_g^+ \end{array}$	278^{+25}_{-25}	243	243	224	240	278
$10_g^+ \rightarrow 8_g^+$	204_{-12}^{+12}	269	269	268	266	330
$2^+_{\beta} o 0^+_{\beta}$	114^{+23}_{-23}	85	83	130	86	116
$4^+_{\beta} \rightarrow 2^+_{\beta}$	170^{+51}_{-51}	128	125	194	144	165
$0^+_{\beta} \rightarrow 2^+_g$	39+2	67	73	51	37	41.2
$2^{+}_{\beta} \rightarrow 0^{+}_{g}$	$1.2^{+0.2}_{-0.2}$	2.1	2.9	3.1	1.6	5.2
$2_{eta}^{r_+} ightarrow 2_{g}^{r_+}$	9_{-2}^{+2}	10	10	9	6	9
$2_{eta}^{r_+} ightarrow4_g^+$	17^{+3}_{-3}	39	42	40	26	26
$4^+_{\beta} \rightarrow 2^+_{\beta}$	$0.12^{+0.02}_{-0.02}$	1.07	1.61	1.64	0.57	5.6
$\begin{array}{c} 4_{\beta}^{+} \rightarrow 2_{g}^{+} \\ 4_{\beta}^{+} \rightarrow 4_{g}^{+} \end{array}$	7^{+1}_{-1}	6	8	8	5	7.2
$4_{\beta}^{+} \rightarrow 6_{g}^{+}$	70^{+13}_{-13}	30	33	46	26	26
$2_{\gamma}^+ \rightarrow 0_g^+$	$3^{+0.8}_{-0.8}$	2.4	8	9.8	5.1	16.3
$2_{\gamma}^{'+} \rightarrow 2_{g}^{\ddot{+}}$	$5.4^{+1.7}_{-1.7}$	3.6	11.9	14.3	7.3	5.4
$2_{\gamma}^{'+} ightarrow 4_{g}^{\circ}$	$2.6^{+2.0}_{-2.0}$	0.2	0.6	0.7	0.4	0.74
$4_{\nu}^{\prime} \rightarrow 2_{g}^{\circ}$	$0.9^{+0.3}_{-0.3}$	1.6	5	6.1	3	28.6
$4^+_{\gamma} \to 4^+_g$	$3.9_{-1.2}^{+1.2}$	5.3	15.5	18.9	9	9.6

Table 22. The same as in table 16, but for ¹⁵⁶Dy. The experimental data are taken from [63].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_g^+ \rightarrow 0_g^+$	$149.3^{+2.5}_{-2.5}$	142	138	111	137	66
$4_g^+ \rightarrow 2_g^+$	261^{+17}_{-17}	225	223	179	219	149
$4_g^+ \rightarrow 2_g^+ 6_g^+ \rightarrow 4_g^+ 8_g^+ \rightarrow 6_g^+$	200^{+15}_{-15}	282	279	235	271	221
$8_g^+ \rightarrow 6_g^+$	289^{+14}_{-14}	323	323	295	316	289
$10_g^+ \rightarrow 8_g^+$	366^{+25}_{-25}	358	358	359	357	354
$12_g^+ \rightarrow 10_g^+$	382^{+22}_{-22}	385	386	425	395	418
$2_{\gamma}^+ ightarrow 0_g^+$	$7.2^{+0.4}_{-0.4}$	6.6	9.9	23.3	11.6	7.2
$2_{\gamma}^{+} \rightarrow 2_{g}^{+}$	$9.4^{+1.0}_{-1.0}$	9.8	14.6	35.1	17.4	9.4
$2_{\gamma}^{+} \rightarrow 4_{g}^{+}$	$12.6^{+1.9}_{-1.9}$	0.5	0.7	1.8	0.9	19.5

In [62], measured data in ¹⁵⁰Nd for intraband transitions ground to ground and beta to beta as well interband transitions to ground band have been reported. These data are described reasonably well by the five approaches as shown in table 21. One remarks the good agreement obtained with the CSM approach. The largest discrepancies with the experimental data are obtained for the transitions $4^+_{\beta} \to 2^+_g$ and $4^+_{\gamma} \to 2^+_g$ which are overestimated by the theoretical results.

