Application of the sextic oscillator potential together with Mathieu and spheroidal functions for triaxial and X(5) type nuclei

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Abstract. The Bohr-Mottelson Hamiltonian is amended with a potential which depends on both β and γ deformation variables and which allows us to separate the β variable from the other variables. The equation for the β variable is quasi-exactly solved for a sextic oscillator with centrifugal barrier potential. Concerning the γ equation, its solutions are the angular spheroidal and Mathieu functions for X(5) type and triaxial nuclei, respectively. The models developed in this way are conventionally called the Sextic and Spheroidal Approach (SSA) and the Sextic and Mathieu Approach (SMA). SSA and SMA was successfully applied for several nuclei, details being presented below.

1 Introduction

A great interest in solving the eigenvalue problem of the Bohr-Mottelson Hamiltonian [1] having a potential which depends on both β and γ variables appeared when nuclei being close to the critical points of some shape phase transition were very well described by analytically solvable equations. The E(5) [2] solution describes the critical point of the transition between spherical and $\gamma-$unstable shape phases, while the one associated to the transition between spherical and axially symmetric shape phases is called X(5) [3]. Other two solutions for critical points were proposed short after that, namely Y(5) [4] and Z(5) [5], for the axial-triaxial shape phase transition and for the prolate-oblate shape phase transition, respectively.

In the this paper, we present new interesting solutions for the Hamiltonian [1], namely, Sextic and Mathieu Approach (SMA) [6–8] and Sextic and Spheroidal Approach (SSA) [9], respectively. SMA represents a realistic tool for the description of triaxial nuclei having axial deformations close to $\pi/6$, while SSA works very well for X(5) type and axial nuclei.

2 New solutions for the generalized Bohr-Mottelson Hamiltonian

The Bohr-Mottelson Hamiltonian [1],

$$\hat{H} = -\frac{\hbar^2}{2B} \left[\frac{1}{\beta^4} \frac{\partial}{\partial\beta}\beta^4\frac{\partial}{\partial\beta} + \frac{1}{\beta^2} \left(\frac{1}{\sin 3\gamma} \frac{\partial}{\partial\gamma} \sin 3\gamma \frac{\partial}{\partial\gamma} - \sum_{k=1}^{3} \frac{\hat{Q}_k^2}{4\sin^2(\gamma - \frac{2\pi}{3}k)} \right) \right] + V(\beta,\gamma), \quad (1)$$

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is amended with a potential which depends on both β and γ deformation variables and which allows us to separate the β variable from the γ variable and the three Euler angles θ_1, θ_2 and θ_3, which are still coupled due to the rotational term. Here, with Q_k are denoted the intrinsic angular momentum components. Further, by performing a second order expansion of the rotational term around $\gamma_0 = 0$ and $\gamma_0 = \pi/6$ for X(5) type nuclei and triaxial nuclei respectively, and then averaging the resulting terms with specific Wigner functions, a complete separation is achieved [6, 9]:

$$\left[-\frac{1}{\beta^4}\frac{d}{\partial\beta}\beta^4\frac{d}{d\beta} + \frac{L(L+1)}{\beta^2} + v_1(\beta) \right] f(\beta) = \varepsilon_\beta f(\beta), \ v_1(\beta) = \frac{2B}{\hbar^2}V_1(\beta), \ \varepsilon_\beta = \frac{2B}{\hbar^2}E_\beta \qquad (2)$$

$$\left[-\frac{1}{\sin 3\gamma}\frac{d}{d\gamma}\sin 3\gamma\frac{d}{d\gamma} - \tilde{W}(L,K,R) + v_2(\gamma) \right]\phi(\gamma) = \tilde{\varepsilon}_\gamma\phi(\gamma), \ v_2(\gamma) = \frac{2B}{\hbar^2}V_2(\gamma), \ \tilde{\varepsilon}_\gamma = \langle\beta^2\rangle\frac{2B}{\hbar^2}E_\gamma. \quad (3)$$

$\tilde{W}(L,K,R)$ contains the terms coming from the rotational term expansion and $L(L+1)$, R and K are the eigenvalues of the total intrinsic angular momentum $\hat{Q}$ and of its projections on the axis 1 and 3.