As for 188,190 Os, the available data are about the intraband transitions ground to ground and gamma to gamma bands as well about the interband transition beta to ground and gamma to ground. They are compared with the results of our calculations in tables 19 and 20. Again, the agreement qualities obtained with the five sets of calculations are comparable with each other. The predictions for the decay probabilities of the transitions $4^+_{\gamma} \rightarrow 2^+_g$ and $6^+_{\gamma} \rightarrow 4^+_g$

Table 23. The same as in table 16, but for ¹⁶⁶Hf. The experimental data are taken from [64].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_g^+ \rightarrow 0_g^+$	128^{+7}_{-7}	98	153	154	155	128
$4_g^+ \rightarrow 2_g^+$	202_{-7}^{+7}	155	212	216	215	203
$6_g^+ \rightarrow 4_g^+$	221^{+13}_{-13}	194	225	232	226	245
$8_g^+ \rightarrow 6_g^+$	280^{+30}_{-30}	223	225	230	225	280
$10^+_g \rightarrow 8^+_g$	250^{+640}_{-110}	246	220	219	218	311
$12_g^+ \rightarrow 10_g^+$	155^{+550}_{-70}	265	213	199	209	351

Table 24. The same as in table 16, but for ¹⁶⁸Hf. The experimental data are taken from [65].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_g^+ \rightarrow 0_g^+$	154^{+7}_{-7}	141	165	176	175	154
$4_g^+ \rightarrow 2_g^+$	244^{+12}_{-12}	223	250	257	255	249
$6_g^+ \rightarrow 4_g^+$	285^{+18}_{-18}	279	294	292	291	304
$8_g^+ \rightarrow 6_g^+$	350^{+50}_{-50}	320	322	318	316	350
$10_g^+ \rightarrow 8_g^+$	370^{+60}_{-60}	354	342	338	338	391
$12_g^+ \rightarrow 10_g^+$	320^{+120}_{-120}	381	356	354	357	438

Table 25. The same as in table 16, but for ¹⁷⁰W. The experimental data are taken from [66].

B(E2)(W.u.)	Exp.	<i>X</i> (5)	ISW	D	SSA	CSM
$2_g^+ \rightarrow 0_g^+$	124^{+3}_{-3}	79	133	126	129	124
$4_g^+ \rightarrow 2_g^+$	179^{+18}_{-18}	125	179	177	179	168
$6_g^+ \rightarrow 4_g^+$	189^{+14}_{-14}	157	184	189	187	182
$8_g^+ \rightarrow 6_g^+$	190^{+50}_{-50}	180	180	187	183	190
$10^+_g \rightarrow 8^+_g$	170^{+40}_{-40}	199	173	175	174	197
$12_g^+ \rightarrow 10_g^+$	160^{+30}_{-30}	214	167	158	162	214

are larger than the corresponding experimental data. Also, the result for $0^+_{\beta} \to 2^+_{\gamma}$ obtained within the CSM is about 6.5 larger than the corresponding experimental value. For some cases, the value of the t_2 obtained through the least-squares procedure is very large. The reason is as follows.

Within the SSA, the t_2 term of the transition operator contributes mainly to the interband transitions, while its matrix elements between states of a given band are very small. However, for the mentioned cases there are only few experimental data for interband transitions, most of the data referring to the intraband transitions. Consequently, the least-squares procedure uses small matrix elements of the intraband transitions which results in obtaining a huge number for t_2 . An equally good description of these cases would be obtained by ignoring the t_2 term. We kept however this term just for the sake of having an unitary approach.

The results for the E2 transitions raise the question: Why the models X(5), ISW, D, SSA predict close results, although the states involved are described by different wavefunctions in the variables β and γ ? It seems that these differences are washed out by the fitting procedure adopted for the strengths of the transition operator. Moreover, the factor function depending

on the Euler angles are common in the mentioned four approaches, thus giving the dominant contribution to the reduced transition probability.

One signature for the triaxiality of the nuclear shape is the equality:

$$E_{2_1^+} + E_{2_2^+} = E_{3_1^+}. (4.1)$$

The departure from this rule, $\Delta E = |E_{2_1^+} + E_{2_2^+} - E_{3_1^+}|$, is equal to 2 and 11 keV for ¹⁸⁸Os and ¹⁹⁰Os, respectively. The magnitude of these deviations was the argument for treating the two isotopes as triaxial nuclei [44]. On the other hand, the ratio $E_{4/2}$ amounts to 2.93 and 3.08 for 188 Os and 190 Os, respectively, which are quite close to the specific value of X(5) nuclei. Given these facts, we asked ourselves whether these nuclei are axially symmetric or behave like a triaxial rigid rotor. In order to answer this question, we compared the rms values of deviations for both energies and B(E2) values provided by the SMA and SSA approaches, respectively. Concerning the excitation energies in the three major bands, the rms values of prediction deviations from the corresponding experimental data yielded by the SMA for ¹⁸⁸Os and ¹⁹⁰Os are 24 and 32 keV, respectively, while the SSA results for these values being 13 and 27 keV, respectively. Therefore, regarding the excitation energies the two isotopes behave more like axially deformed nuclei. However, comparing the results for the reduced transition probabilities, it comes out that the triaxial rigid rotor behavior is favored. Indeed, the rms values for the SMA approach applied to ¹⁸⁸Os and ¹⁹⁰Os are 13 and 16 W.u., respectively, while those corresponding to the SSA are 16 and 17 W.u., respectively. Remarkable is the fact that the differences of the rms values characterizing the two approaches, SMA and SSA, are quite small. Therefore, one could conclude that the two investigations, from [44] and from here, indicate that the two nuclei might be equally well described by both approaches.

5. Conclusions

Here, we summarize the main results obtained by this work. We selected ten nuclei characterized by a ratio $R_{4_g^+/2_g^+}$ close to 2.9 which is specific to the so-called X(5) nuclei. Spectra of these nuclei are described by a new approach which treats the beta variable by the Schrödinger equation associated with a sextic oscillator plus a centrifugal potential. For the variable γ , one finds a differential equation which is satisfied by the spheroidal function. The excitation energies are obtained by summing the eigenvalues provided by the differential equations for the β and γ variables, respectively, while the corresponding functions are used to calculate the E2 transition probabilities. The results are compared with the corresponding experimental data as well as with those obtained through other formalisms called X(5), ISW, D and CSM which were previously used by the present authors to describe the spectroscopic properties of other X(5) like nuclei.

Note that while the formalisms X(5), ISW, D and SSA treat the energies and transition probabilities using the intrinsic coordinates and the rotation matrix function, the CSM is a quadrupole boson approach and therefore the mentioned observables are calculated with the collective coordinates which are specific to the laboratory frame.

A comparison of the rms values yielded by the five approaches shows that the D, CSM and SSA approaches produce the best agreement with the experimental energies. Concerning the E2 transitions, one may show that all five sets of results quantitatively describe the experimental situation in a comparable manner with a slight advantage for SSA and CSM. Since the formalisms ISW, D, and SSA differ from each other by the way the variable beta is treated, otherwise the γ equation being the same, the transition probabilities produced by the three approaches exhibit similar agreement with the experimental data. The SSA method produces very good agreement with the experimental energies for 188 Os, 150 Nd and 168 Hf.

Table 5 shows that these nuclei have the largest deformations and moreover for the first two nuclei the ratio $R_{4^+/2^+}$ has the values 3.08 and 3.11, respectively, which deviate most from the X(5) value. The quoted ratio for 150 Nd is 2.93, which is close to the X(5) value but its deformation is the largest one.