Making the change of function $f(\beta) = \beta^{-2}\varphi(\beta)$ and considering a sextic oscillator potential,

$$v_1^\pm(\beta) = (b^2 - 4ac^\pm)\beta^2 + 2ab\beta^4 + a^2\beta^6 + u_0^\pm, \ c^\pm = \frac{L}{2} + \frac{5}{4} + M, \ M = 0, 1, 2, ..., \qquad (4)$$

Eq. (2) is exactly solved. The meaning of the term *quasi-exactly* is given in Ref. [6]. Here, c is a constat which has two different values, one for L even and other for L odd, while $u_0^\pm$ are constants that are fixed such that the potential for L odd to have the same minimum energy with the potential for L even. The solutions for the β variable are

$$\varphi_{n_\beta,L}^{(M)}(\beta) = N_{n_\beta,L}P_{n_\beta,L}^{(M)}(\beta^2)\beta^{L+1}e^{-\frac{a}{4}\beta^4 - \frac{b}{2}\beta^2}, \ E_\beta = \frac{\hbar^2}{2B}\left[b(2L+3) + \lambda_{n_\beta}^{(M)}(L) + u_0^\pm \right], \ n_\beta = 0, 1, ..., M, \quad (5)$$

where $P_{n_\beta,L}^{(M)}(\beta^2)$ are polynomials in β^2 of n_β order and eigenvectors of the equation:

$$\left[-\frac{d^2}{d\beta^2} - \frac{4s-1}{\beta}\frac{d}{d\beta} + 2b\beta\frac{d}{d\beta} + 2a\beta^2\left(\beta\frac{d}{d\beta} - 2M\right) \right]P_{n_\beta,L}^{(M)} = \lambda_{n_\beta}^{(M)}P_{n_\beta,L}^{(M)}, \ s = \frac{L}{2} + \frac{3}{4}. \qquad (6)$$

Concerning Eq. (3), its potential is chosen such that to exhibit minima in $\gamma = 0$ and $\gamma = \pi/6$, namely, $v_2(\gamma) = u_1\cos 3\gamma + u_2\cos^2 3\gamma$. Performing a second order expansion around $\gamma_0 = 0$ in $\sin 3\gamma$ of $v_2(\gamma)$ and then substituting $x = \cos 3\gamma$ and $S(x) \equiv \phi(\gamma)$, we obtain the spheroidal equation [10]:

$$\left[(1 - x^2)\frac{d^2}{dx^2} - 2x\frac{d}{dx} + \lambda_{m_\gamma,n_\gamma} - c^2x^2 - \frac{m_\gamma^2}{1 - x^2} \right]S_{m_\gamma,n_\gamma}(x) = 0, \ m_\gamma = \frac{K}{2}, \ n_\gamma = 0, 1, 2, ... \qquad (7)$$

$$\lambda_{m_\gamma,n_\gamma} = \frac{1}{9}\left(\tilde{\varepsilon}_\gamma - \frac{u_1}{2} + \frac{L(L+1) + 11K^2 + 22}{12} \right), \ c^2 = \frac{1}{9}\left[\frac{u_1}{2} + u_2 - \frac{2}{27}(L(L+1) - K^2 - 2) \right]. \qquad (8)$$

For triaxial nuclei, expanding $v_2(\gamma)$ around $\frac{\pi}{6}$, after some steps, we get the Mathieu equation [8]:

$$\left(\frac{d^2}{dy^2} + a - 2q\cos 2y \right)M(y) = 0, \ \ y = 3\gamma, \ \ \phi(\gamma) = \frac{|M_{L,R,n_\gamma}(3\gamma)|^2}{|\sin 3\gamma|} \qquad (9)$$

$$q = \frac{1}{36}\left[\frac{10}{9}L(L+1) - \frac{13}{12}R^2 + u_1 - \frac{9}{4} \right], \ a = \frac{1}{9}\left(\tilde{\varepsilon}_\gamma + \frac{3}{4}R^2 + \frac{5}{2} \right) - 2q. \qquad (10)$$

The γ functions are normalized to unity with the integration measure $|\sin 3\gamma|d\gamma$:

$$\frac{3(2n_\gamma + 1)(n_\gamma - m_\gamma)!}{2(n_\gamma + m_\gamma)!}\int_0^{\frac{\pi}{3}}|S_{m_\gamma,n_\gamma}(\cos 3\gamma)|^2|\sin 3\gamma|d\gamma = 1, \ \frac{6}{\pi}\int_0^{\frac{\pi}{3}}\frac{|M_{L,R,n_\gamma}(3\gamma)|^2}{|\sin 3\gamma|}|\sin 3\gamma|d\gamma = 1. \ (11)$$

The total energy of the system is obtained by adding the eigenvalues of the β and γ equations.