The sextic potential for the β ensures a more realistic description of the excited states where the excitation of the beta degree of freedom is important. This is best seen in the excellent agreement of the calculated excitation energies in the beta and gamma bands with the corresponding experimental data.

The final conclusion is that the SSA, proposed in this paper, proves to be a suitable tool for a realistic description of the X(5) like nuclei.

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References

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[1] Bohr A 1952 Mat.-Fys. Medd. K. Dan. Vidensk. Selsk. 16 1
     Bohr A and Mottelson B 1953 Mat.-Fys. Medd. K. Dan. Vidensk. Selsk. 27 1
 [2] Faessler A and Greiner W 1962 Z. Phys. 168 425
     Faessler A and Greiner W 1962 Z. Phys. 170 105
     Faessler A and Greiner W 1964 Z. Phys. 177 190
     Faessler A. Greiner W and Sheline R 1965 Nucl. Phys. 70 33
 [3] Gneuss G, Mosel U and Greiner W 1969 Phys. Lett. B 30 397
 [4] Hess P, Maruhn J and Greiner W 1981 Phys. Rev. C 23 2335
     Hess P, Maruhn J and Greiner W 1981 J. Phys. G: Nucl. Phys. 7 737
 [5] Raduta A A, Ceausescu V, Gheorghe A and Dreizler R M 1981 Phys. Lett. B 99 444
     Raduta A A, Ceausescu V, Gheorghe A and Dreizler R M 1982 Nucl. Phys. A 381 253
 [6] Raduta A A, Ceausescu V and Faessler A 1987 Phys. Rev. C 36 2111
 [7] Raduta A A, Lima C and Faessler A 1983 Z, Phys. A 313 69
 [8] Raduta A A, Raduta Al H and Faessler A 1997 Phys. Rev. C 55 1747
     Raduta A A, Ionescu D and Faessler A 2002 Phys. Rev. C 65 064322
 [9] Raduta A A and Sabac C 1983 Ann. Phys., NY 148 1
[10] Raduta A A 2004 Recent Research Developments in Nuclear Physics vol 1 pp 1-70 ISBN:81-7895-124-X
        (managing editor: S G Pandalai, Transworld Research Network)
[11] Ceausescu V and Raduta A A 1976 Ann. Phys., NY 100 94
[12] Raduta A A, Simkovic F and Faessler A 2000 J. Phys. G: Nucl. Part. Phys. 26 793
[13] Wilets L and Jean M 1956 Phys. Rev. 102 788
[14] Davydov A S and Filippov G F 1958 Nucl. Phys. 8 237
[15] Arima A and Iachello F 1976 Ann. Phys., NY 99 253
     Arima A and Iachello F 1979 Ann. Phys., NY 123 468
[16] Iachello F and Arima A 1987 The Interacting Boson Model (Cambridge: Cambridge University Press)
[17] Ginocchio J H and Kirson M W 1980 Phys. Rev. Lett. 44 1744
[18] Dieperink A E L, Scholten O and Iachello F 1980 Phys. Rev. Lett. 44 1767
[19] McCutchan E A, Zamfir N V and Casten R F 2004 Phys. Rev. C 69 064306
[20] McCutchan E A and Zamfir N V 2005 Phys. Rev. C 71 054306
[21] Iachello F 2000 Phys. Rev. Lett. 85 3580
[22] Iachello F 2001 Phys. Rev. Lett. 87 052502
[23] Casten R F and Zamfir N V 2000 Phys. Rev. Lett. 85 3584
[24] Casten R F and Zamfir N V 2001 Phys. Rev. Lett. 87 052503
[25] Zamfir N V et al 2002 Phys. Rev. C 65 044325
[26] Clark R M et al 2004 Phys. Rev. C 69 064322
```

[28] Bonatsos D, Lenis D, Minkov N, Petrellis D, Raychev P P and Terziev P A 2003 arXiv:nucl-th/0312121v1 Bonatsos D, Lenis D, Minkov N, Petrellis D, Raychev P P and Terziev P A 2004 Phys. Lett. B 584 40

[27] Zhang D-l and Liu Y-x 2002 Phys. Rev. C 65 057301

- [29] Davidson P M 1932 Proc. R. Soc. 135 459
- [30] Raduta A A, Gheorghe A and Faessler A 2005 J. Phys. G: Nucl. Part. Phys. 31 337
- [31] Leyvraz F and Heiss W D 2005 Phys. Rev. Lett. 95 050402
- [32] Gilmore R and Feng D H 1978 Nucl. Phys. A 301 189
- [33] Caprio M A, Cejnar P and Iachello F 2008 Ann. Phys., NY 323 1106
- [34] Fortunato L 2005 Eur. J. Phys. A 26 1-30
- [35] Gheorghe A, Raduta A A and Ceausescu V 1978 Nucl. Phys. A 296 228
- [36] Raduta A A, Gheorghe A and Ceausescu V 1978 Nucl. Phys. A 311 118
- [37] Chacon E, Moshinski M and Sharp R T 1976 J. Math. Phys. 17 668
- [38] Corrigan T M, Margetan F J and Williams S A 1976 Phys. Rev. C 14 2279
- [39] Margetan T M and Williams S A 1982 Phys. Rev. C 25 1602
- [40] Gheorghe A, Raduta A A and Ceausescu V 1998 Nucl. Phys. A 637 201
- [41] Gheorghe A, Raduta A A and Faessler A 2007 Phys. Lett. B 648 171
- [42] Raduta A A, Gheorghe A C, Buganu P and Faessler A 2009 Nucl. Phys. A 819 46
- [43] Cejnar P, Jolie J and Casten R F 2010 Rev. Mod. Phys. 82 2155
- [44] Raduta A A and Buganu P 2011 Phys. Rev. C 83 034313
- [45] Ushveridze A G 1994 Quasi-Exactly Solvable Models in Quantum Mechanics (Bristol: IOP)
- [46] Lévai G and Arias J M 2004 Phys. Rev. C 69 014304
- [47] Rose M E 1957 Elementary Theory of Angular Momentum (New York: Wiley)
- [48] Bijker R, Casten R F, Zamfir N V and McCutchan E A 2003 Phys. Rev. C 68 064304
- [49] Meyer U, Raduta A A and Faessler A 1998 Nucl. Phys. A 637 321
- [50] Raduta A A, Ceausescu V and Dreizler R M 1976 Nucl. Phys. A 272 11
- [51] Greiner W and Maruhn J Nuclear Models (Berlin: Springer) p 132 ISBN 3-540-59180-X
- [52] Raduta A A, Budaca R and Faessler A 2012 Ann. Phys., NY 327 671-704
- [53] Basunia M S 2006 Nucl. Data Sheets 107 791
- [54] Melon B 2011 Investigation of the X(5)-structure in 176Os using absolute transition probabilities Inaugural-Dissertation zur Erlangung des Doktorgrades der Mathematisch-Naturwissenschaftlochen Facultät der Universität zu Köln
- [55] Kibédi T, Dracoulis G D, Byrne A P and Davidson P M 1994 Nucl. Phys. A 567 183
- [56] Möller O et al 2005 Phys. Rev. C 72 034306
- [57] Achterberg E, Capurro O A and Marti G V 2009 Nucl. Data Sheets 110 1473
- [58] Wu S-C and Niu H 2003 Nucl. Data Sheets 100 483
- [59] Singh B 2002 Nucl. Data Sheets 95 387
- [60] Singh B 2003 Nucl. Data Sheets 99 275
- [61] Dermateosian E and Tuli J K 1995 Nucl. Data Sheets 75 827
- [62] Krücken R et al 2002 Phys. Rev. Lett. 88 232501
- [63] Reich C W 2003 Nucl. Data Sheets 99 753
- [64] Baglin C M 2008 Nucl. Data Sheets 109 1103
- [65] Baglin C M 2010 Nucl. Data Sheets 111 1807
- [66] Baglin C M 2002 Nucl. Data Sheets 96 611