The reduced E2 transition probabilities are determined using the following transition operator:

$$T_{2\mu}^{(E2)} = t_1\beta\left[\cos\gamma D_{\mu0}^2 + \frac{\sin\gamma}{\sqrt{2}}(D_{\mu2}^2 + D_{\mu,-2}^2)\right] + t_2\sqrt{\frac{2}{7}}\beta^2\left[-\cos 2\gamma D_{\mu0}^2 + \frac{\sin 2\gamma}{\sqrt{2}}(D_{\mu2}^2 + D_{\mu,-2}^2)\right]. \quad (12)$$

For triaxial nuclei, in Eq. (12) γ is substituted with $\gamma - 2\pi/3$. The argument is justified by the fact that $\gamma - 2\pi/3$ defines the axis 1 of the principal inertial ellipsoid. Indeed, the transformation from the laboratory to the intrinsic frame is a rotation defined by the matrix D_{MR}^L, where M and R are eigenvalues of the operator $\hat{Q}_1$. The models developed in this way are conventionally called the Sextic and Spheroidal Approach (SSA) [9] and the Sextic and Mathieu Approach (SMA) [6–8], respectively.

3 Numerical results

In Refs. [6–8], the SMA was successfully applied for several triaxial nuclei, ^{188,190,192}Os, ^{228,230}Th, ^{180}Hf and ^{182}W, chosen according to a certain signature of the rigid triaxial rotor. In Ref. [9], a good agreement of the SSA results with experimental data of several X(5) candidate nuclei as ^{176,178,180,188,190}Os, ^{150}Nd, ^{170}W, ^{156}Dy and ^{166,168}Hf, was obtained. In Ref. [9], the SSA results were compared with those yielded by X(5), ISW [10], D [10] and Coherent State Model (CSM) [11]. From space reasons we present here only one nucleus for each of the models, SMA and SSA. For SMA, the numerical results for ^{192}Os are shown in Fig. 1 and Table 1. Both, energy spectrum and reduced probability transitions are very well explained by the SMA and CSM. Also, the staggering behavior [15] of the γ band is reproduced by the SMA. Comparing the results for ^{188}Os presented in Tables 2, 3, we can see that the best agreement with experimental data, for both energy spectrum and E2 transition probabilities, is obtained with SSA.

Figure 1: Excitation energies, given in keV, for ground, β and γ bands for ^{192}Os, calculated with SMA and CSM, are compared with the experimental data [12, 13].

$B(E2; J_i^+ \to J_f^+)$	Exp.	SMA	CSM
$2_g^+ \to 0_g^+$	0.424	0.424	0.236
$4_g^+ \to 2_g^+$	0.497	0.632	0.449
$6_g^+ \to 4_g^+$	0.660	0.858	0.611
$8_g^+ \to 6_g^+$	0.754	1.030	0.754
$10_g^+ \to 8_g^+$	0.688	1.175	0.887
$4_\gamma^+ \to 2_\gamma^+$	0.298	0.261	0.277
$6_\gamma^+ \to 4_\gamma^+$	0.336	0.352	0.595
$8_\gamma^+ \to 6_\gamma^+$	0.314	0.549	0.814
$2_\gamma^+ \to 0_g^+$	0.037	0.006	0.192
$2_\gamma^+ \to 2_g^+$	0.303	0.303	0.055
$2_\gamma^+ \to 4_g^+$	0.024	0.000	0.000
$4_\gamma^+ \to 2_g^+$	0.002	0.004	0.274
$4_\gamma^+ \to 4_g^+$	0.203	0.068	0.137
$4_\gamma^+ \to 6_g^+$	0.018	0.000	0.000
$6_\gamma^+ \to 4_g^+$	0.000	0.002	0.357
$6_\gamma^+ \to 6_g^+$	0.171	0.042	0.171

Table 1: E2 transition probabilities, given in (eb)^2, for ^{192}Os, calculated with SMA and CSM, are compared with the experimental data [12, 13].

^{188}Os	Exp.	X(5)	ISW	D	SSA	CSM
2_g^+	155	179	179	151	152	150
4_g^+	478	519	519	479	476	468
6_g^+	940	970	970	945	935	934
8_g^+	1515	1516	1516	1512	1501	1535
10_g^+	2170	2149	2150	2156	2154	2264
12_g^+	2856	2867	2868	2860	2877	3116
0_β^+	1086	1009	1007	1120	1063	1164
2_β^+	1305	1331	1328	1270	1330	1305
4_β^+		1910	1907	1599	1808	1621
6_β^+		2636	2632	2064	2421	2096
8_β^+		3474	3470	2632	3132	2717
10_β^+		4412	4407	3276	3920	3475
2_γ^+	633	631	631	627	641	665
3_γ^+	790	786	785	773	791	790
4_γ^+	966	972	971	959	969	956
5_γ^+	1181	1185	1185	1180	1172	1157
6_γ^+	1425	1423	1423	1432	1434	1399
7_γ^+	1686	1685	1684	1709	1674	1669
8_γ^+		1969	1969	2009	2008	1983
9_γ^+		2275	2275	2329	2273	2318
10_γ^+		2602	2603	2666	2670	2701
r.m.s. [keV]		27	27	16	13	36

Table 2: The energy spectrum, given in keV, of the ground, β and γ bands for ^{188}Os yielded by the X(5), ISW, D, SSA and CSM are compared with the experimental data taken from Ref. [14].

B(E2)(W.u.)	Exp.	X(5)	ISW	D	SSA	CSM
$2_g^+ \rightarrow 0_g^+$	79^{+2}_{-2}	74	72	79	82	42
$4_g^+ \rightarrow 2_g^+$	133^{+8}_{-8}	118	115	121	123	87
$6_g^+ \rightarrow 4_g^+$	138^{+8}_{-8}	147	144	147	145	125
$8_g^+ \rightarrow 6_g^+$	161^{+11}_{-11}	169	166	174	162	161
$10_g^+ \rightarrow 8_g^+$	188^{+25}_{-25}	187	184	203	178	195
$0_\beta^+ \rightarrow 2_g^+$	$0.95^{+0.08}_{-0.08}$	47	48	33	21	0.95
$0_\beta^+ \rightarrow 2_\gamma^+$	$4.3^{+0.5}_{-0.5}$	5.2	5.2	1.9	1.5	44
$4_\gamma^+ \rightarrow 2_\gamma^+$	47^{+10}_{-10}	47	50	52	56	14
$4_\gamma^+ \rightarrow 3_\gamma^+$	320^{+120}_{-120}	112	117	120	132	43
$6_\gamma^+ \rightarrow 4_\gamma^+$	70^{+30}_{-30}	107	111	114	118	31
$2_\gamma^+ \rightarrow 0_g^+$	$5^{+0.6}_{-0.6}$	8.4	10.9	10.8	9.9	5
$2_\gamma^+ \rightarrow 2_g^+$	16^{+2}_{-2}	13	17	16	14	10.4
$2_\gamma^+ \rightarrow 4_g^+$	34^{+5}_{-5}	0.65	0.85	0.80	0.73	1.4
$4_\gamma^+ \rightarrow 2_g^+$	$1.29^{+0.19}_{-0.19}$	5.7	7.1	6.7	6.1	1.7
$4_\gamma^+ \rightarrow 4_g^+$	19^{+3}_{-3}	18	23	20	19	10.7
$4_\gamma^+ \rightarrow 6_g^+$	16^{+7}_{-7}	2	2	2	2	5
$6_\gamma^+ \rightarrow 4_g^+$	$0.21^{+0.11}_{-0.11}$	5.3	6.4	5.8	5.3	0.9
$6_\gamma^+ \rightarrow 6_g^+$	>9.4	21	25	23	20	8.3

Table 3: The reduced E2 transition probabilities determined with the X(5), ISW, D, SSA and CSM models for ^{188}Os are compared with the corresponding experimental data taken from Ref. [14].

4 Conclusions

SSA and SMA represent realistic tools for the description of X(5) type and triaxial nuclei. The Mathieu and spheroidal functions are periodic, defined on bounded intervals and normalized to unity with the integration measure $|\sin 3\gamma|d\gamma$, preserving in this way the hermiticity of the initial γ Hamiltonian. CSM works very well also for nuclei being in critical points of the shape phase transitions.

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